
Scattering of dark Pions



Yannick Dengler, Axel Maas, Fabian Zierler

ACHT 2025

Scattering of dark Pions



Yannick Dengler, Axel Maas, Fabian Zierler

ACHT 2025

Outline

- ❖ Self-Interacting Dark Matter

Scattering of dark Pions



Yannick Dengler, Axel Maas, Fabian Zierler

ACHT 2025

Outline

- ❖ Self-Interacting Dark Matter
- ❖ Minimal Realisation: $\mathrm{Sp}(4)_c$ with $N_f = 2$

Scattering of dark Pions



Yannick Dengler, Axel Maas, Fabian Zierler

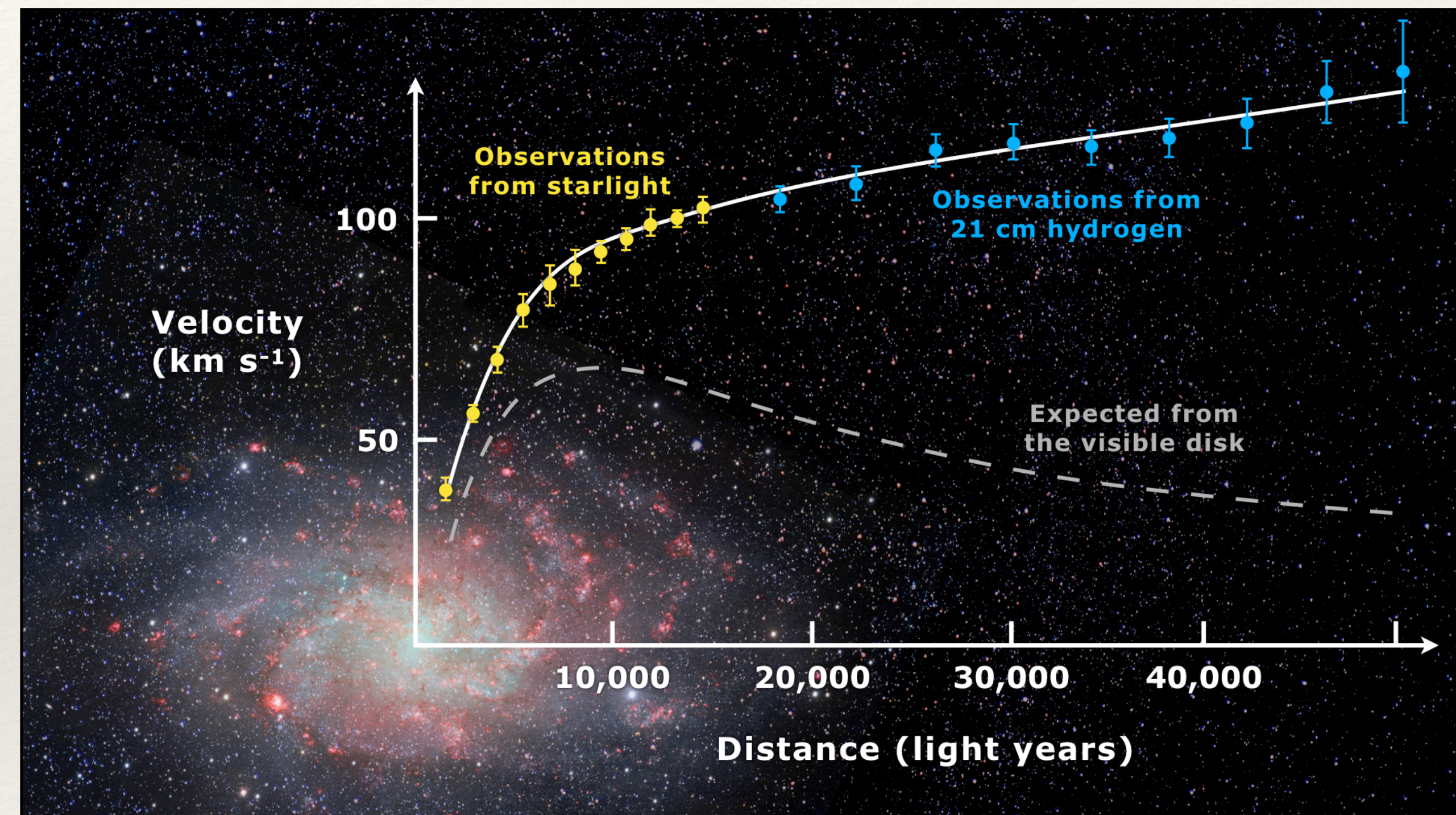
ACHT 2025

Outline

- ❖ Self-Interacting Dark Matter
- ❖ Minimal Realisation: $\mathrm{Sp}(4)_c$ with $N_f = 2$
- ❖ Lattice Results

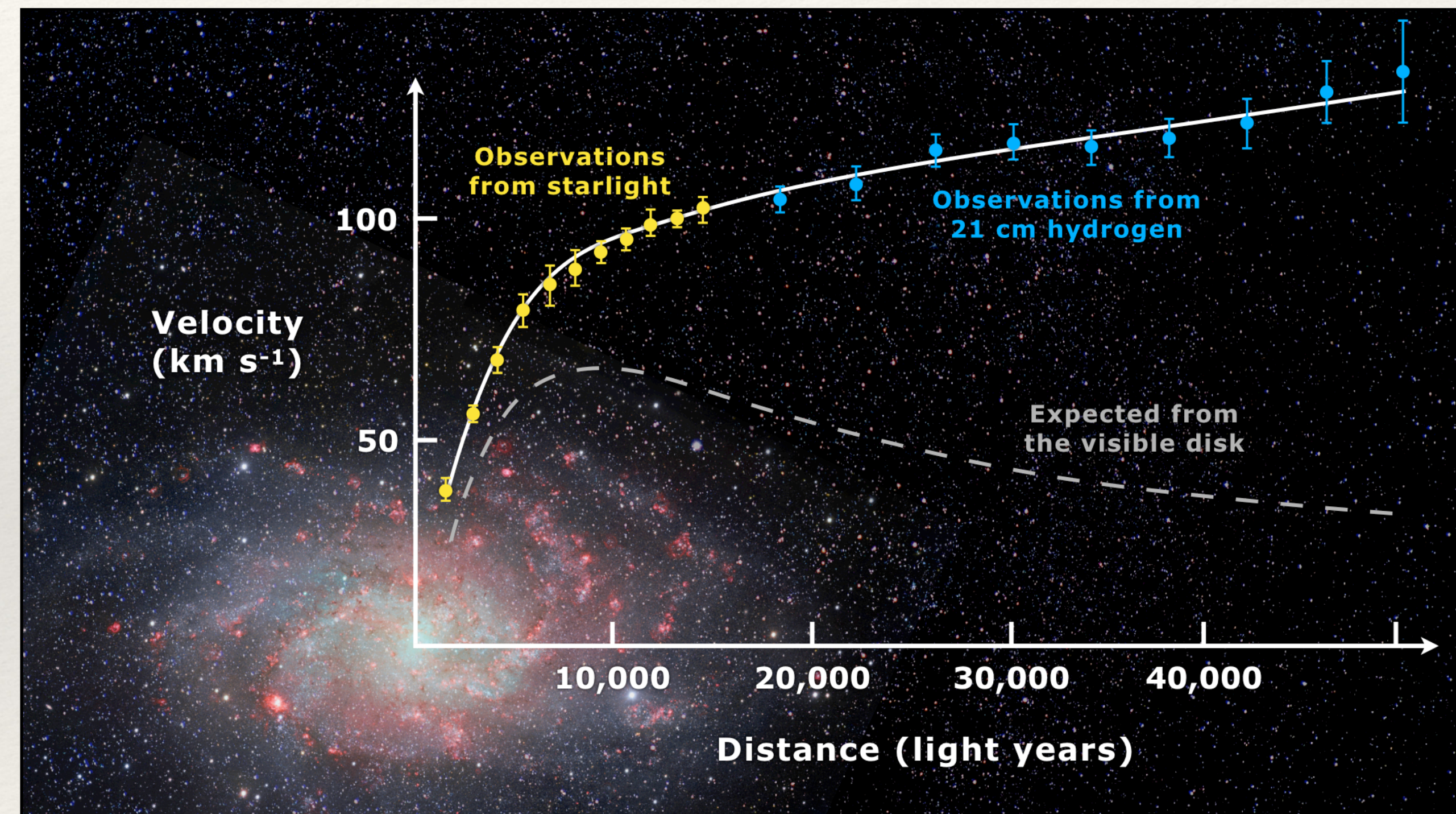
Dark Matter

- ❖ Collection of astronomical phenomena
 - ❖ Motion of objects, Large scale structure, gravitational lensing, ...
 - ❖ In conflict with our current understanding



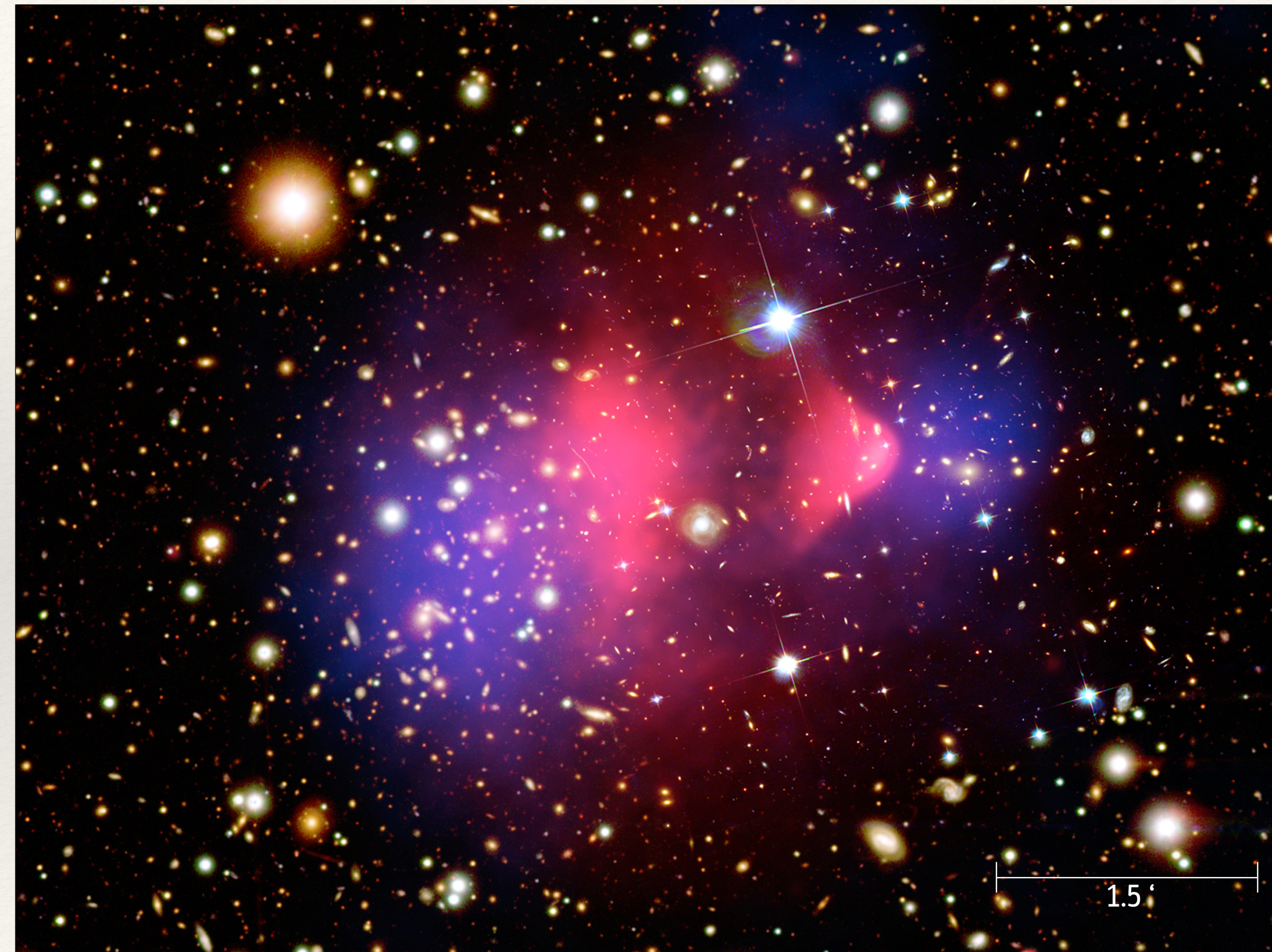
Dark Matter

- ❖ Collection of astronomical phenomena
 - ❖ Motion of objects, Large scale structure, gravitational lensing, ...
 - ❖ In conflict with our current understanding
- ❖ Explanations:
 - ❖ Modified Gravity
 - ❖ Not observable sort of matter



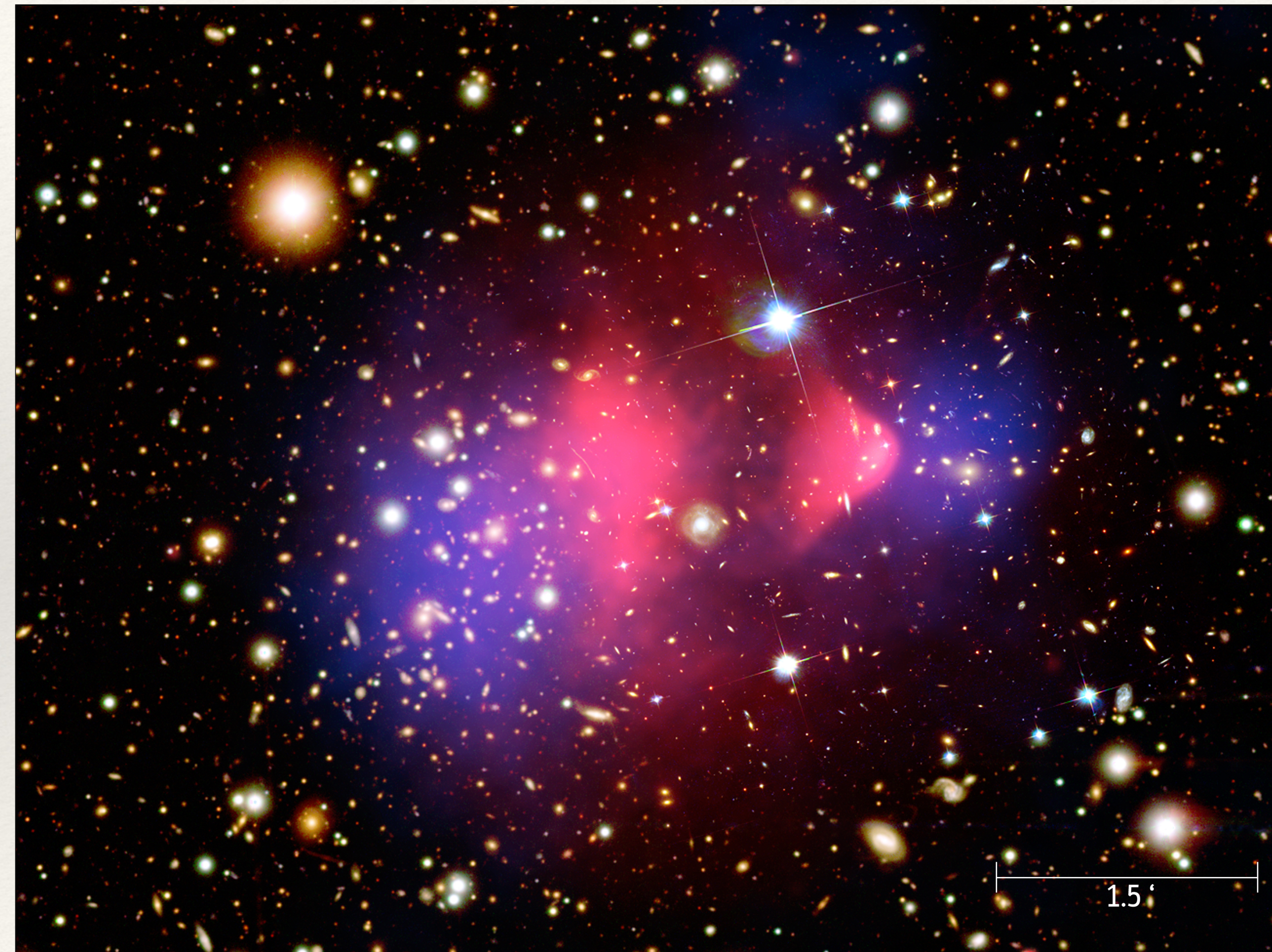
Dark Matter Particle

- ❖ Evidence for particle DM:
 - ❖ i.e. "Bullet cluster"



