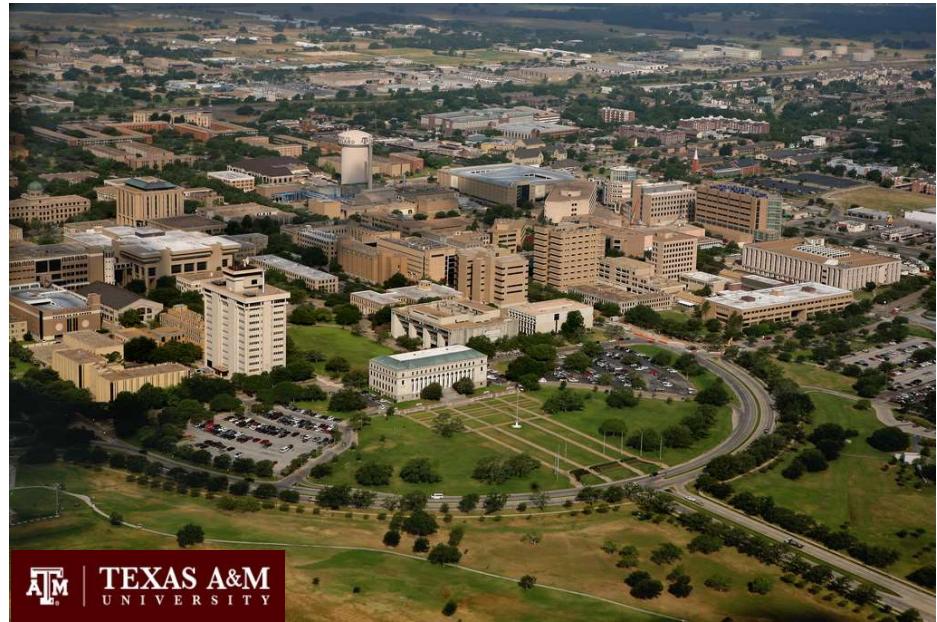


From Landau-Zener transitions to Hawking radiation

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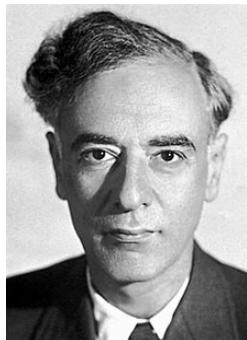
Hagler Institute for Advanced Study at Texas A&M
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Overview

- Landau-Zener transitions
- Logarithmic phase singularity
- Inverted harmonic oscillator
- Experiment

The Landau–Zener formula made simple

Eric P Glasbrenner^{1,*}  and Wolfgang P Schleich^{1,2} 



$$H \equiv \hbar \begin{pmatrix} -\alpha t & g \\ g & \alpha t \end{pmatrix}$$



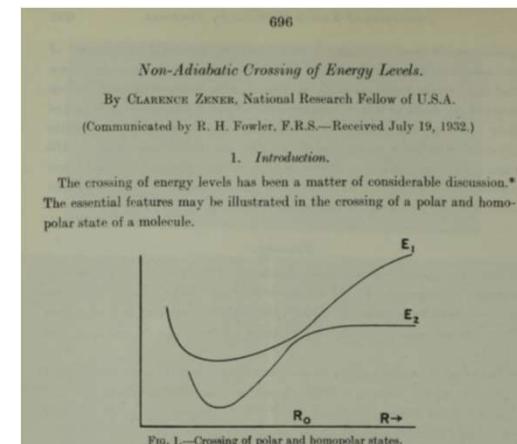
9. A THEORY OF ENERGY TRANSFER. II

The results of I are applied to: 1, predissociation; 2, excitation of oscillations during an optical transition; 3, the case of closely approaching potential curves. The question of the cross-section is discussed.

THE results obtained in I¹ can be applied to a number of effects. It must be added, however, that during the collision of similar atoms it is necessary to add to the selection rules developed in I the postulate of the change in parity which is evident from the fact that the electron-spin components which are normal to the line connecting the nuclei change their sign when mirrored.

1. The most direct generalisation is the predissociation. For the dissociation probability during an oscillation we can use here expression (21) directly. This can also be written in the following form:

$$\omega = \frac{4\pi \hbar D^2}{(F_1 - F_2) m^2 r^4} \cdot \frac{j^2}{v}. \quad (1)$$



Formulation of the problem

$$H \equiv \hbar \begin{pmatrix} -\alpha t & g \\ g & \alpha t \end{pmatrix}$$

$$\dot{a}(t) = -ig e^{-i\alpha t^2/2} b(t)$$

$$\dot{b}(t) = -ig e^{i\alpha t^2/2} a(t)$$

$$\dot{a}(t) = -g^2 e^{-i\alpha t^2/2} \int_{-\infty}^t dt' e^{i\alpha t'^2/2} a(t')$$

One-line derivation

$$\dot{a}_M(t) = -g^2 e^{-i\alpha t^2/2} \int_{-\infty}^t dt' e^{i\alpha t'^2/2} a_M(t')$$

$$I = \int_{-\infty}^{\infty} dt e^{-i\alpha t^2/2} \int_{-\infty}^0 dt' e^{i\alpha t'^2/2} \\ + \int_{-\infty}^{\infty} dt e^{-i\alpha t^2/2} \int_0^t dt' e^{i\alpha t'^2/2}$$

$$a_M(t \rightarrow \infty) = \exp\left(-\pi \frac{g^2}{\alpha}\right) = a(t \rightarrow \infty)$$

Validity of Markov approximation

$$\dot{a}(t) = -g^2 e^{-i\alpha t^2/2} \int_{-\infty}^t dt' e^{i\alpha t'^2/2} a(t')$$

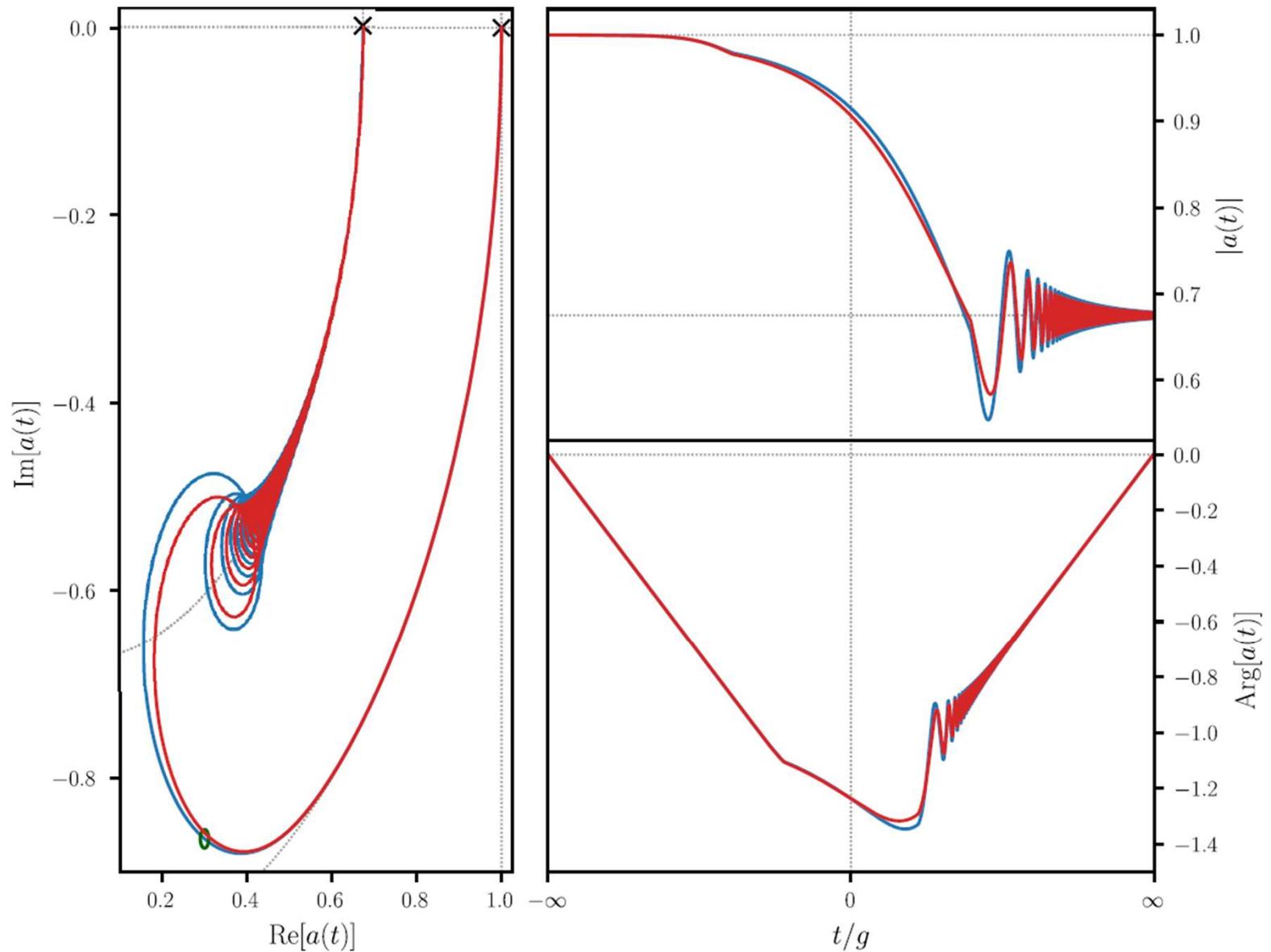
$$\dot{a}(t) = -g^2 \int_0^\infty d\tau e^{-i\alpha \tau t} e^{i\alpha \tau^2/2} a(t - \tau)$$

Landau-Zener transition: jump between circular paths

$$\dot{a}(t) = -g^2 \int_0^\infty d\tau e^{-i\alpha\tau t} e^{i\alpha\tau^2/2} a(t-\tau)$$

$$\dot{a}_S(t) \approx -g^2 \int_0^\infty d\tau e^{-i\alpha t\tau} a_S|(t)$$

$$a_S(t) = \exp \left[-\pi \frac{g^2}{\alpha} \Theta(t) \right] \exp \left[i \frac{g^2}{\alpha} \mathcal{P} \int_{-T}^t dt' \frac{1}{t'} \right]$$