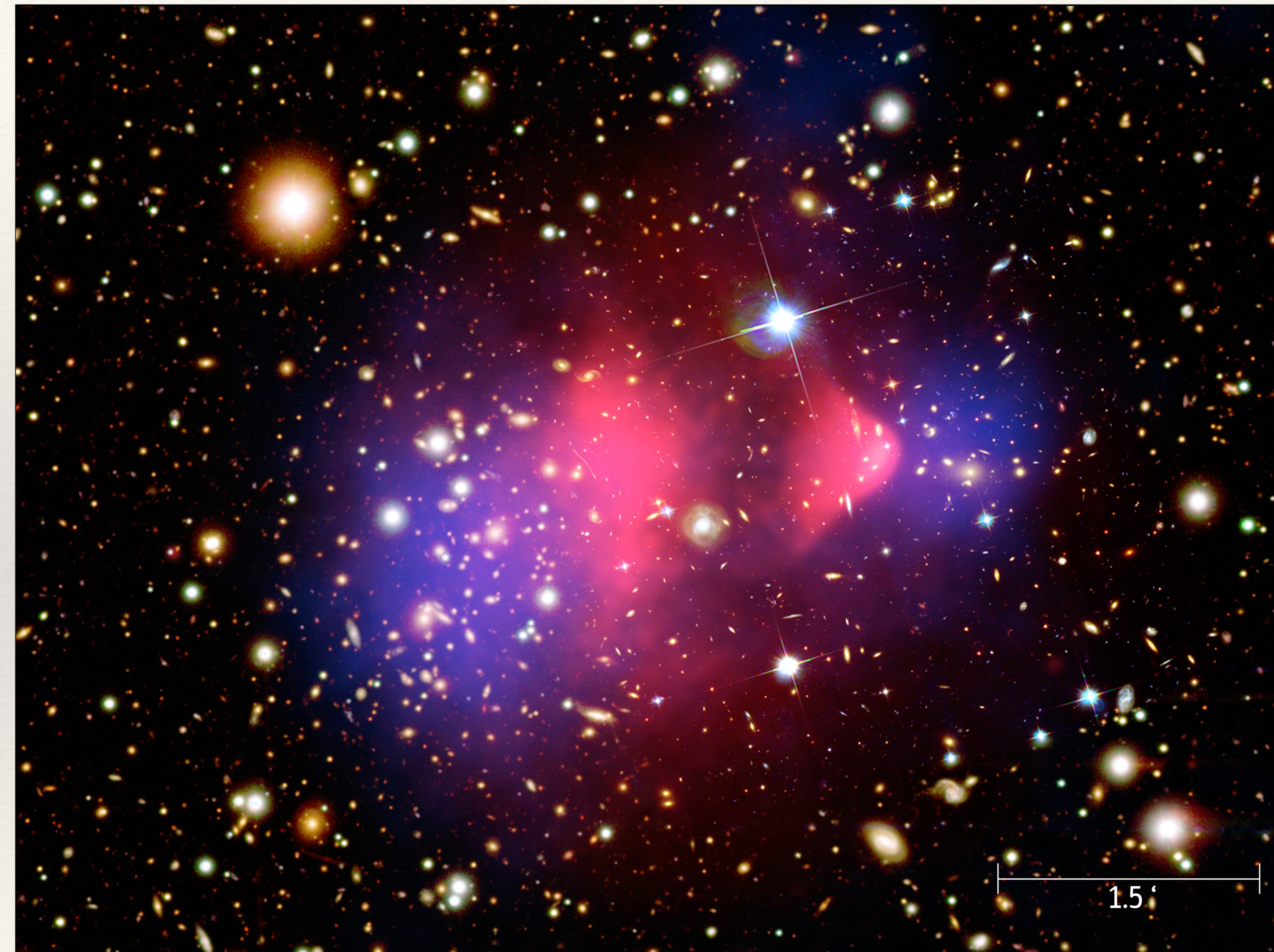
Dark Matter Particle

- ❖ Evidence for particle DM:
 - ❖ i.e. "Bullet cluster"
- ❖ Properties:
 - ❖ Massive, stable, "invisible"



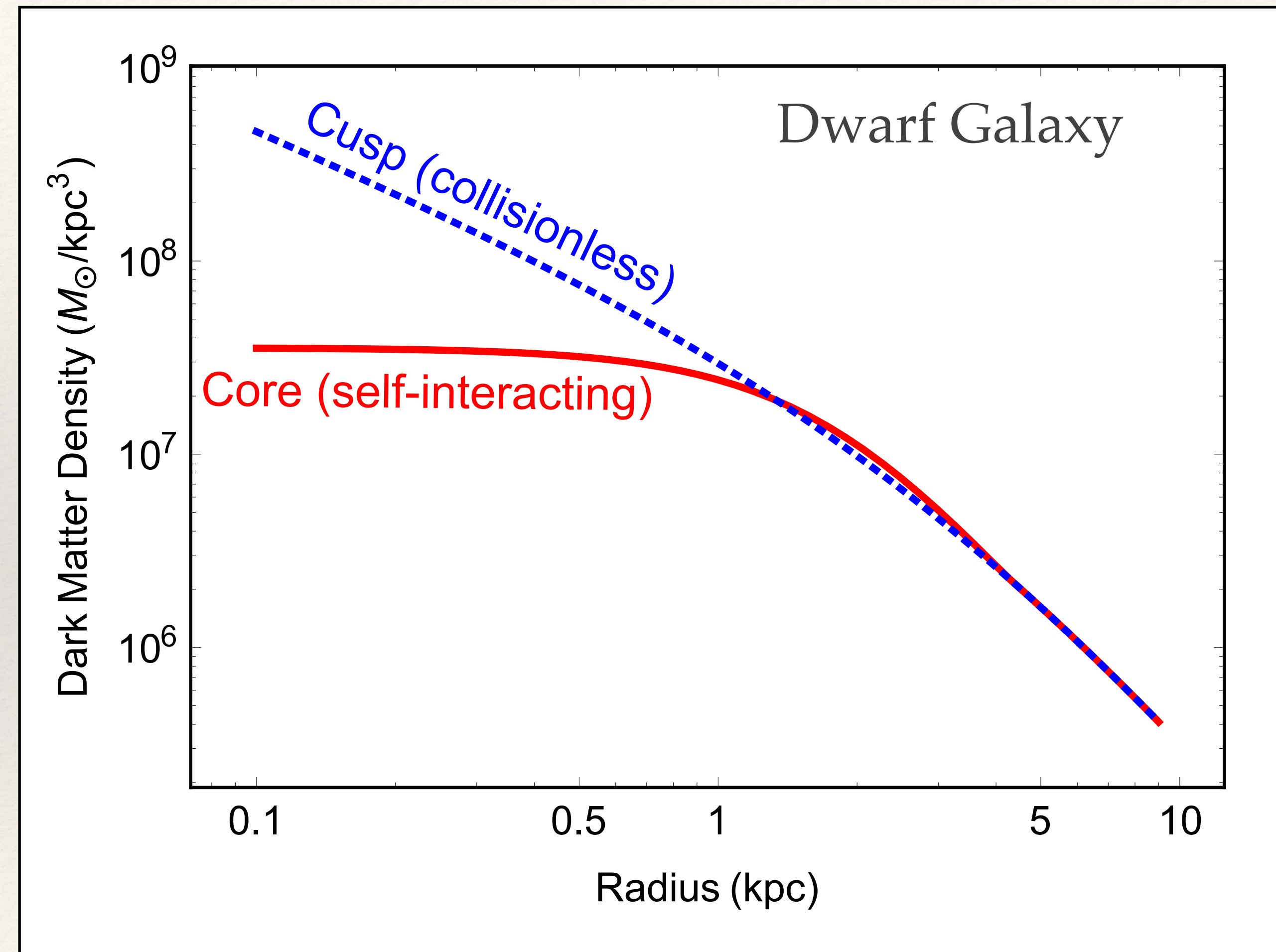
Dark Matter Particle

- ❖ Evidence for particle DM:
 - ❖ i.e. "Bullet cluster"
- ❖ Properties:
 - ❖ Massive, stable, "invisible"
- ❖ Interaction?
 - ❖ With SM: no (low)
 - ❖ Self: Maybe



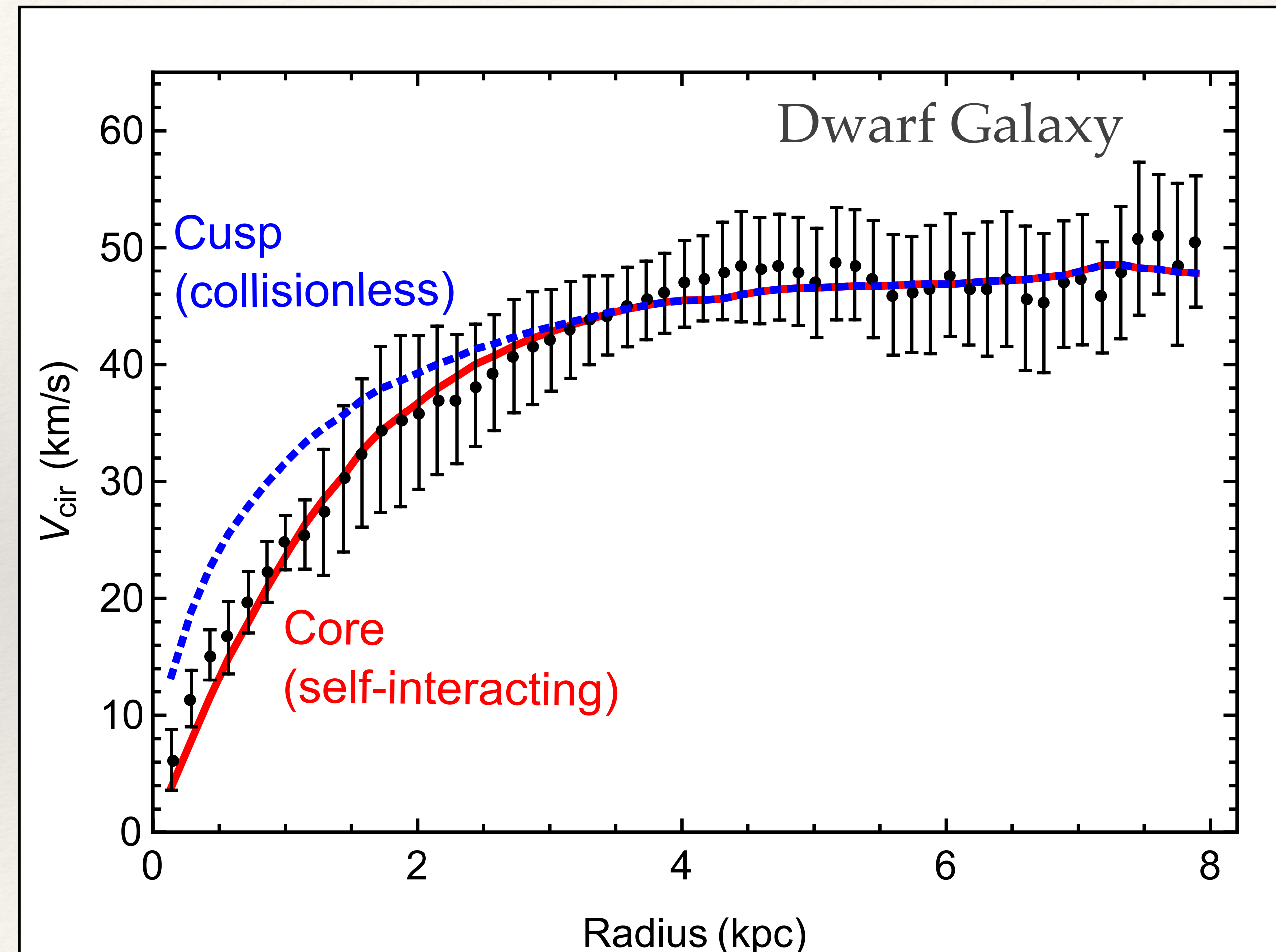
Self-interaction

- ❖ "Small structure problems"
- ❖ *Diversity, too-big-to-fail, missing satellites, cusp vs. core*



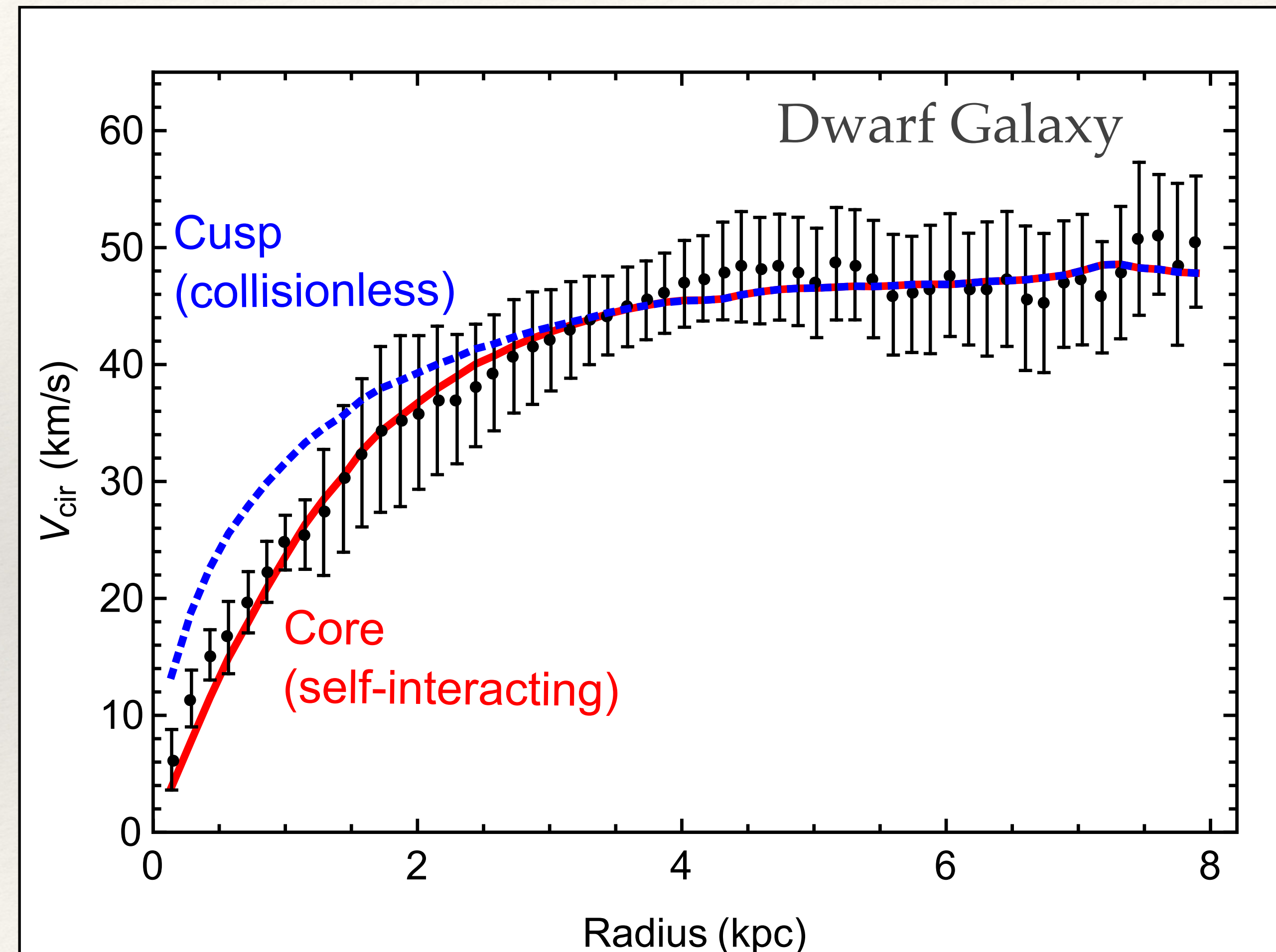
Self-interaction

- ❖ "Small structure problems"
 - ❖ *Diversity, too-big-to-fail, missing satellites, cusp vs. core*
- ❖ Core-like shape preferred
 - ❖ Hints towards self-interaction



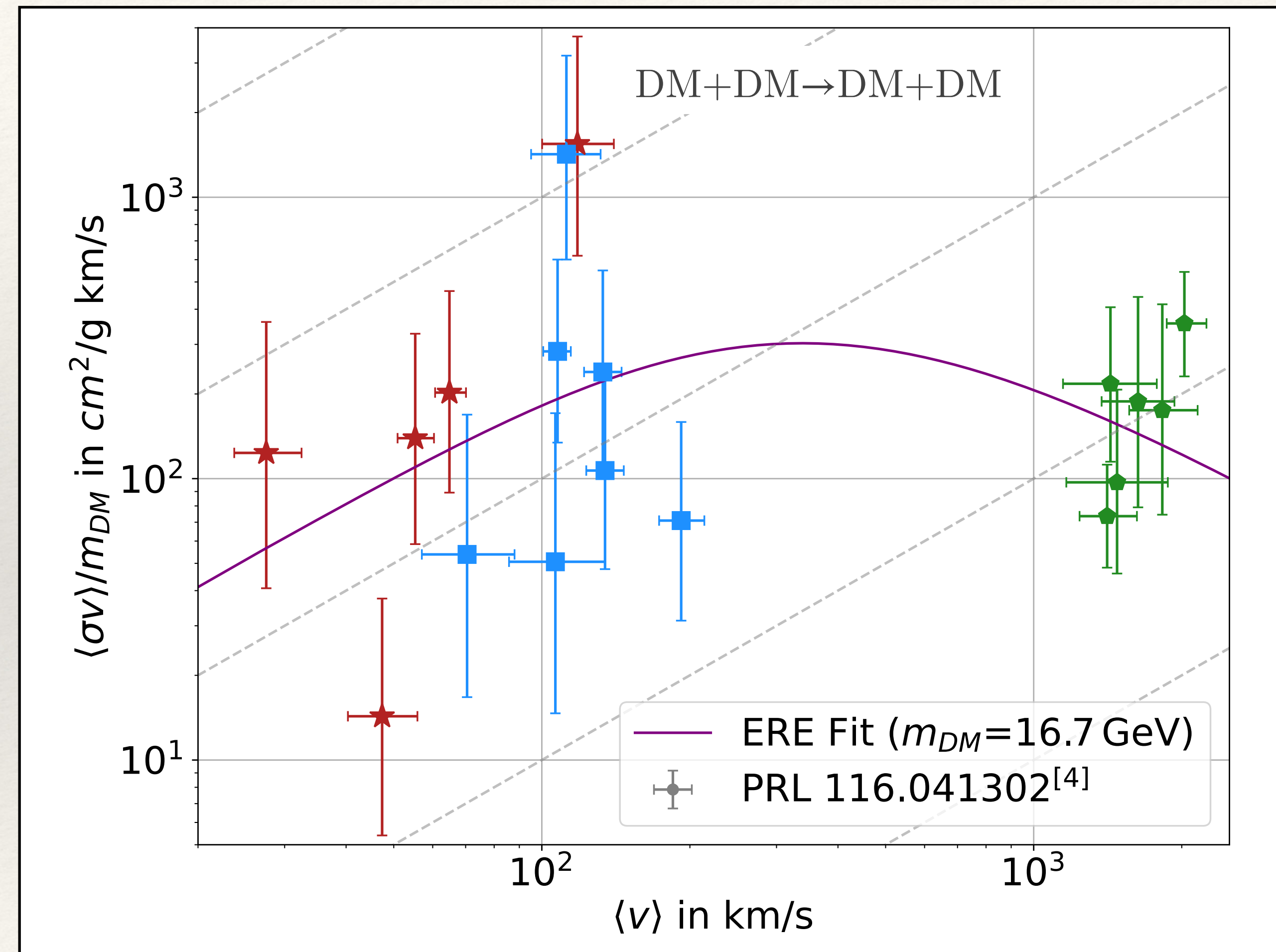
Self-interaction

- ❖ "Small structure problems"
 - ❖ *Diversity, too-big-to-fail, missing satellites, cusp vs. core*
- ❖ Core-like shape preferred
 - ❖ Hints towards self-interaction
- ❖ Upper bounds on cross-section from the bullet cluster