Stückelberg oscillations

$$\dot{a}(t) = -g^2 \int_0^\infty d\tau e^{-i\alpha\tau t} e^{i\alpha\tau^2/2} a(t-\tau)$$



Ernst Carl Gerlach
Stückelberg v. Breidenbach

$$\dot{a}(t) = -g^2 e^{-i\alpha t^2/2} \int_0^\infty d\tau e^{i\alpha(\tau-t)^2/2} a(t-\tau)$$

Alternative approach

$$\dot{a}(t) = -ig e^{-i\alpha t^2/2} b(t)$$

$$\dot{b}(t) = -ig e^{i\alpha t^2/2} a(t)$$

$$\ddot{a}(t) + i\alpha t \dot{a}(t) + g^2 a(t) = 0$$

$$\dot{a}_S(t) = -g^2 \frac{1}{\varepsilon + i\alpha t} a_S(t) = -g^2 \frac{\varepsilon - i\alpha t}{\varepsilon^2 + (\alpha t)^2} a_S(t)$$

Black hole explosions?

From this it follows that the number of particles emitted in this wave packet mode is $(\exp(2\pi\omega/\kappa) - 1)^{-1}$ times the number of particles that would have been absorbed from a similar wave packet incident on the black hole from I^- . But this is just the relation between absorption and emission cross sections that one would expect from a body with a temperature in geometric units of $\kappa/2\pi$. Similar results hold for

S. W. HAWKING

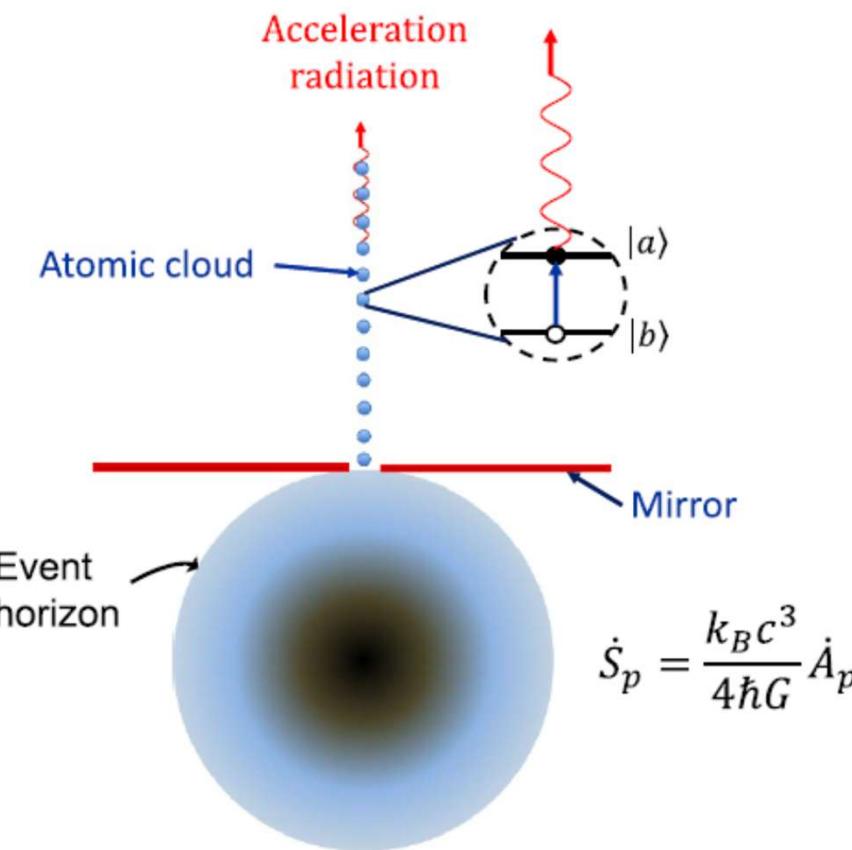
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Quantum optics approach to radiation from atoms falling into a black hole

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$$P_{exc} = \frac{g^2}{\omega^2} \left| \int_0^\infty dx e^{-2i\nu \ln x} e^{-ix(1+\frac{2\nu}{\omega})} \right|^2$$

$$= \frac{4\pi g^2 \nu}{\omega^2 (1 + \frac{2\nu}{\omega})^2} \frac{1}{e^{4\pi\nu} - 1}$$

Logarithmic phase singularity

$$\psi(r) = (r - r_0)^{i\kappa} = e^{i\kappa \ln(r - r_0)}$$

$$(r - r_0) \frac{d\psi}{dr} = i\kappa\psi$$



Tunneling of an energy eigenstate through a parabolic barrier viewed from Wigner phase space



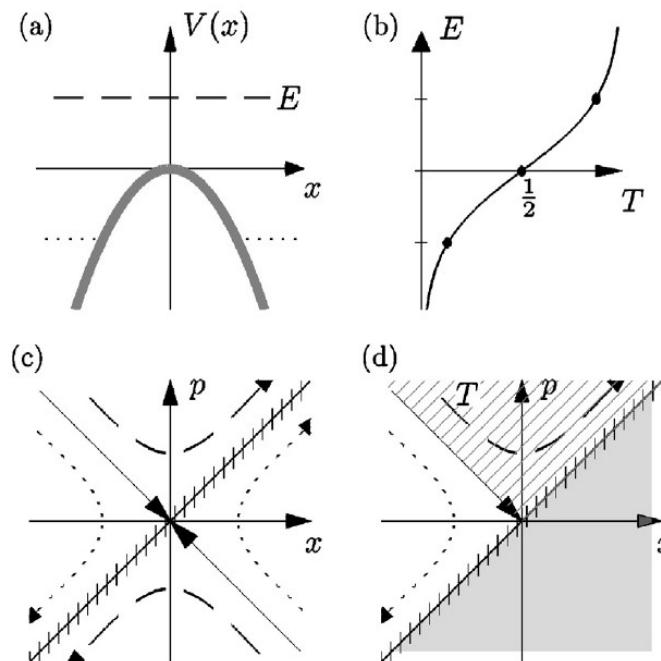
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Inverted oscillator: energy eigenstates

$$\hat{H} \equiv \frac{\hat{p}^2}{2m} - \frac{m\omega^2}{2}\hat{x}^2$$

$$\hat{\xi} \equiv \frac{1}{\sqrt{2}} \left(\hat{x} - \frac{\hat{p}}{m\omega} \right) \quad \hat{\pi} \equiv \frac{1}{\sqrt{2}} (\hat{p} + m\omega\hat{x})$$

$$\hat{H} = -\omega\hat{\xi}\hat{\pi} + \frac{i}{2}\hbar\omega$$

$$u(\xi) = \frac{N_E}{\sqrt{|\xi|}} \exp \left(-i \frac{E}{\hbar\omega} \ln |\xi| \right)$$

The logarithmic phase singularity in the inverted harmonic oscillator

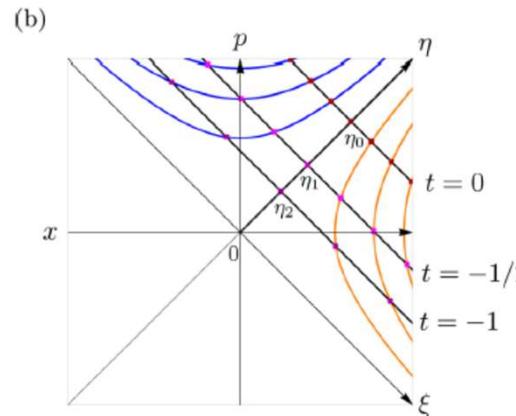
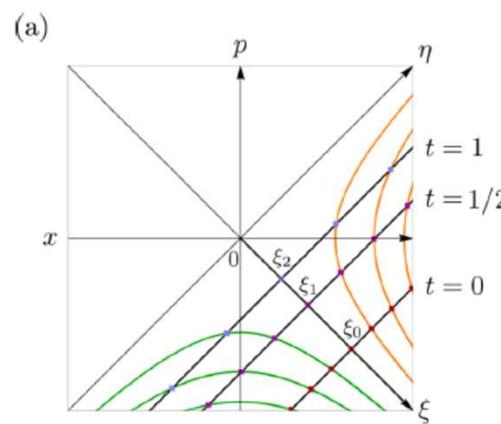
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Fermi-Dirac statistics

$$\langle \eta | \Psi_{\varepsilon}^{+} \rangle = \frac{1}{2\pi} \int_0^{\infty} d\xi \frac{\exp(-i\varepsilon \ln |\xi|)}{\sqrt{\xi}} \exp(-i\xi \eta)$$

$$|\langle \eta | \Psi_{\varepsilon}^{+} \rangle|^2 = \frac{1}{2\pi|\eta|} \left[\frac{\Theta(\eta)}{1 + e^{2\pi\varepsilon}} + \frac{\Theta(-\eta)}{1 + e^{-2\pi\varepsilon}} \right]$$