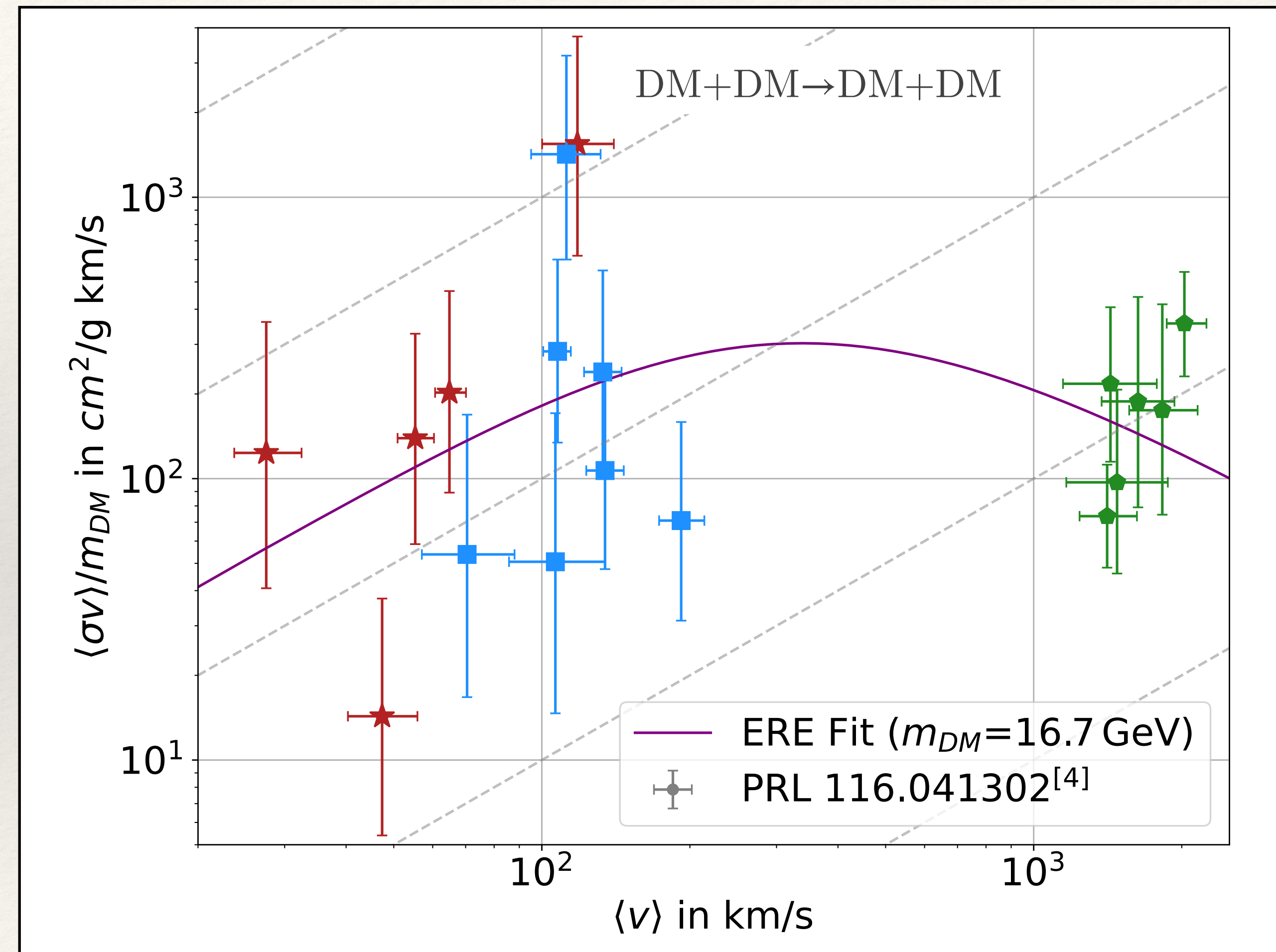
Velocity-dependent cross-section

- ❖ "Dark matter halos as particle colliders"
- ❖ Larger object $\rightarrow$ larger $\langle v_i \rangle$ & $\langle v \rangle$



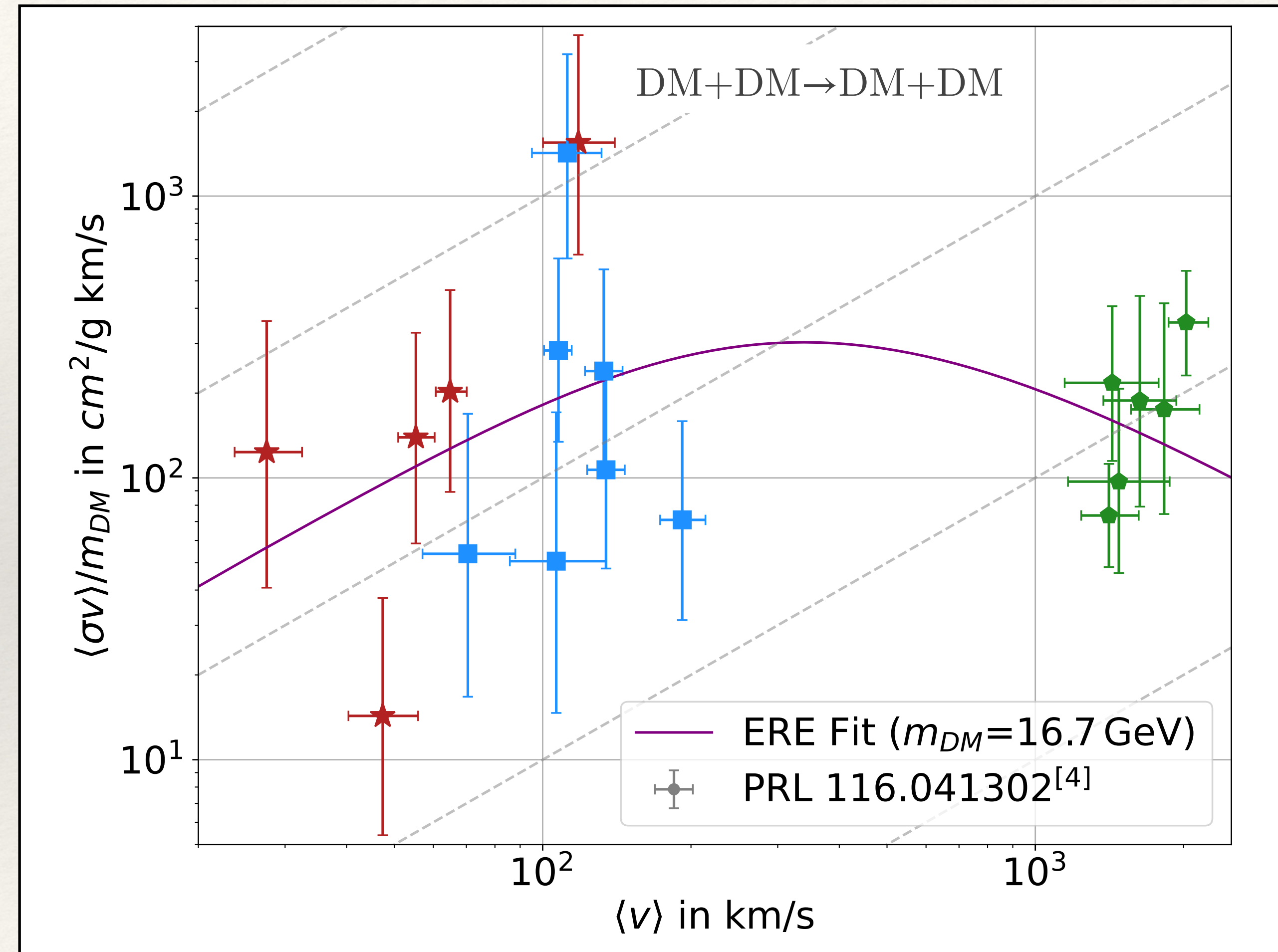
Velocity-dependent cross-section

- ❖ "Dark matter halos as particle colliders"
 - ❖ Larger object $\rightarrow$ larger $\langle v_i \rangle$ & $\langle v \rangle$
- ❖ Fit halo shape to simulations



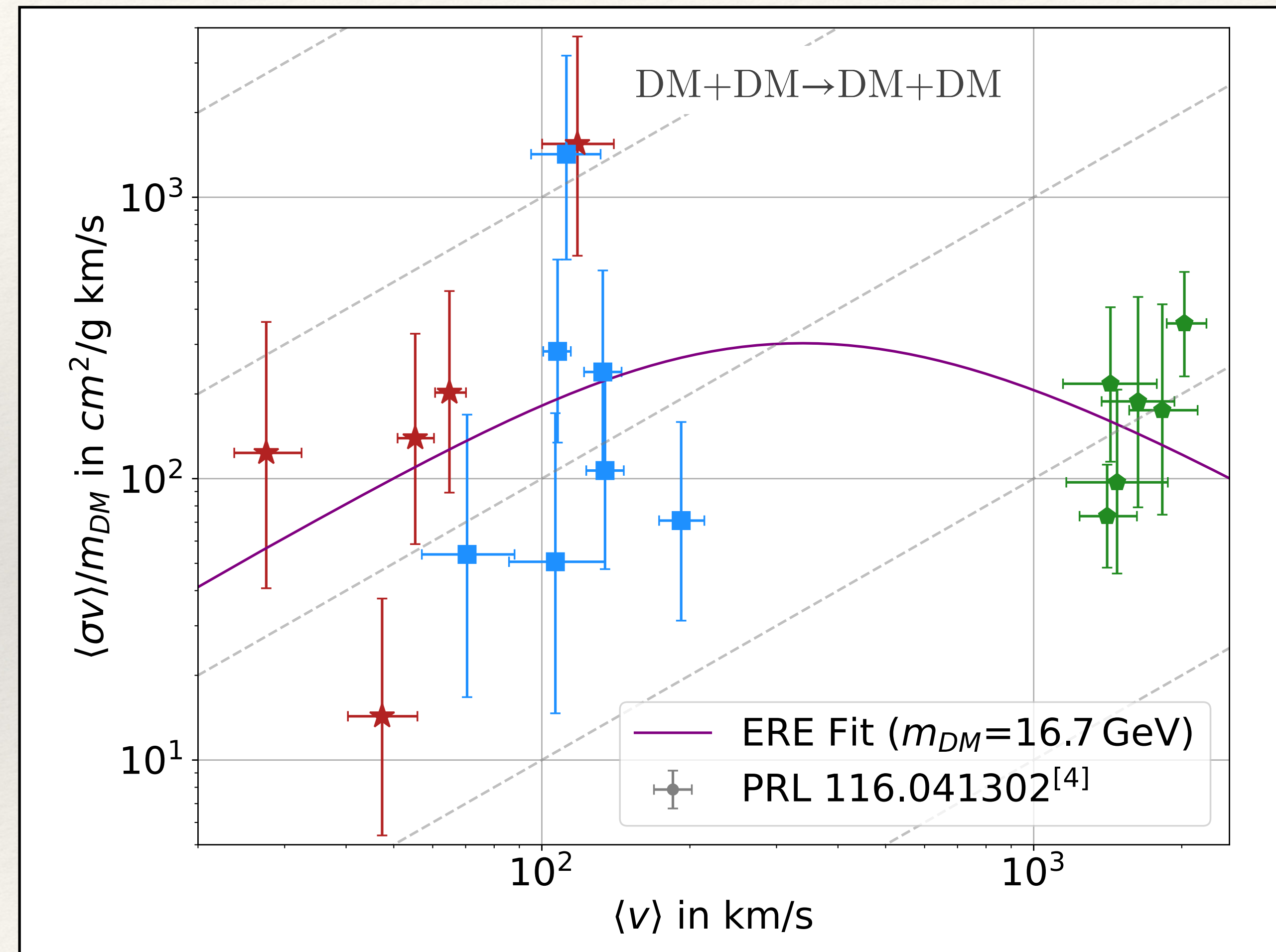
Velocity-dependent cross-section

- ❖ "Dark matter halos as particle colliders"
 - ❖ Larger object $\rightarrow$ larger $\langle v_i \rangle$ & $\langle v \rangle$
- ❖ Fit halo shape to simulations
- ❖ Mild velocity dependence @ non-relativistic velocities



Velocity-dependent cross-section

- ❖ "Dark matter halos as particle colliders"
 - ❖ Larger object $\rightarrow$ larger $\langle v_i \rangle$ & $\langle v \rangle$
- ❖ Fit halo shape to simulations
- ❖ Mild velocity dependence @ non-relativistic velocities
- ❖ Thermalized dark matter

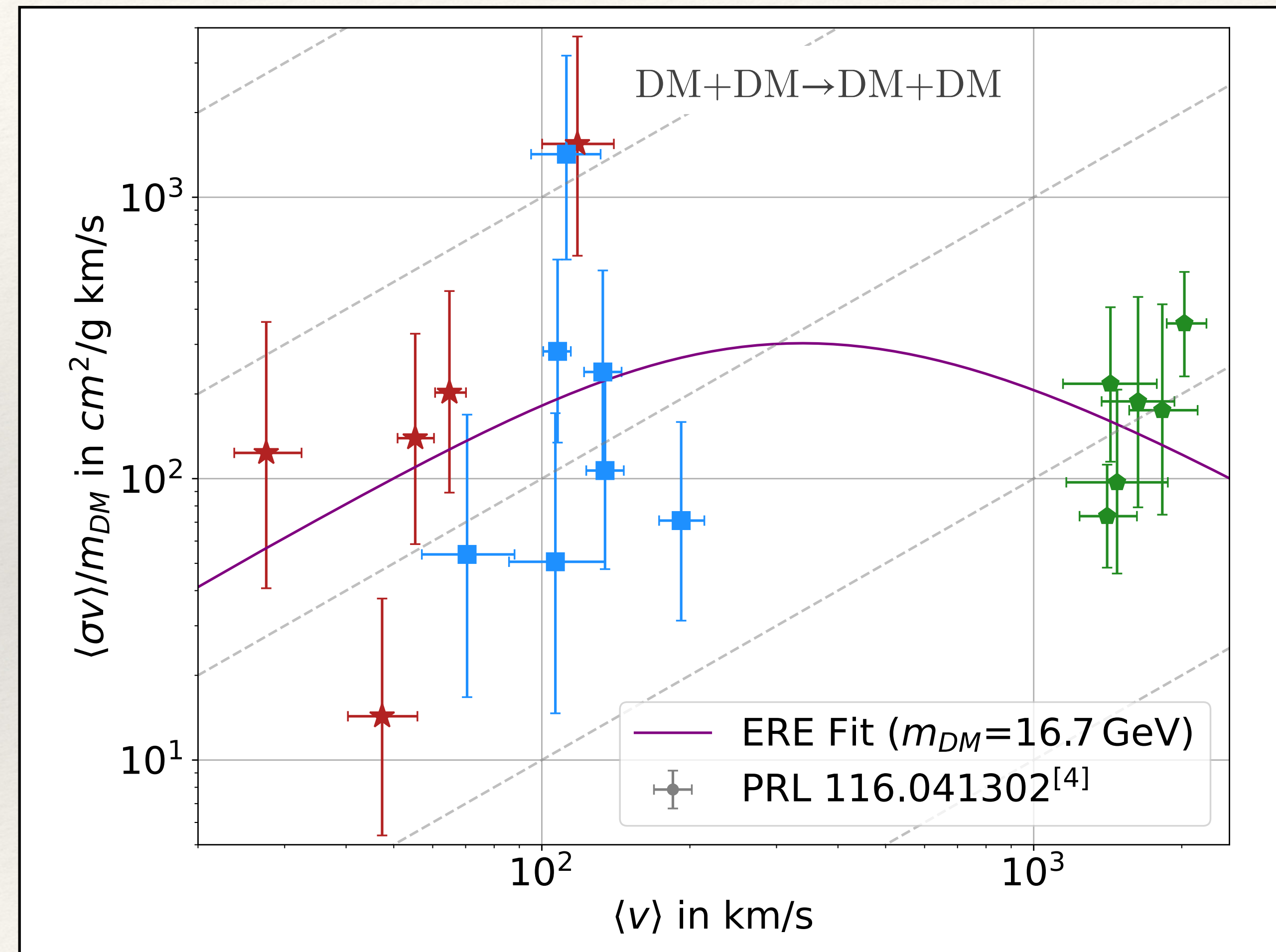


Velocity-dependent cross-section

- ❖ "Dark matter halos as particle colliders"
 - ❖ Larger object $\rightarrow$ larger $\langle v_i \rangle$ & $\langle v \rangle$
- ❖ Fit halo shape to simulations
- ❖ Mild velocity dependence @ non-relativistic velocities
- ❖ Thermalized dark matter
- ❖ Only $\sigma(s)$ needed