Physica Scripta



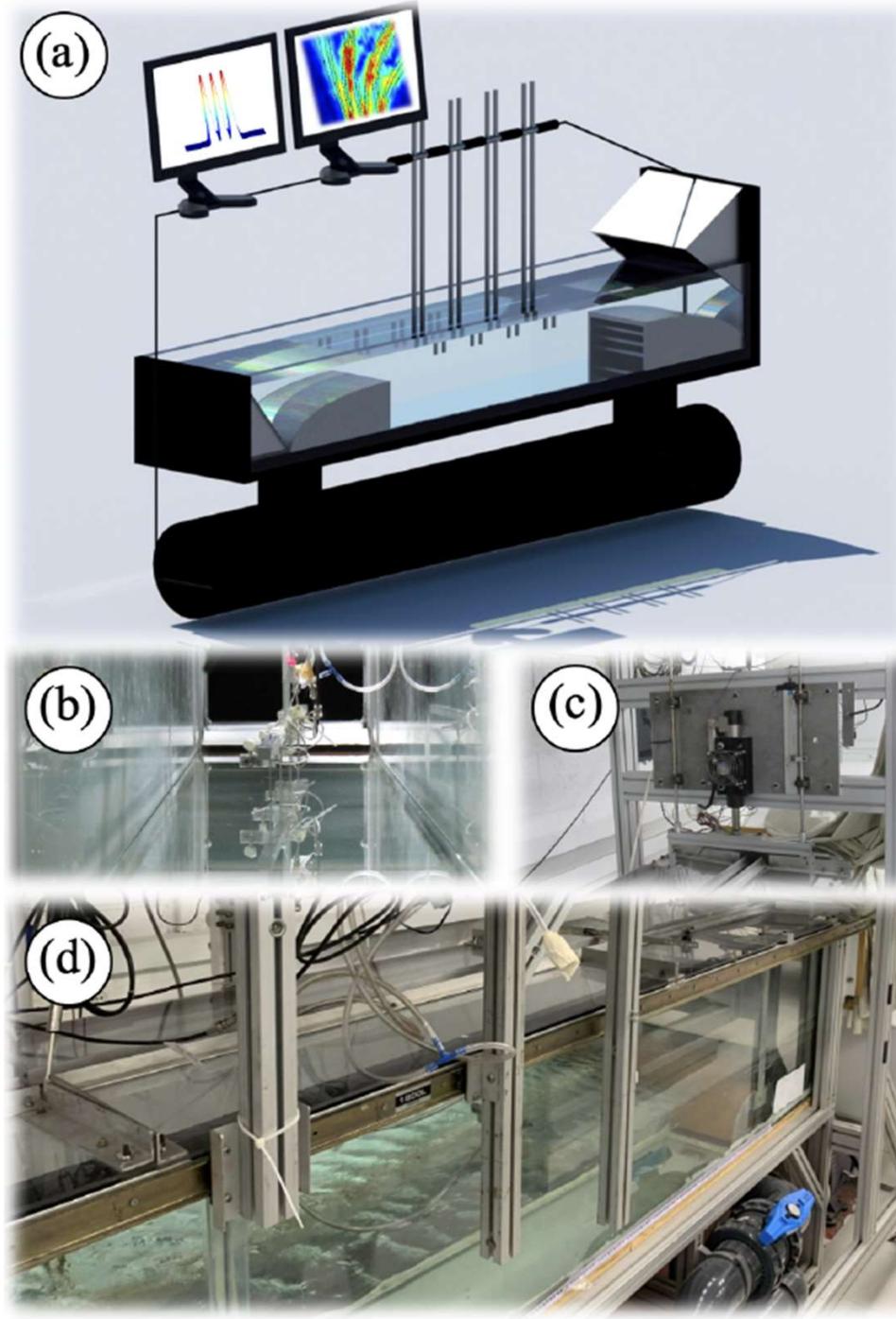
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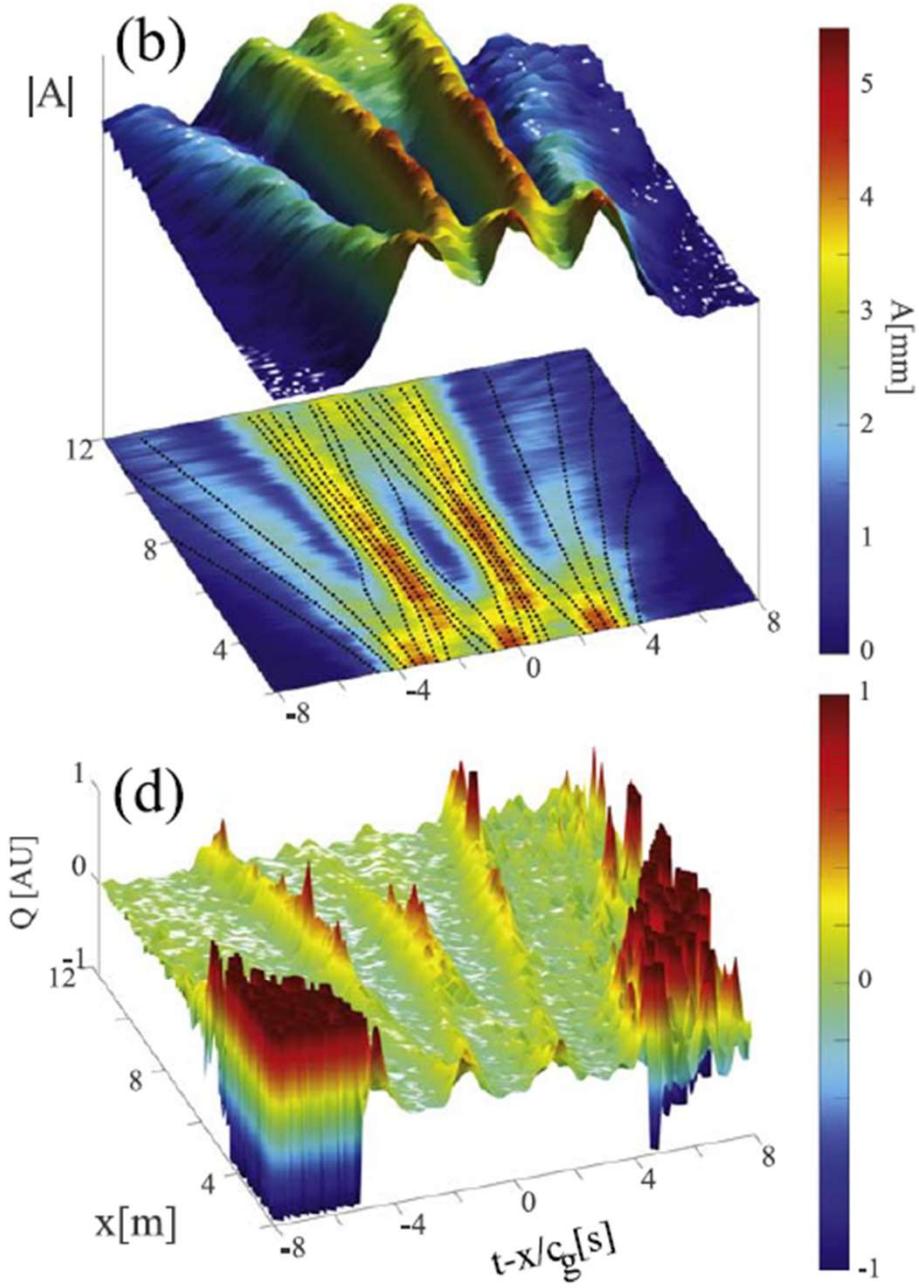
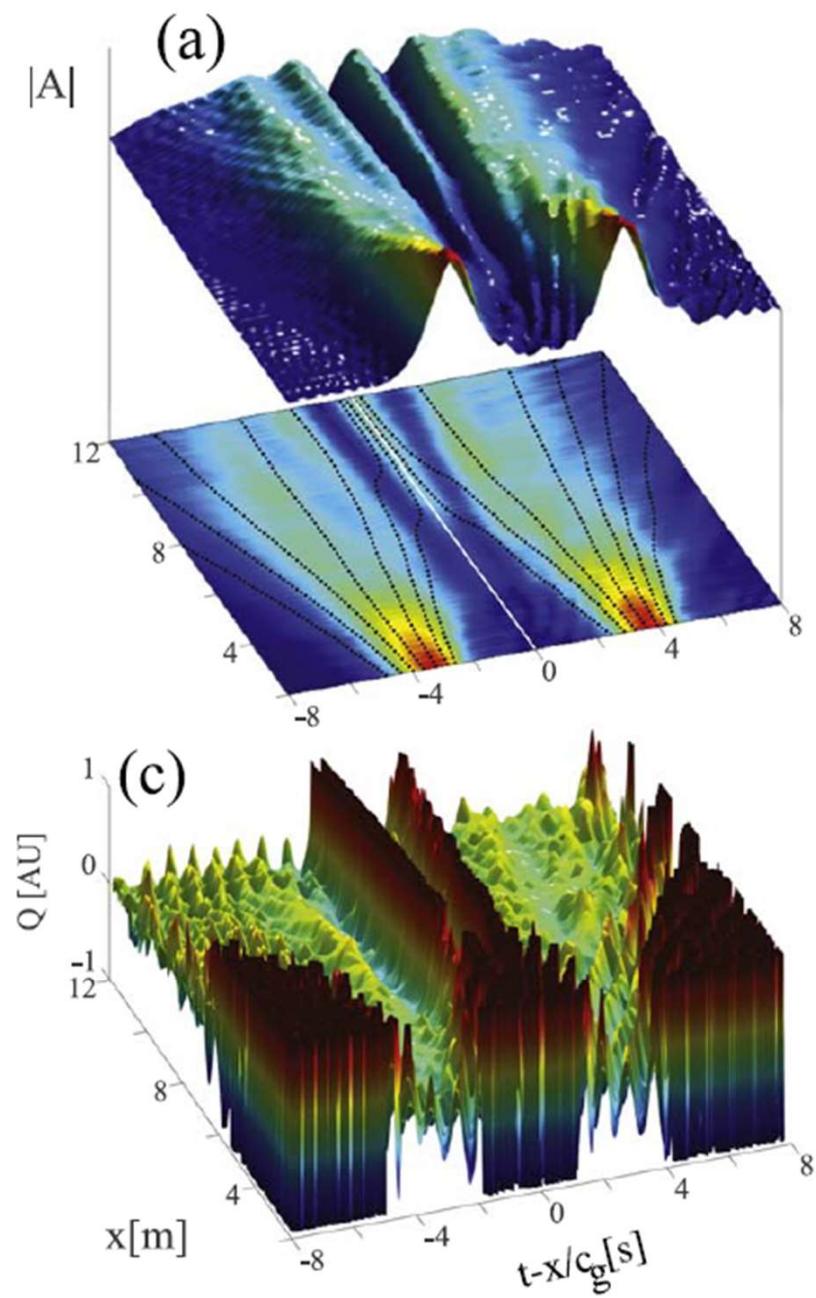
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Observation of Bohm trajectories and quantum potentials of classical waves