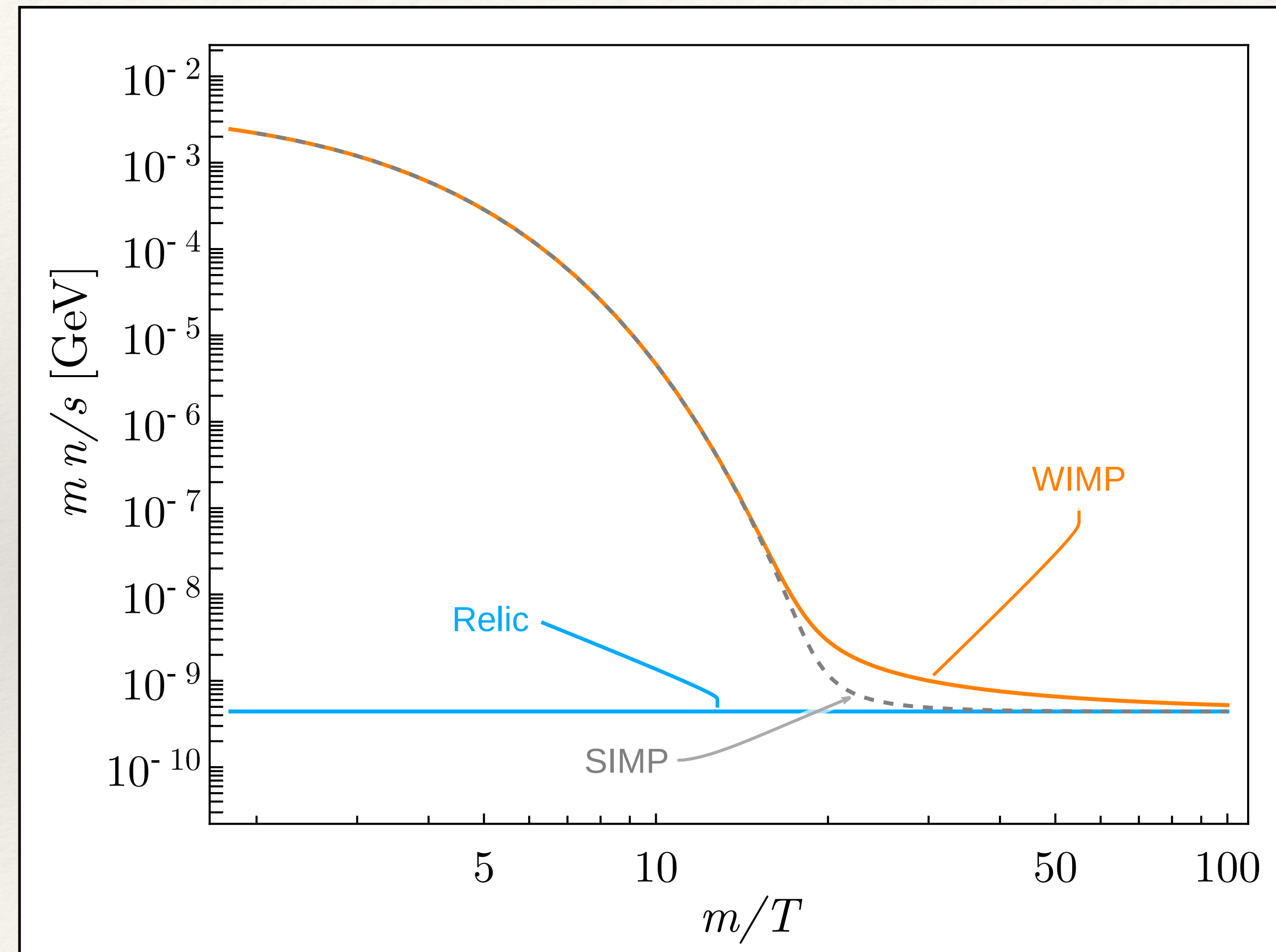


Can be done on the lattice



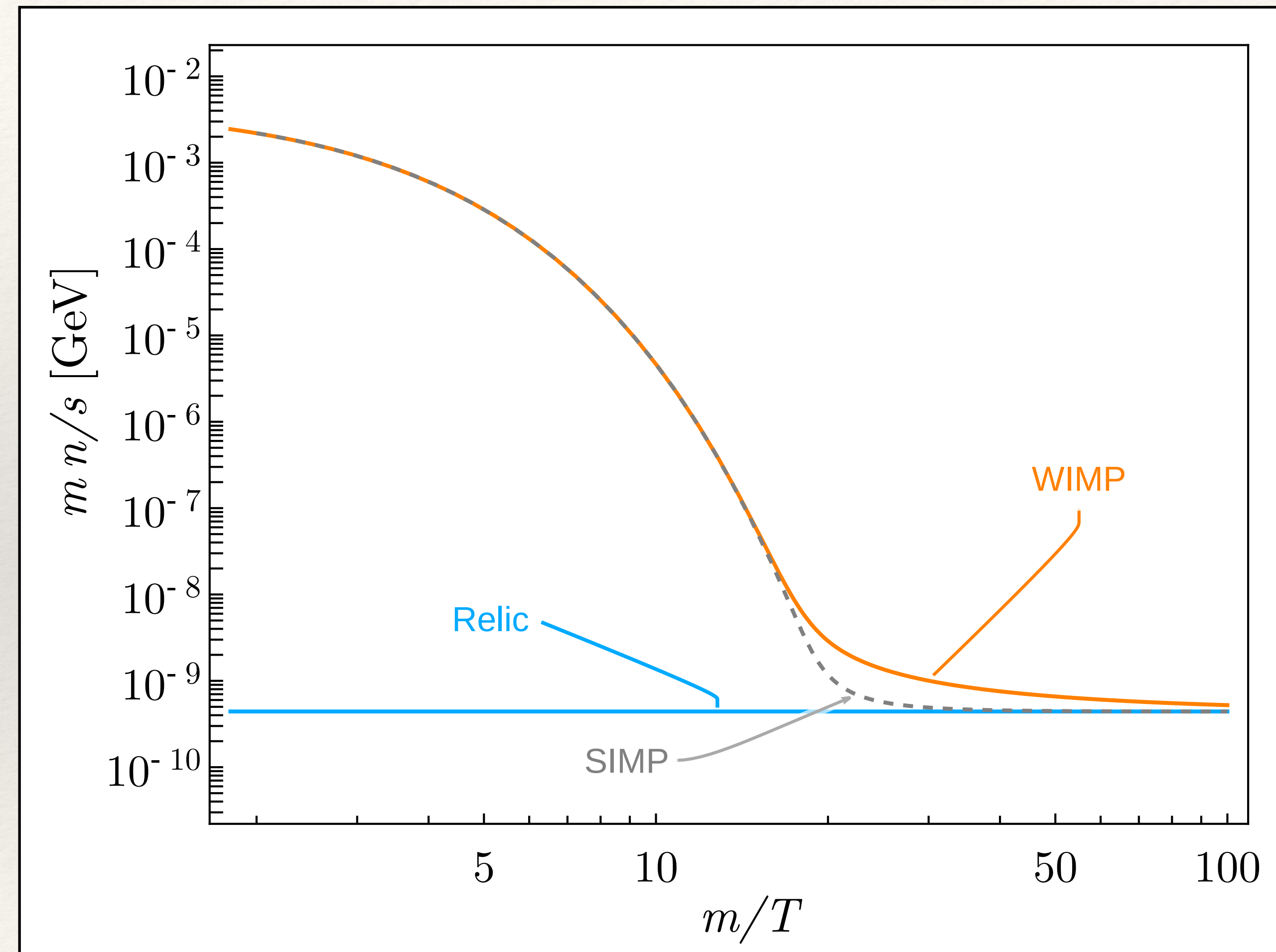
Relic density

- ❖ Possibility: Dark matter as a thermal relic from the early universe



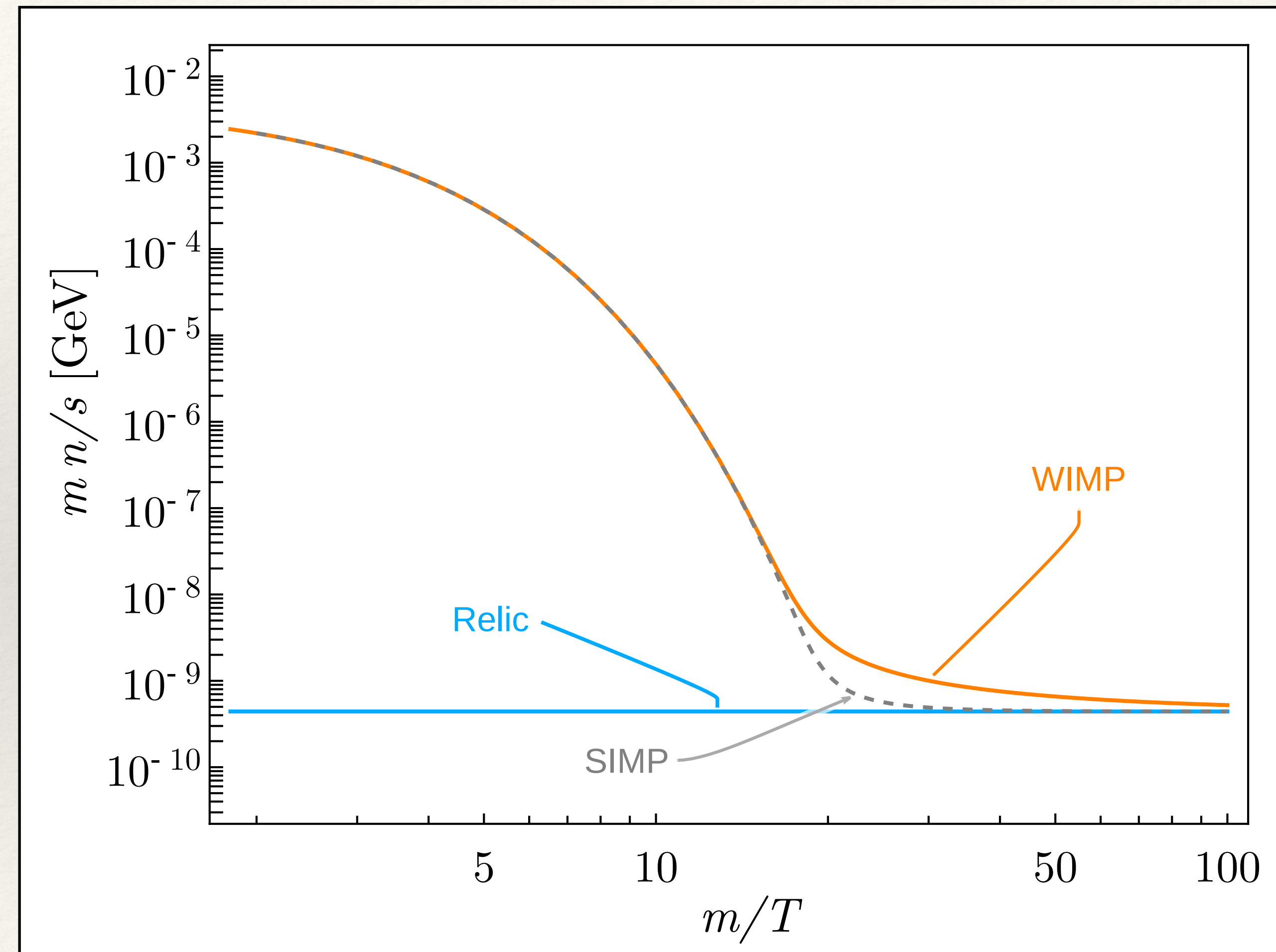
Relic density

- ❖ Possibility: Dark matter as a thermal relic from the early universe
- ❖ Handle on the dark matter abundance



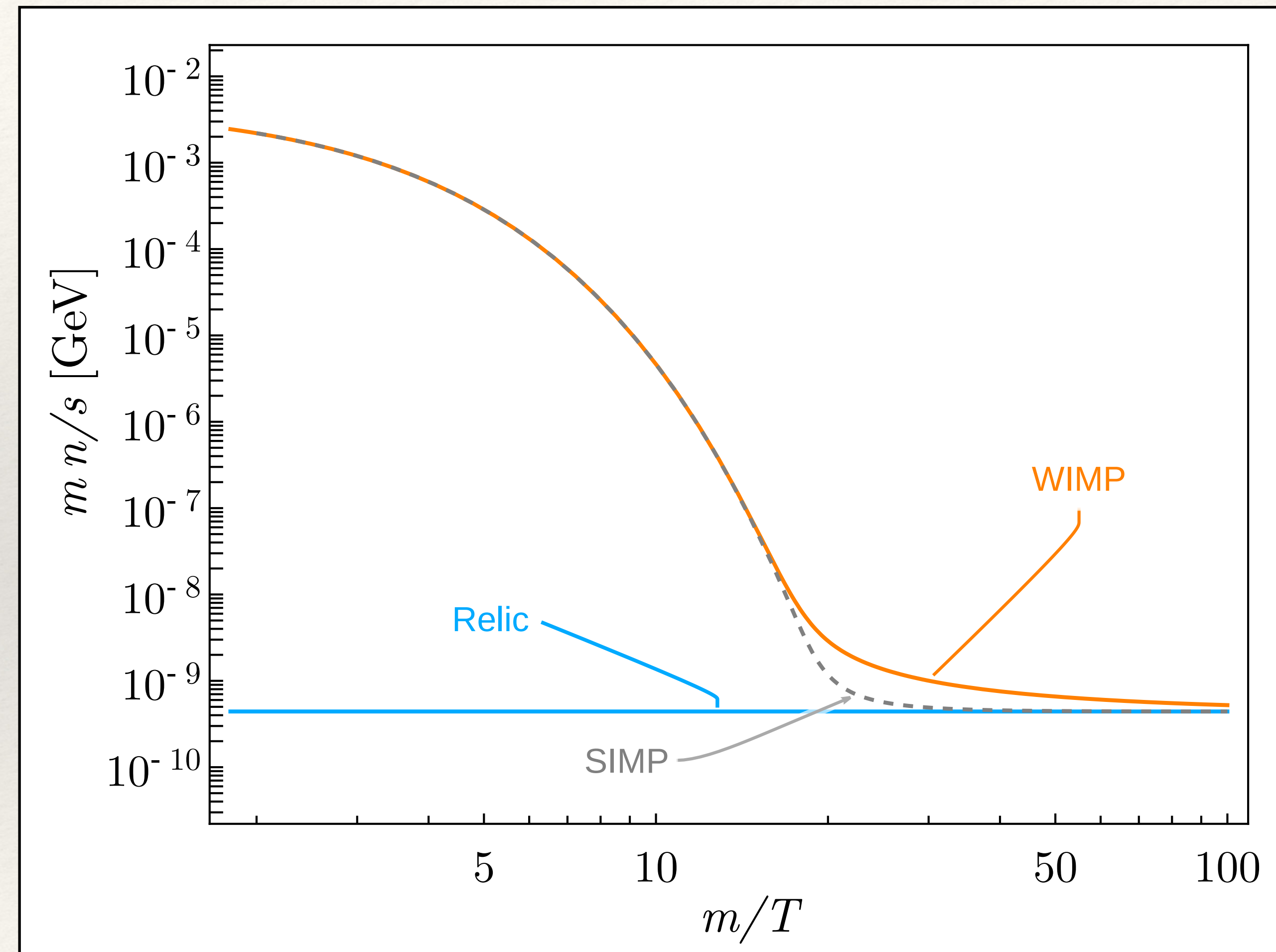
Relic density

- ❖ Possibility: Dark matter as a thermal relic from the early universe
- ❖ Handle on the dark matter abundance
- ❖ Solve Boltzmann equations
 - ❖ Temperature decreases $\rightarrow$ interaction "freezes out"