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16 October 2022Georgi Gary Rozenman^{1,2} , Denys I Bondar³ , Wolfgang P Schleich^{4,5} , Lev Shemer⁶ and Ady Arie²

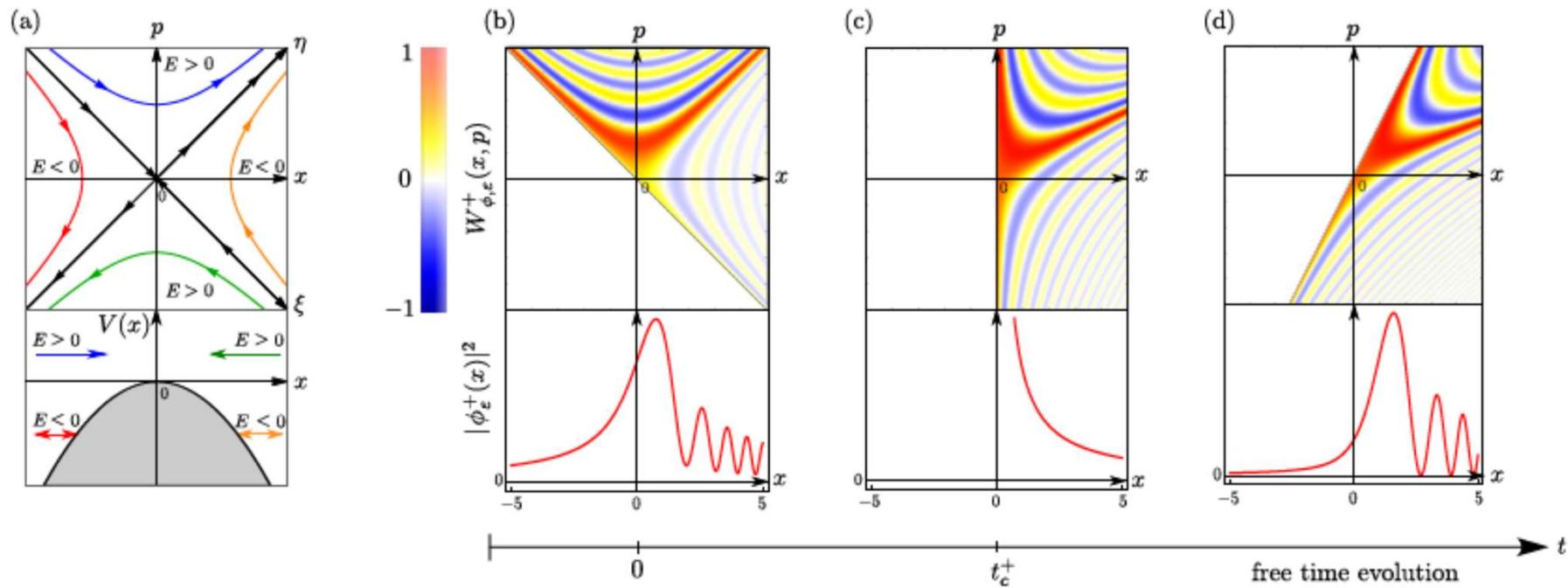
Quantity	Quantum mechanics	Surface gravity water waves
Complex-valued function	wave function $\psi(x, t)$	surface envelope $A(\tau, \xi)$
Propagation coordinate	t	ξ
Transverse coordinate	x	τ
Wave equation	$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2}$	$i \frac{\partial A}{\partial \xi} = \frac{\partial^2 A}{\partial \tau^2}$
Guiding equation	$\frac{d}{dt} \bar{x}(t) = \frac{\hbar}{m} \text{Im} \left\{ \frac{1}{\psi(\bar{x}(t), t)} \frac{\partial}{\partial \bar{x}} \psi \right\}$	$\frac{d}{d\xi} \bar{\tau}(\xi) = 2 \text{Im} \left\{ \frac{1}{A(\xi, \bar{\tau}(\xi))} \frac{\partial}{\partial \bar{\tau}} A \right\}$
Quantum potential	$Q = -\frac{\hbar^2}{2m} \frac{1}{ \psi } \frac{\partial^2}{\partial x^2} \psi $	$Q = -\frac{1}{ A } \frac{\partial^2}{\partial \tau^2} A $