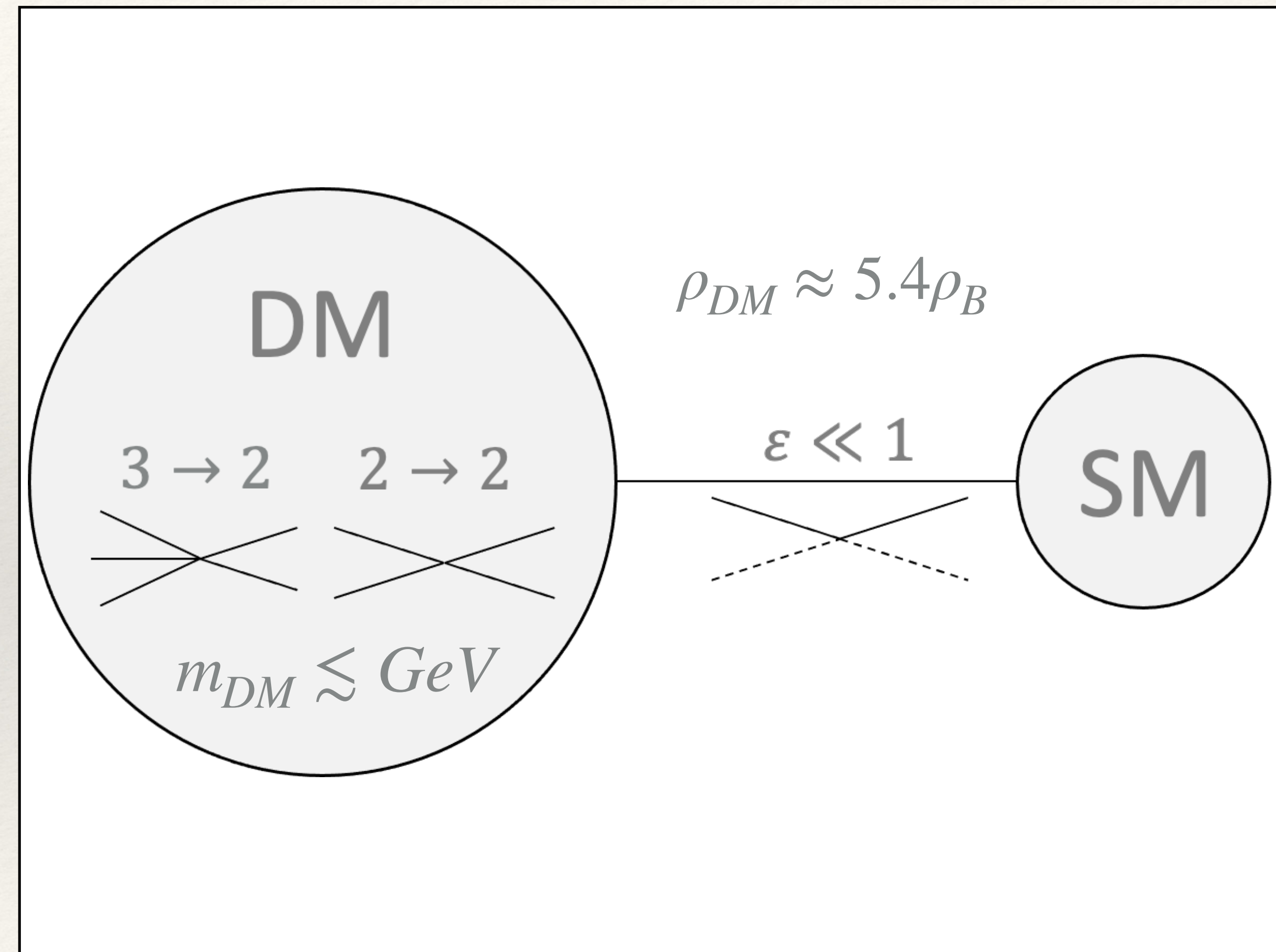
Relic density

- ❖ Possibility: Dark matter as a thermal relic from the early universe
- ❖ Handle on the dark matter abundance
- ❖ Solve Boltzmann equations
 - ❖ Temperature decreases $\rightarrow$ interaction "freezes out"
- ❖ Example:
 - ❖ WIMP: $\text{DM} + \text{DM} \rightarrow \text{SM} + \text{SM}$



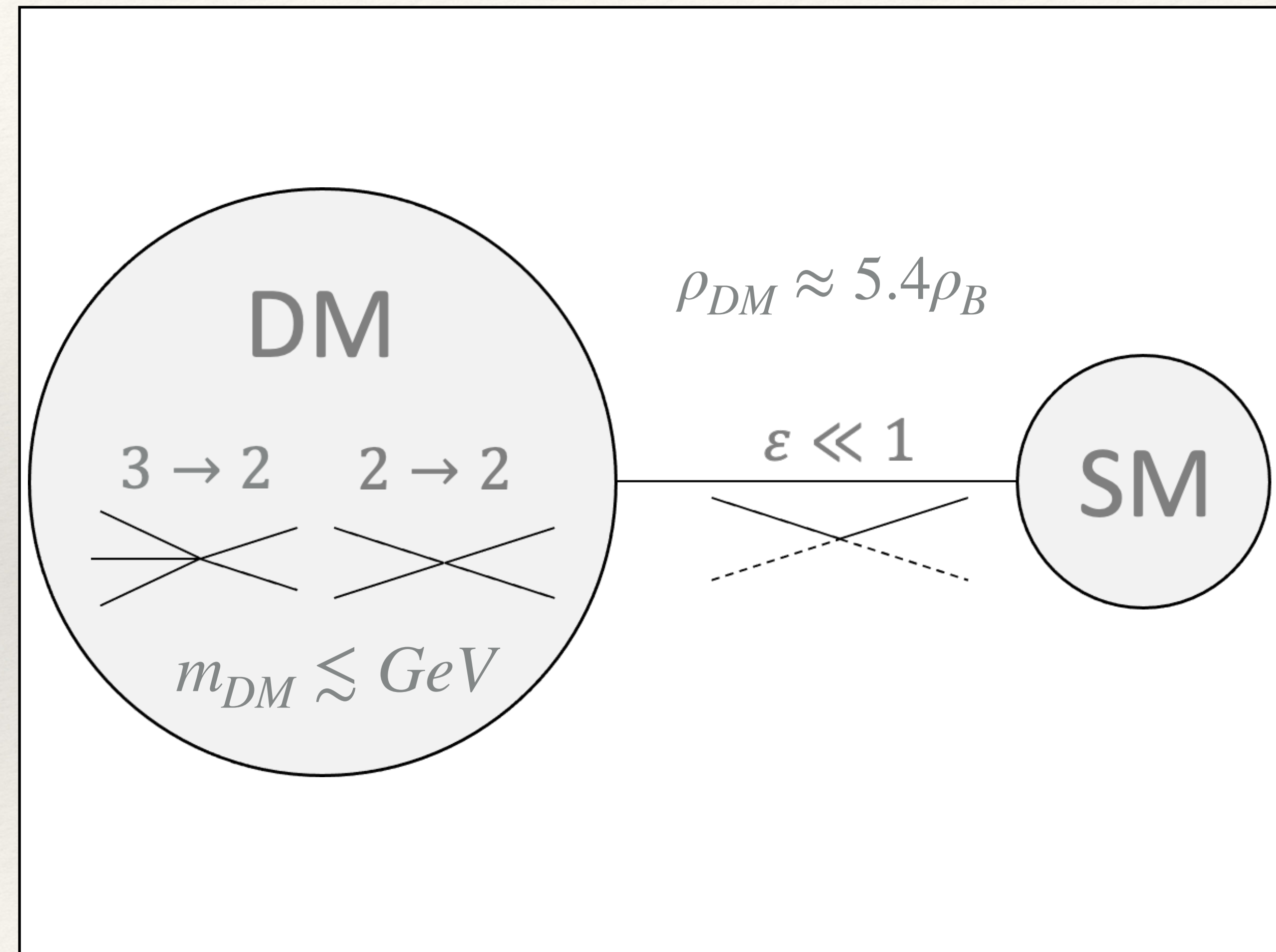
Strongly Interacting Massive Particles

- ❖ Alternative freeze-out paradigm



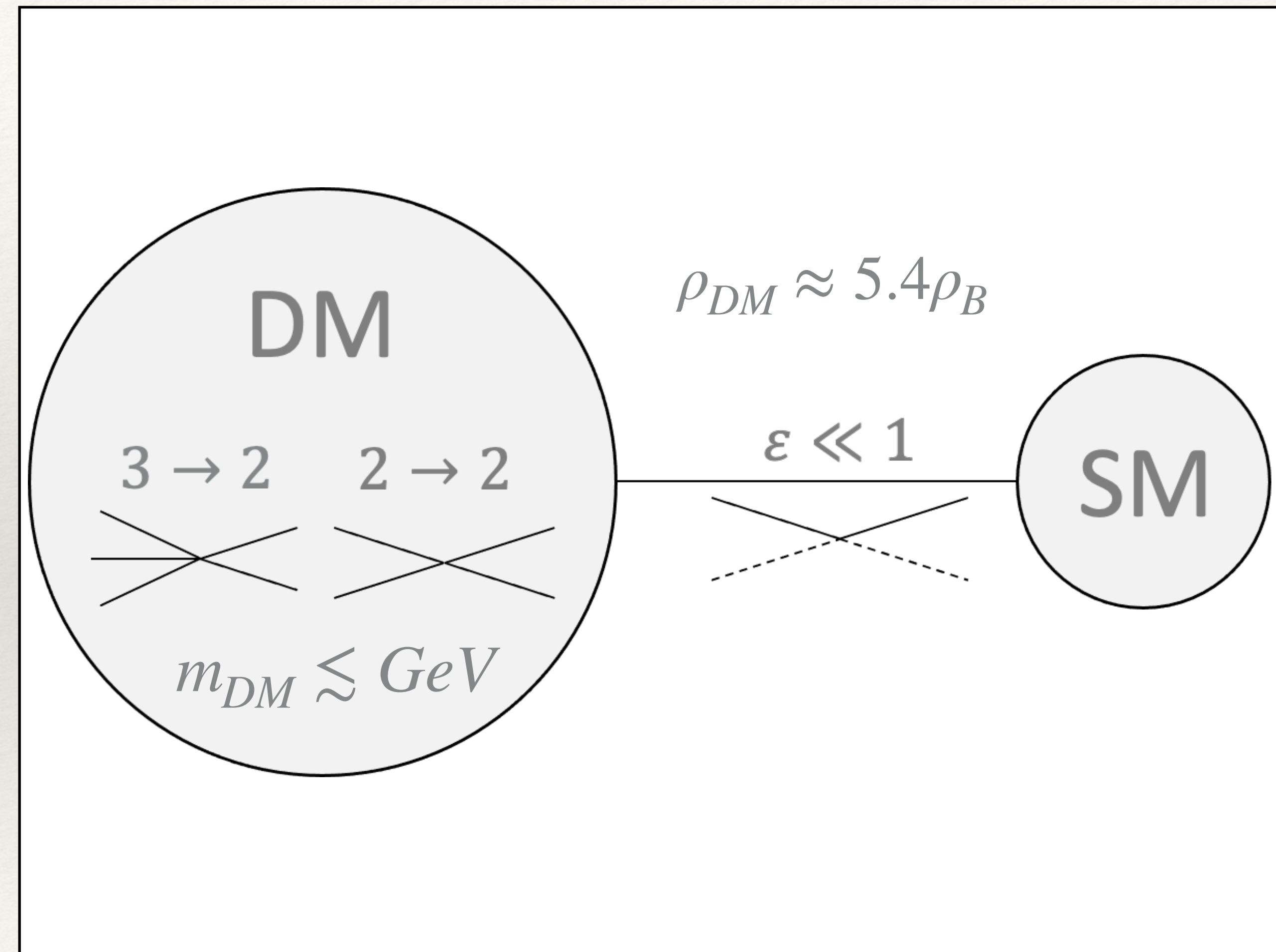
Strongly Interacting Massive Particles

- ❖ Alternative freeze-out paradigm
- ❖ Number lowering process in the dark sector
 - ❖ Addresses self-interaction



Strongly Interacting Massive Particles

- ❖ Alternative freeze-out paradigm
- ❖ Number lowering process in the dark sector
 - ❖ Addresses self-interaction
- ❖ Coupling to the SM sector needed to prevent heat-up
 - ❖ Mediator enables direct detection



UV realisation

- ❖ Strong coupling arises *naturally* in confining gauge theories

- ❖ $\mathcal{L} = -\frac{1}{2}F_{\mu\nu}F^{\mu\nu} + \bar{q}_i(i\gamma^\mu D_\mu - m_i)q_i$

UV realisation

- ❖ Strong coupling arises *naturally* in confining gauge theories

- ❖ $\mathcal{L} = -\frac{1}{2}F_{\mu\nu}F^{\mu\nu} + \bar{q}_i(i\gamma^\mu D_\mu - m_i)q_i$

- ❖ Symmetry depends on representation
 - ❖ Fundamental, adjoint, antisymmetric, ...

UV realisation

- ❖ Strong coupling arises *naturally* in confining gauge theories

- $$\mathcal{L} = -\frac{1}{2}F_{\mu\nu}F^{\mu\nu} + \bar{q}_i(i\gamma^\mu D_\mu - m_i)q_i$$

- ❖ Symmetry depends on representation

- ❖ Fundamental, adjoint, antisymmetric, ...

Symmetry of the UV Lagrangian

Representation of gauge group	Flavour symmetry
Complex	U(2)xU(2)
Real	U(4)
Pseudoreal	U(4)

UV realisation

- ❖ Strong coupling arises *naturally* in confining gauge theories

- $$\mathcal{L} = -\frac{1}{2}F_{\mu\nu}F^{\mu\nu} + \bar{q}_i(i\gamma^\mu D_\mu - m_i)q_i$$

- ❖ Symmetry depends on representation

- ❖ Fundamental, adjoint, antisymmetric, ...

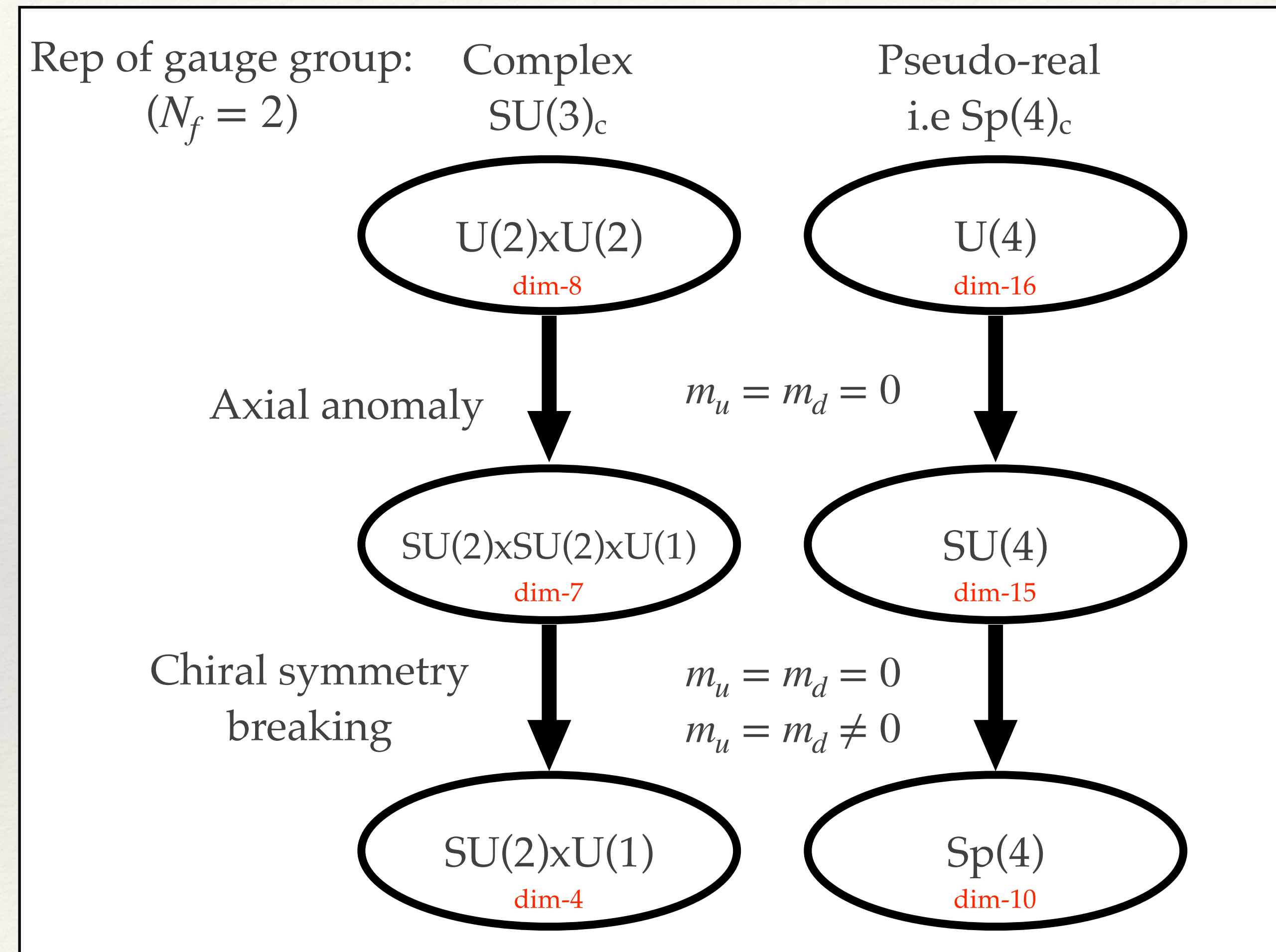
- ❖ Also different breaking patterns

Symmetry of the UV Lagrangian

Representation of gauge group	Flavour symmetry
Complex	U(2)xU(2)
Real	U(4)
Pseudoreal	U(4)

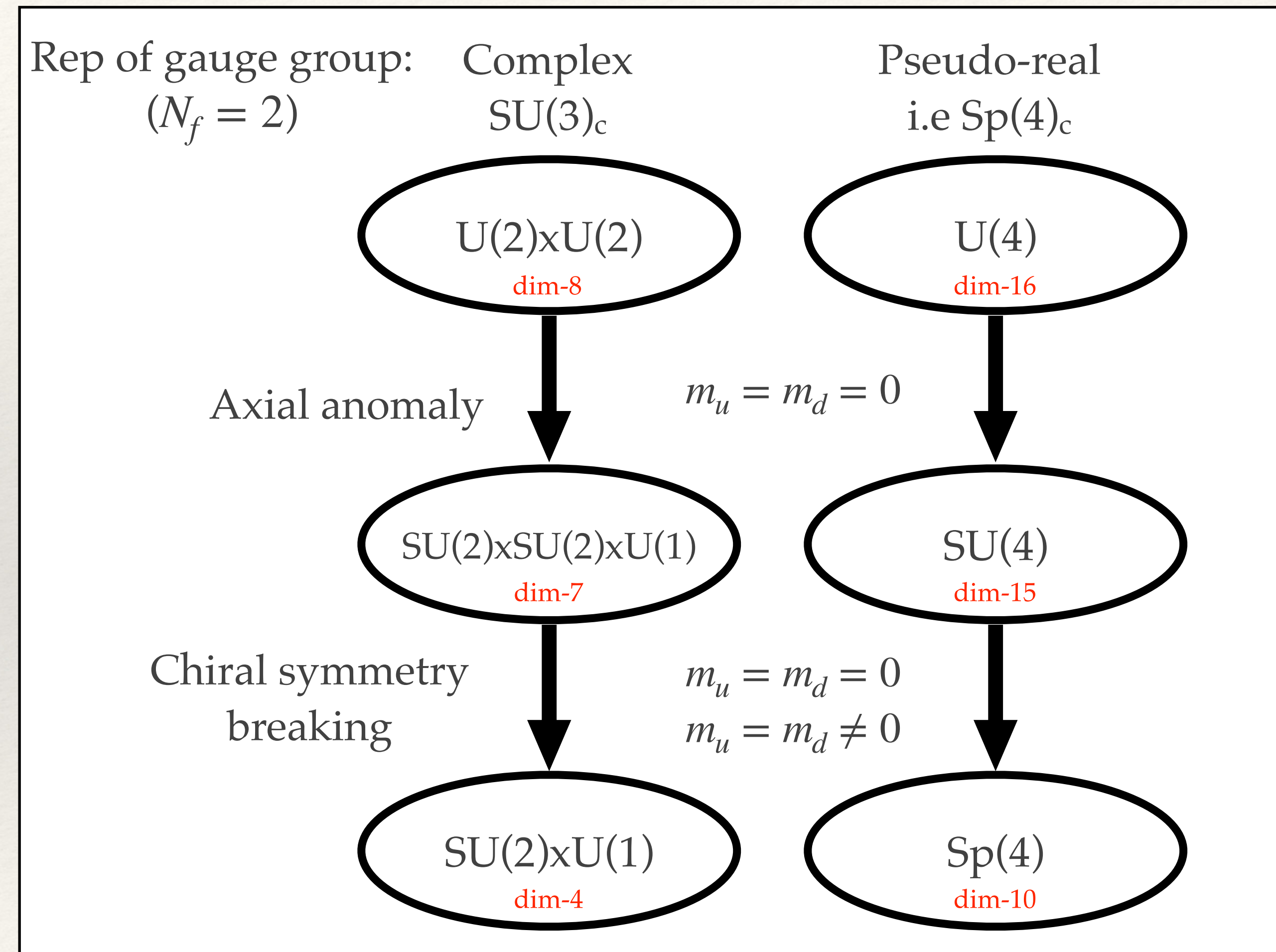
Minimal realisation

- ❖ Pseudo-real rep of gauge group with $N_f = 2$
- ❖ $Sp(4)$ flavour symmetry



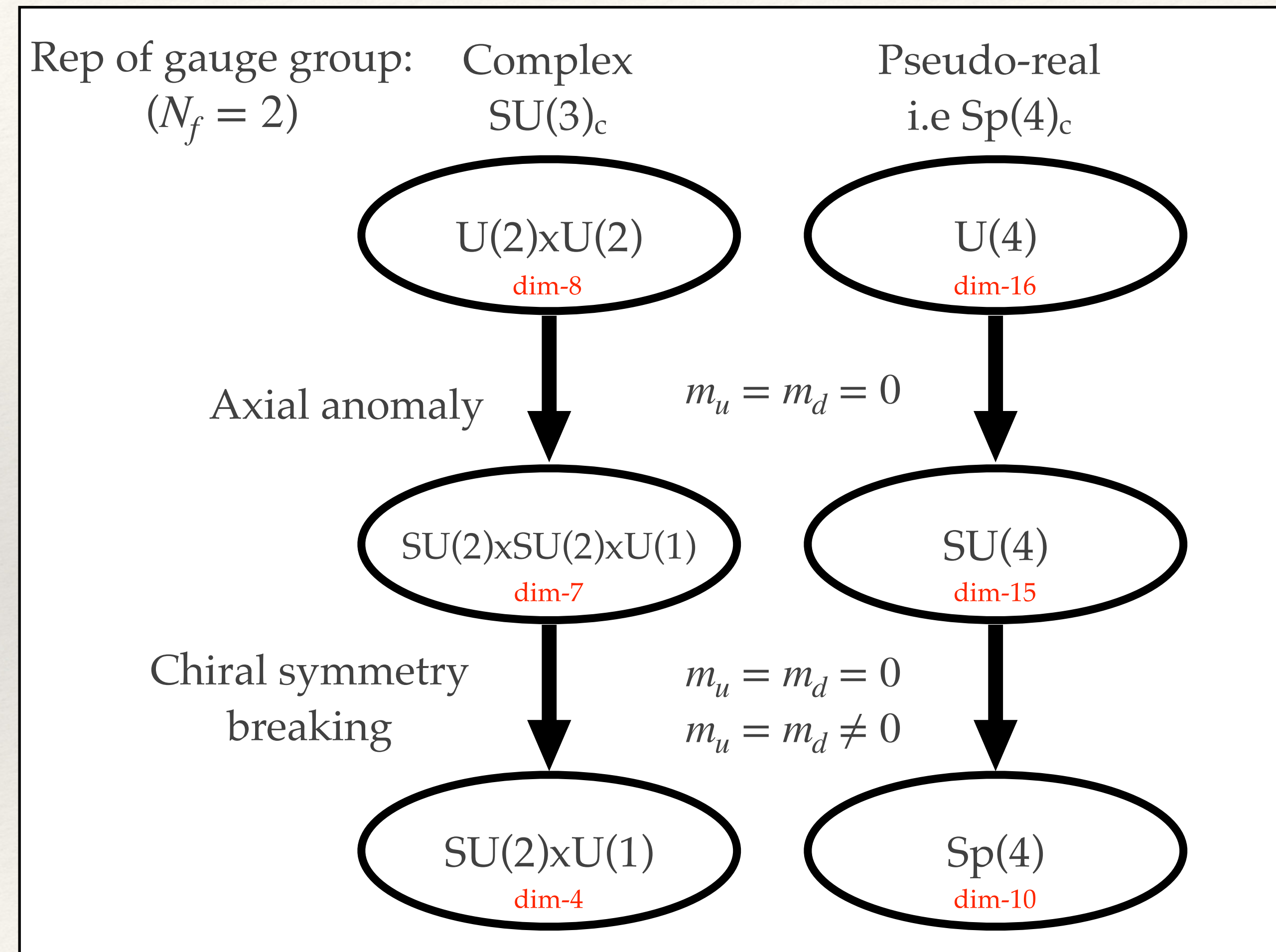
Minimal realisation

- ❖ Pseudo-real rep of gauge group with $N_f = 2$
 - ❖ $Sp(4)$ flavour symmetry
- ❖ Mixing of left- and right handed components (Weyl-fermions)
 - ❖ Symmetry is enlarged



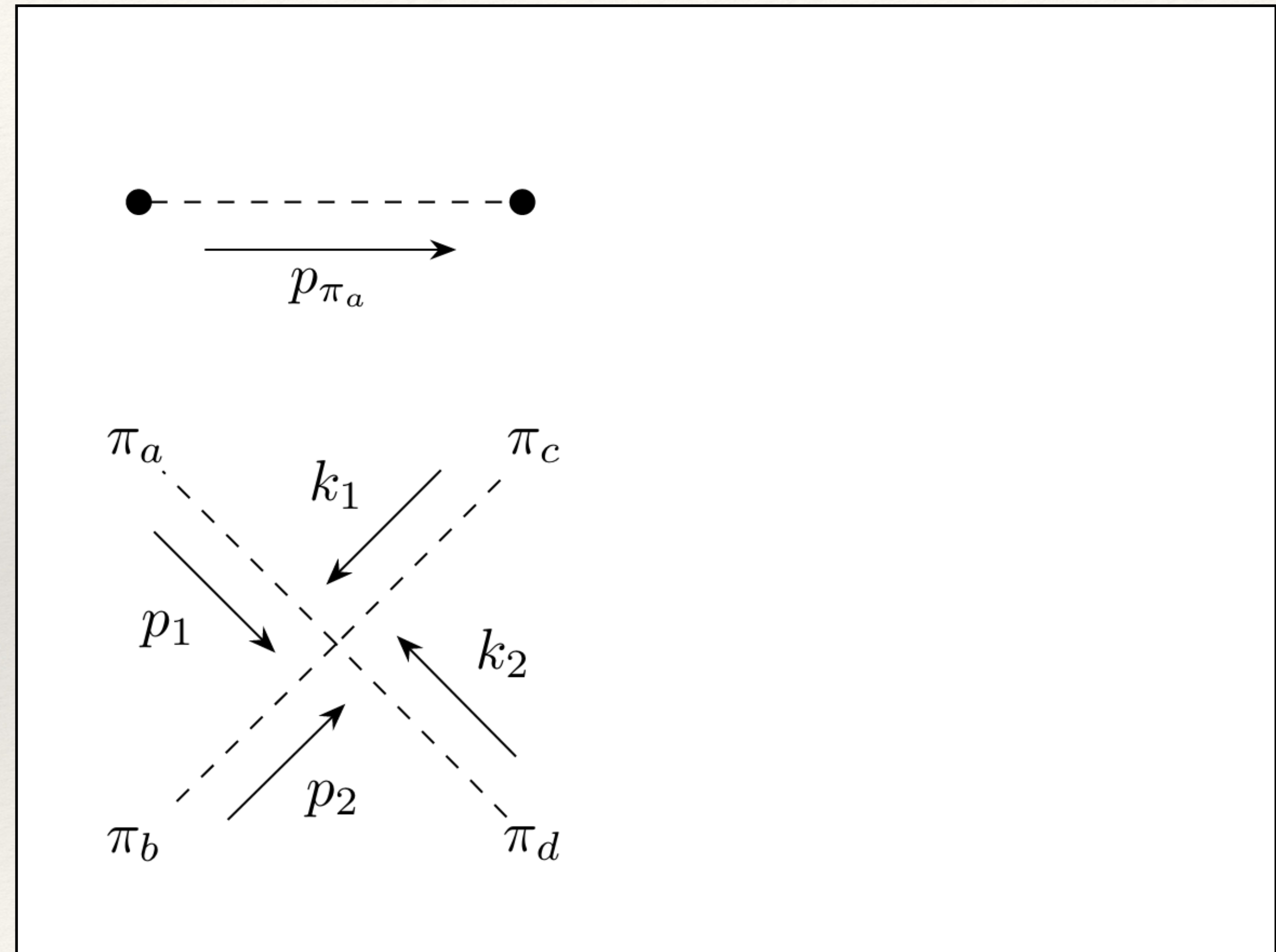
Minimal realisation

- ❖ Pseudo-real rep of gauge group with $N_f = 2$
 - ❖ $Sp(4)$ flavour symmetry
- ❖ Mixing of left- and right handed components (Weyl-fermions)
 - ❖ Symmetry is enlarged
- ❖ Result: 5 pNGBs
 - ❖ $3 \rightarrow 2$ process possible
 - ❖ WZW description in ChPT



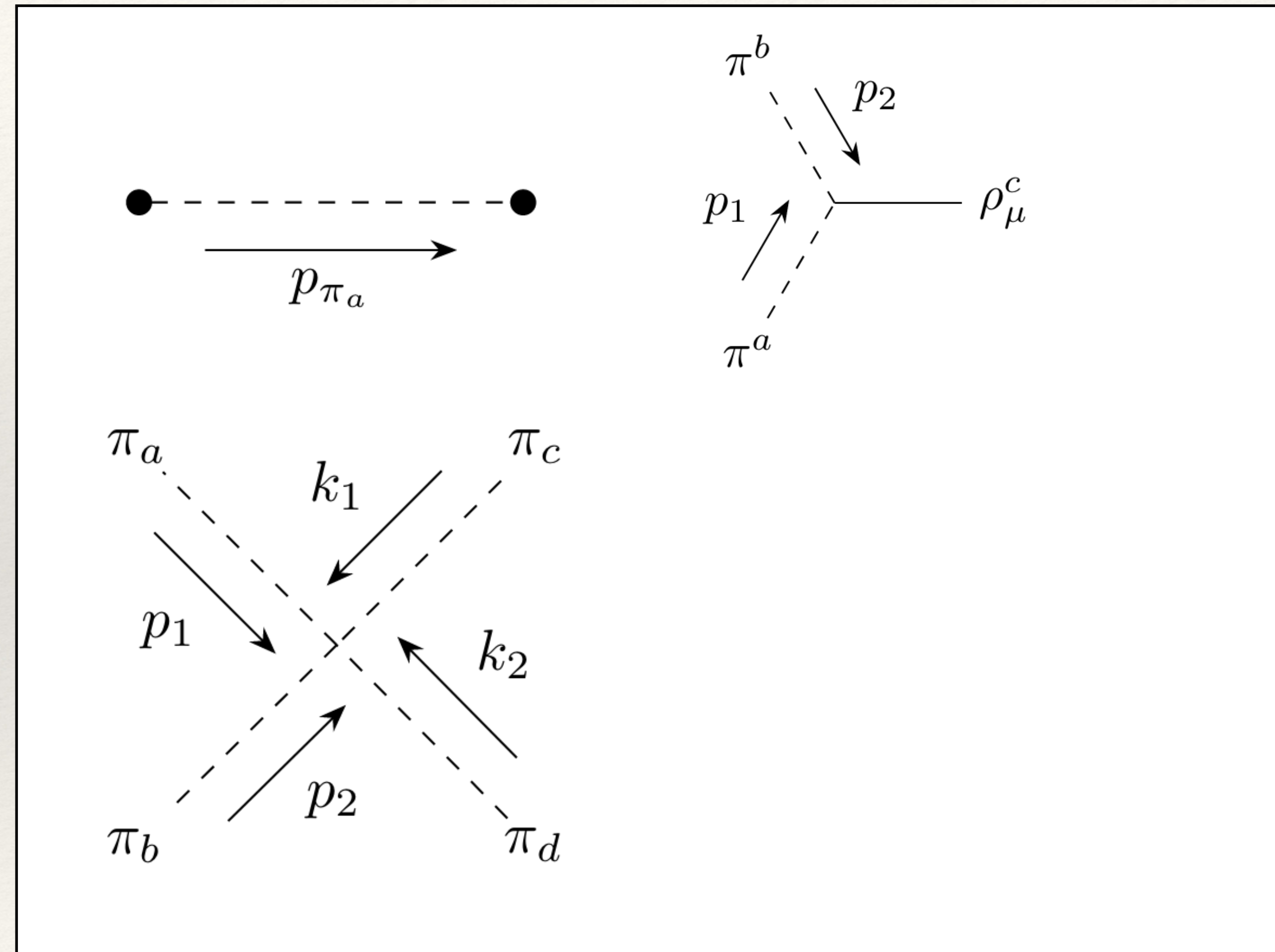
$Sp(4)$ ChPT

- ❖ Effective description in terms of 5 *dark Pions*



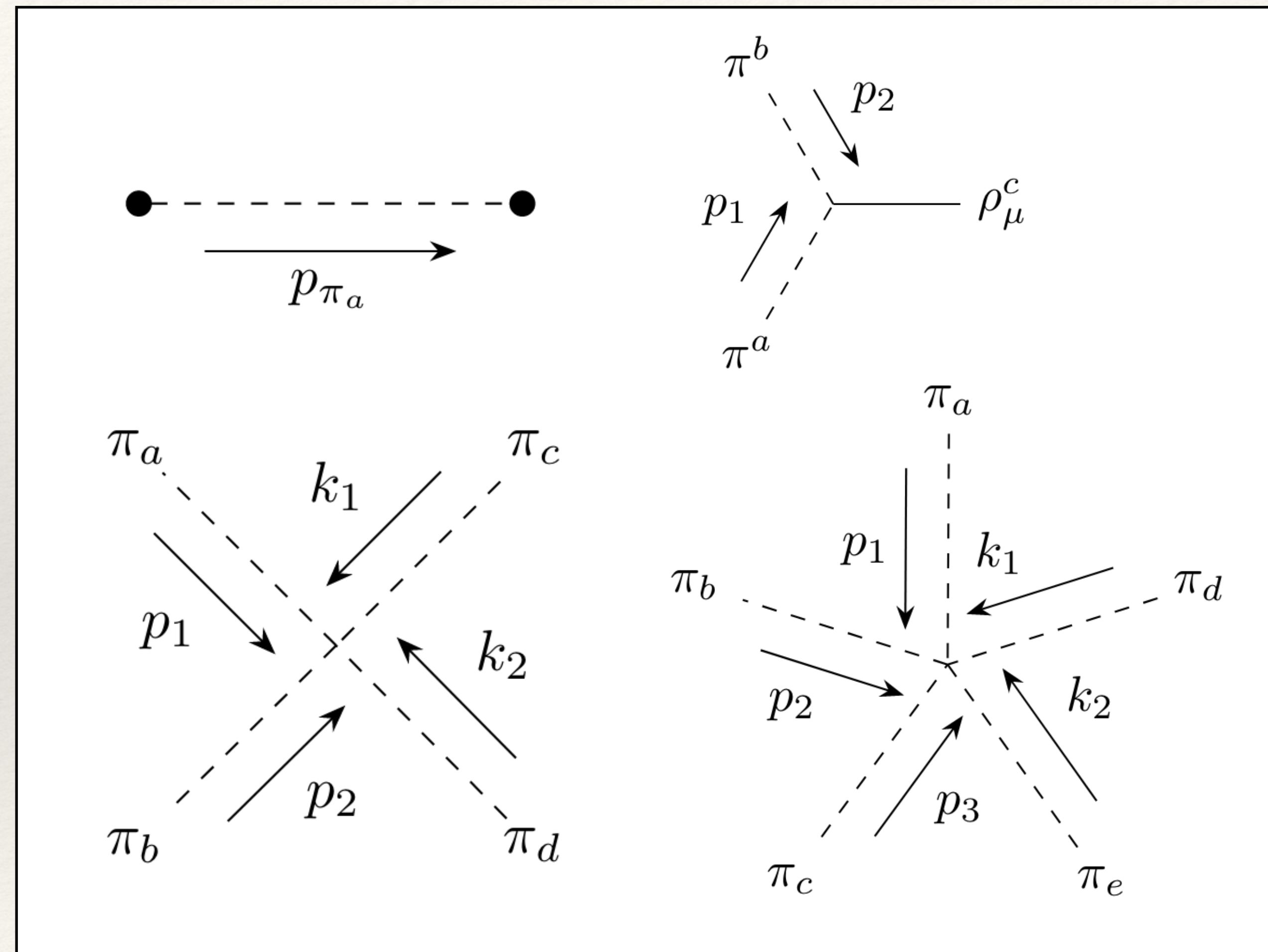
Sp(4) ChPT

- ❖ Effective description in terms of 5 *dark Pions*
- ❖ Include a vector particle and a mediator to the standard model