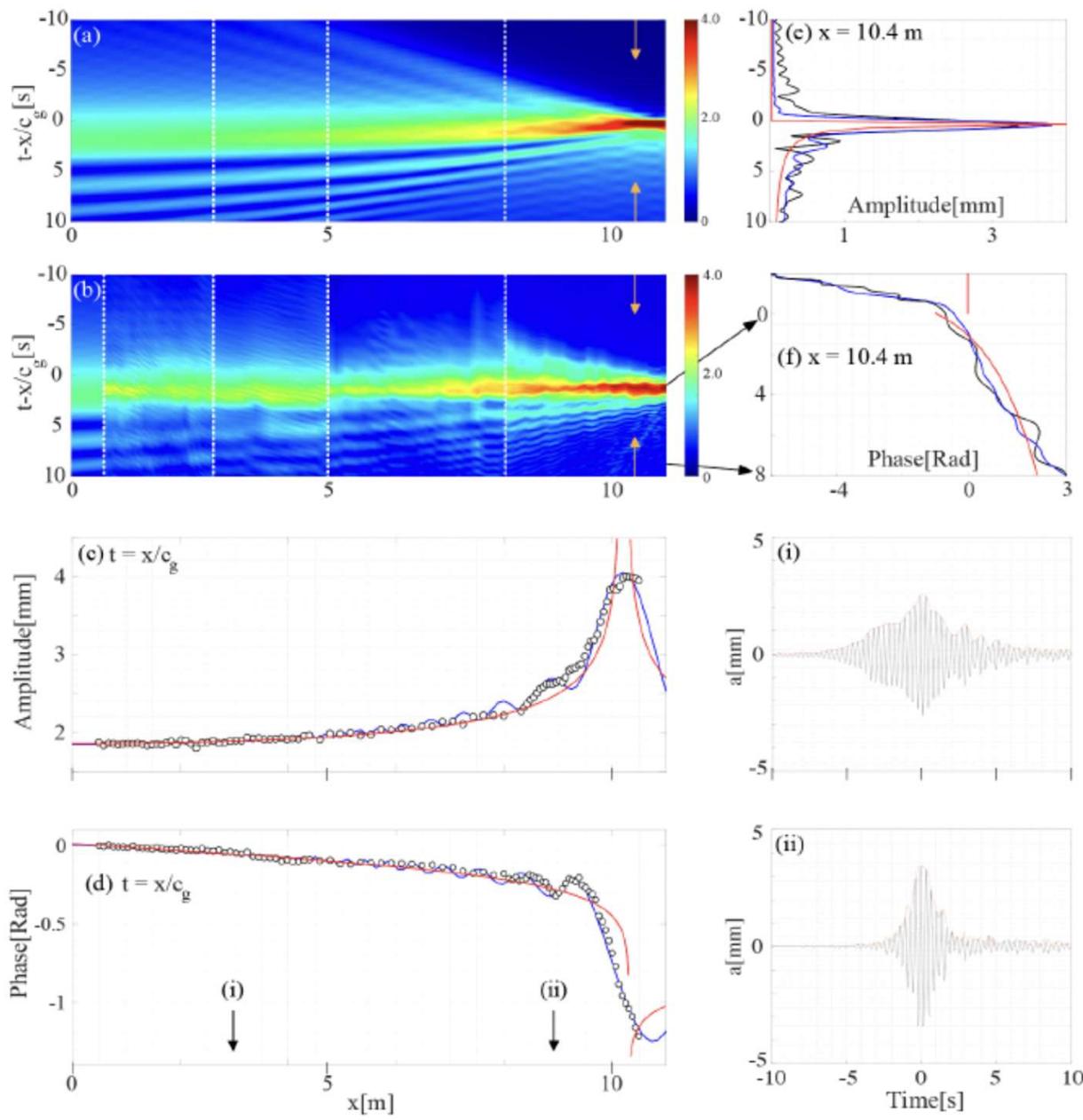


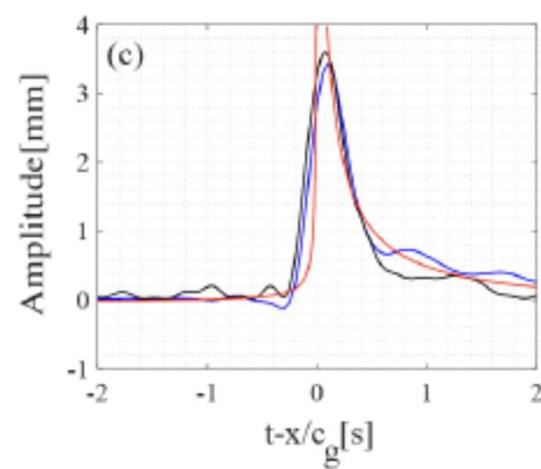
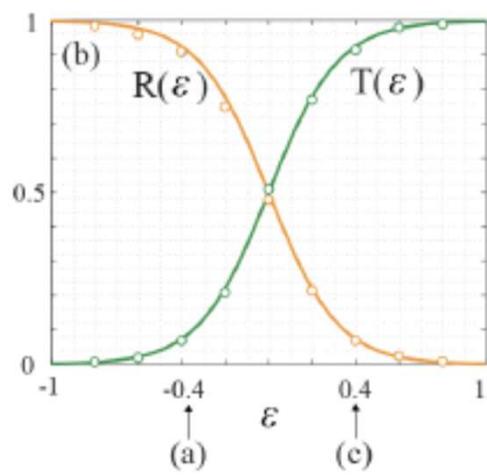
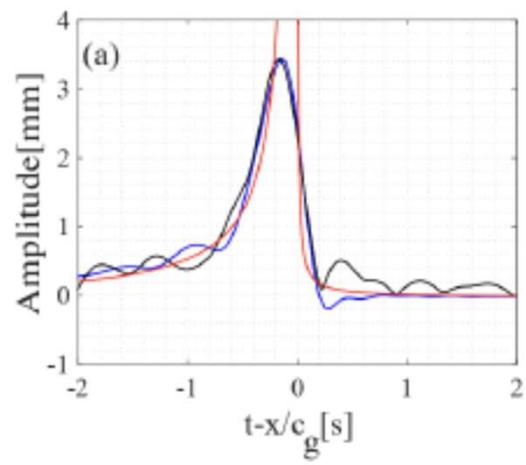


Observation of a phase space horizon with surface gravity water waves

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Maxim A. Efremov^{3,4}, Lev Shemer⁵, Wolfgang P. Schleich^{4,6}, and Ady Arie^{2,*}







Summary

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