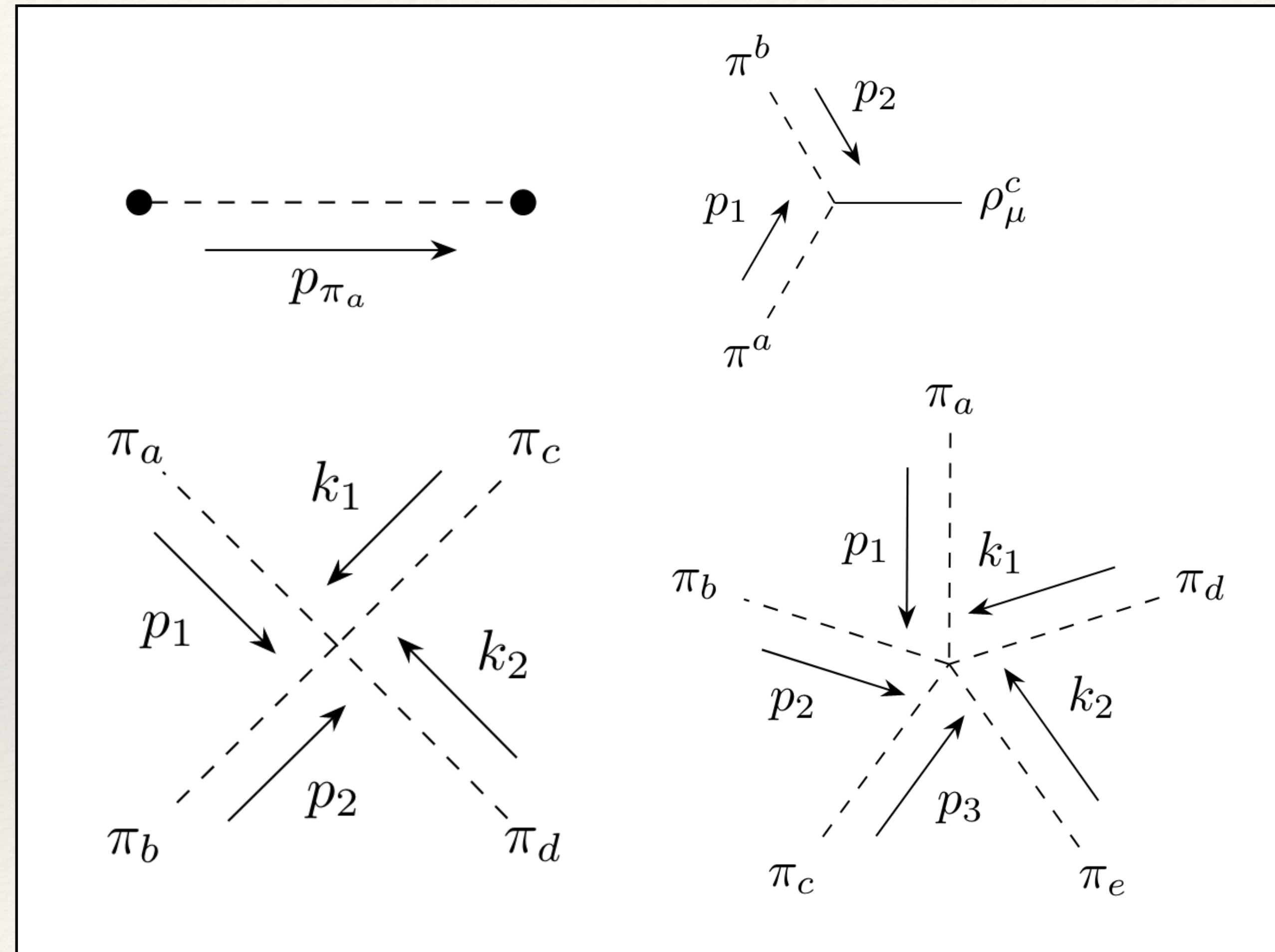
Sp(4) ChPT

- ❖ Effective description in terms of 5 *dark Pions*
- ❖ Include a vector particle and a mediator to the standard model
- ❖ Include $3 \rightarrow 2$ via Wess-Zumino-Witten term

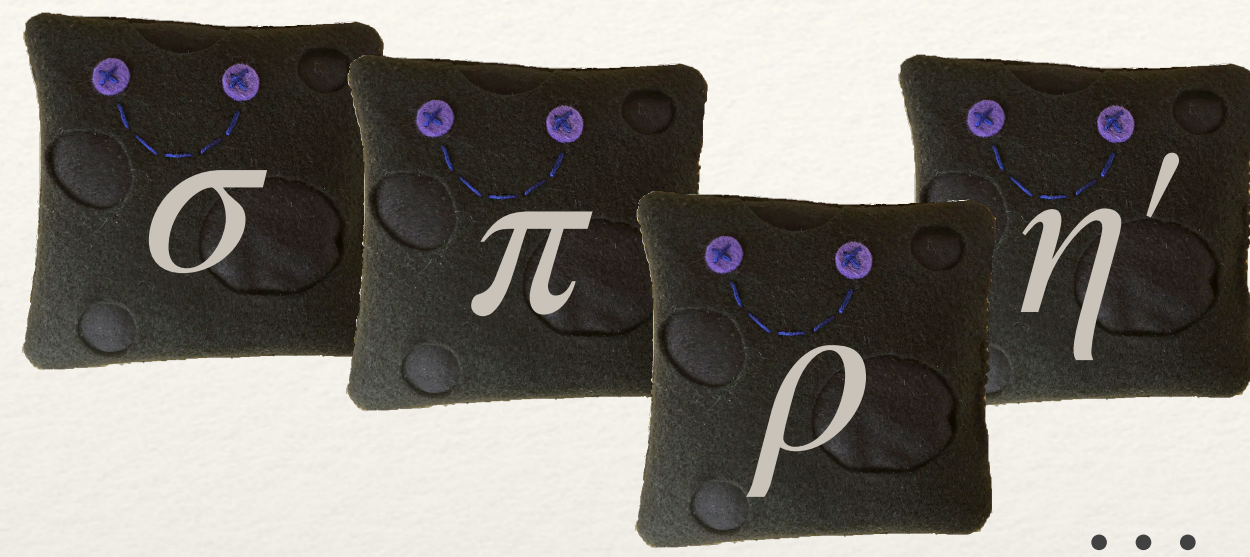


Sp(4) ChPT

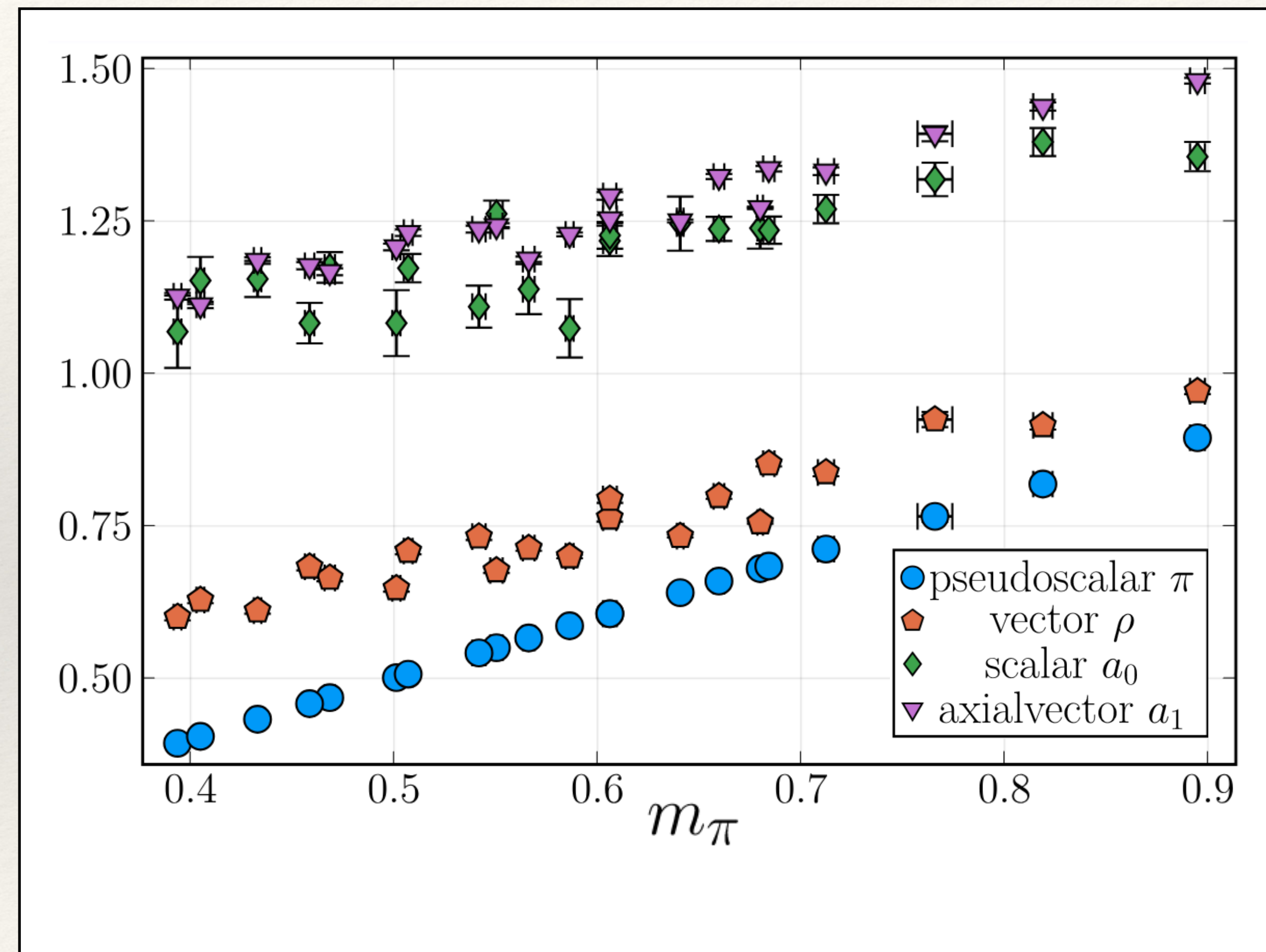
- ❖ Effective description in terms of 5 *dark Pions*
- ❖ Include a vector particle and a mediator to the standard model
- ❖ Include $3 \rightarrow 2$ via Wess-Zumino-Witten term
- ❖ Relies on low energy constants
 - ❖ Masses, scattering length, ...



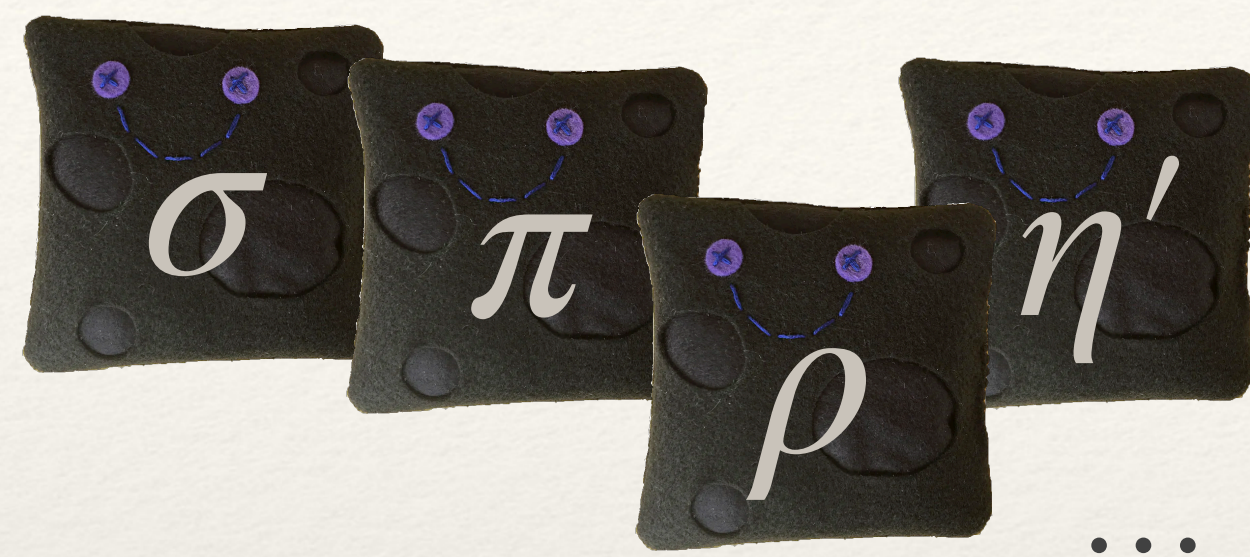
Particle phenomenology



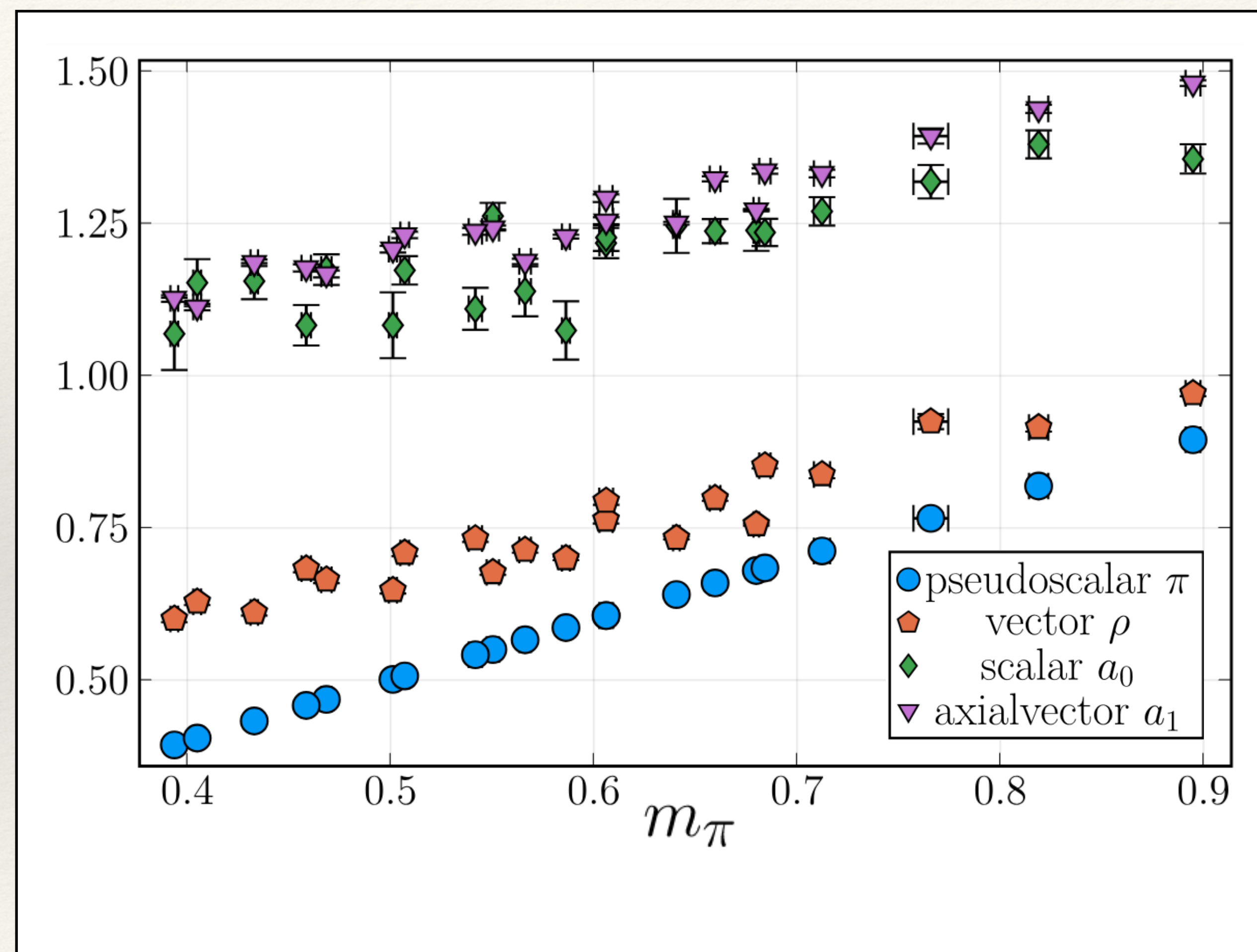
❖ Zoo of dark hadrons



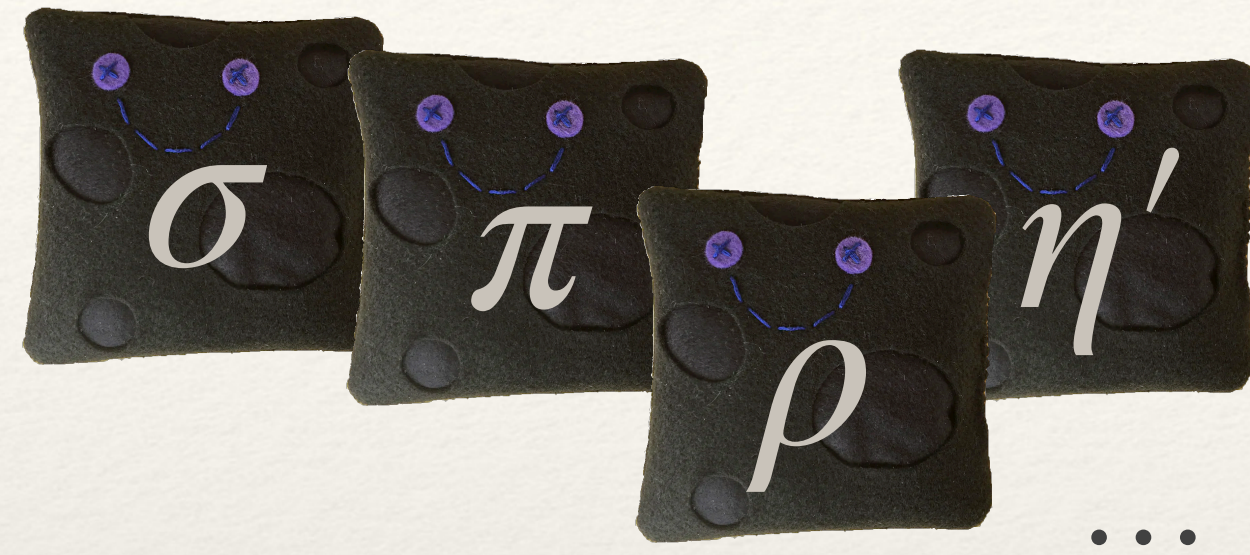
Particle phenomenology



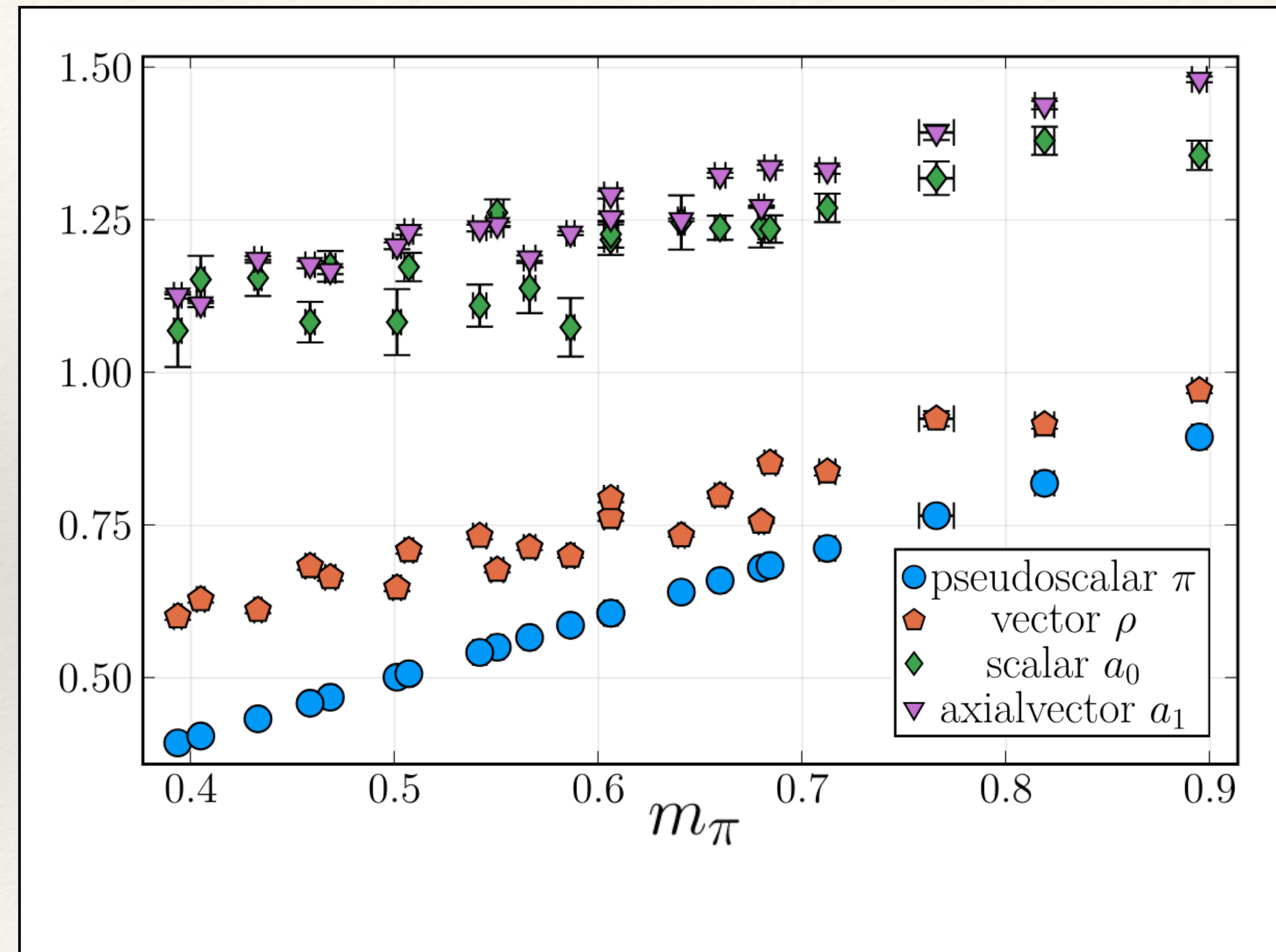
- ❖ Zoo of dark hadrons
- ❖ 5 Pions & 10 Rhos lightest non-singlets



Particle phenomenology



- ❖ Zoo of dark hadrons
- ❖ 5 Pions & 10 Rhos lightest non-singlets
- ❖ No fermionic bound states



Scattering phenomenology

- ❖ 14-dim:
- ❖ (Probably) contributes most to $\pi\pi$ -scattering
- ❖ 14 out of 25 possible combinations of Pions

$$Sp(4)_f$$

$$5 \otimes 5 = 1 \oplus 10 \oplus 14$$

$$\pi\pi \rightarrow \pi\pi \text{ (I=0,1,2)}$$

$$\pi\pi \rightarrow \rho \text{ (I=1)}$$

$$\pi\pi \rightarrow \pi\pi\pi \text{ (I=1)}$$

etc.

Scattering phenomenology

- ❖ 1-dim:
- ❖ (Probably) no large contribution to $\pi\pi$ -scattering
- ❖ Mixes in other scattering channel
- ❖ Numerically challenging

$$Sp(4)_f$$

$$5 \otimes 5 = \mathbf{1} \oplus 10 \oplus 14$$

$$\pi\pi \rightarrow \pi\pi \text{ (I=0,1,2)}$$

$$\pi\pi \rightarrow \rho \text{ (I=1)}$$

$$\pi\pi \rightarrow \pi\pi\pi \text{ (I=1)}$$

etc.

Scattering phenomenology

- ❖ 10-dim:
 - ❖ Mixing with the Rho
 - ❖ $\pi\pi\pi \rightarrow \pi\pi$
- ❖ Work in progress

$$Sp(4)_f$$

$$5 \otimes 5 = 1 \oplus \mathbf{10} \oplus 14$$

$$\pi\pi \rightarrow \pi\pi \text{ (I=0,1,2)}$$

$$\pi\pi \rightarrow \rho \text{ (I=1)}$$

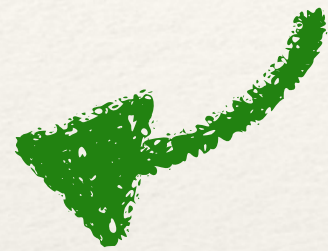
$$\pi\pi \rightarrow \pi\pi\pi \text{ (I=1)}$$

etc.

Phenomenology of scattering channels

Done 

❖ 14-dim:



❖ Makes up most $\pi\pi$ scattering (14/25)

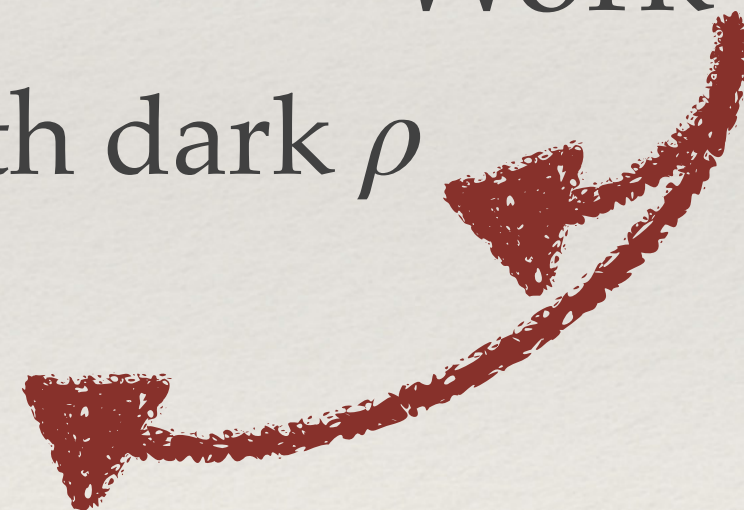
❖ Easiest on the lattice

❖ 10-dim:

Work in progress

❖ Mixing with dark ρ

❖ $\pi\pi\pi \rightarrow \pi\pi$



❖ 1-dim:

❖ Mixing with other states

$Sp(4)_f$

$$5 \otimes 5 = 1 \oplus 10 \oplus 14$$

$$\pi\pi \rightarrow \pi\pi \text{ (I=0,1,2)}$$

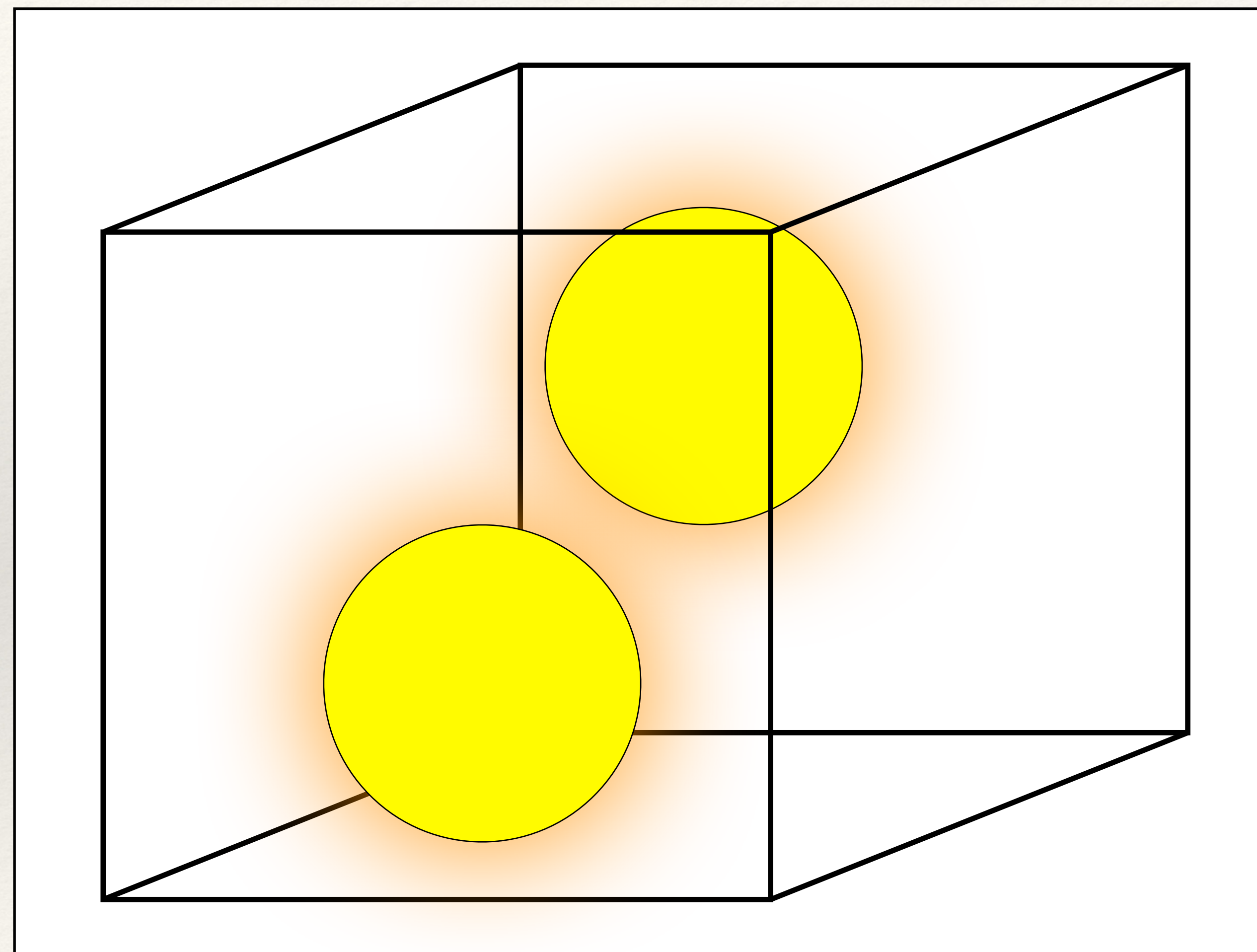
$$\pi\pi \rightarrow \rho \text{ (I=1)}$$

$$\pi\pi \rightarrow \pi\pi\pi \text{ (I=1)}$$

etc.

Scattering on the lattice

- ❖ Relate finite volume energy levels with infinite volume scattering properties
 - ❖ "Lüscher quantization condition"

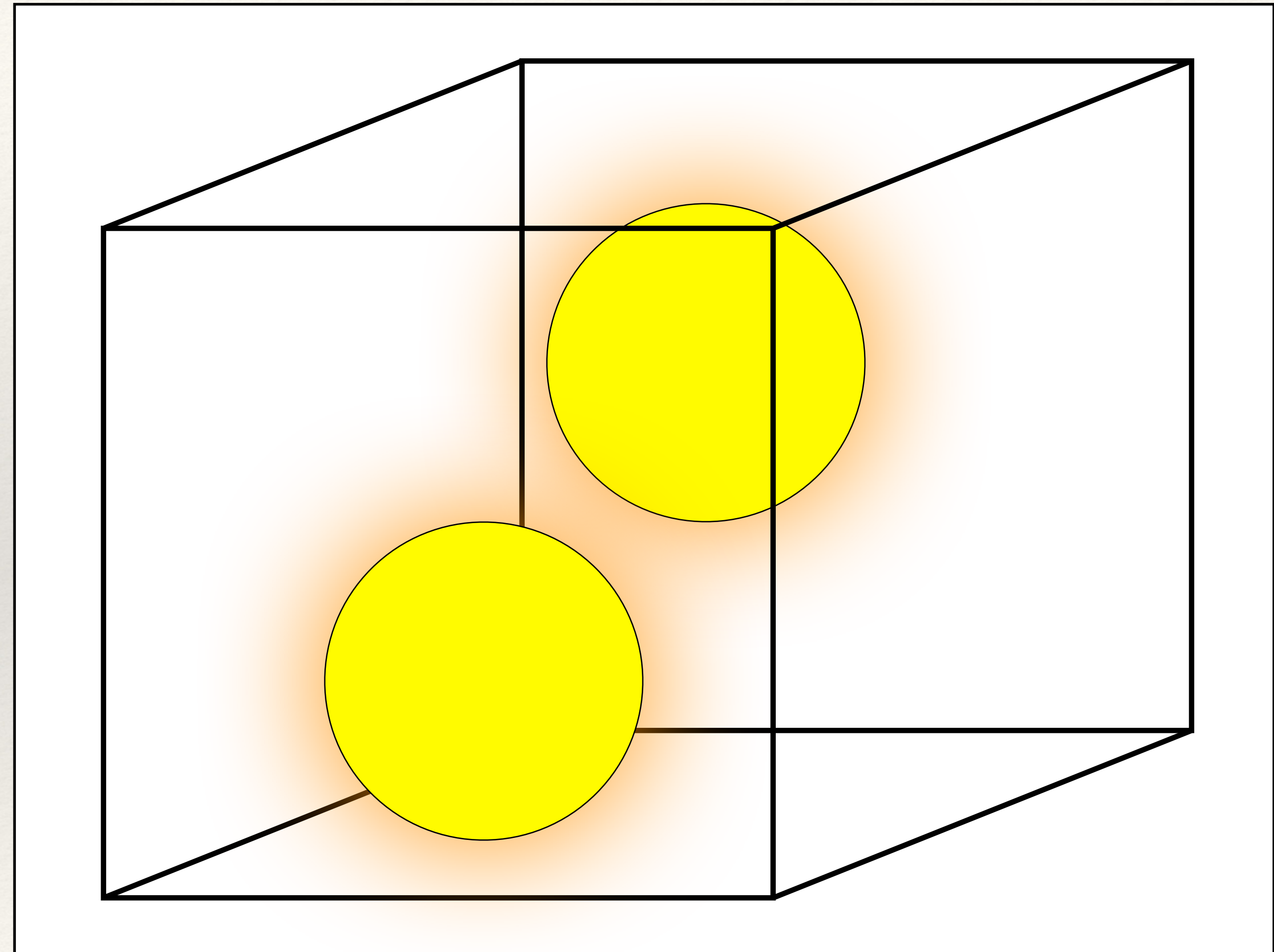


Lüscher et al.: Commun. Math. Phys. 104/105 (1986)

Briceño et al.: Rev.Mod.Phys. 90 (2018) 2, 025001

Scattering on the lattice

- ❖ Relate finite volume energy levels with infinite volume scattering properties
 - ❖ "Lüscher quantization condition"
- ❖ $\tan(\delta(\sqrt{s})) = f(E, \vec{P}, L) |_{E=E(L)}$

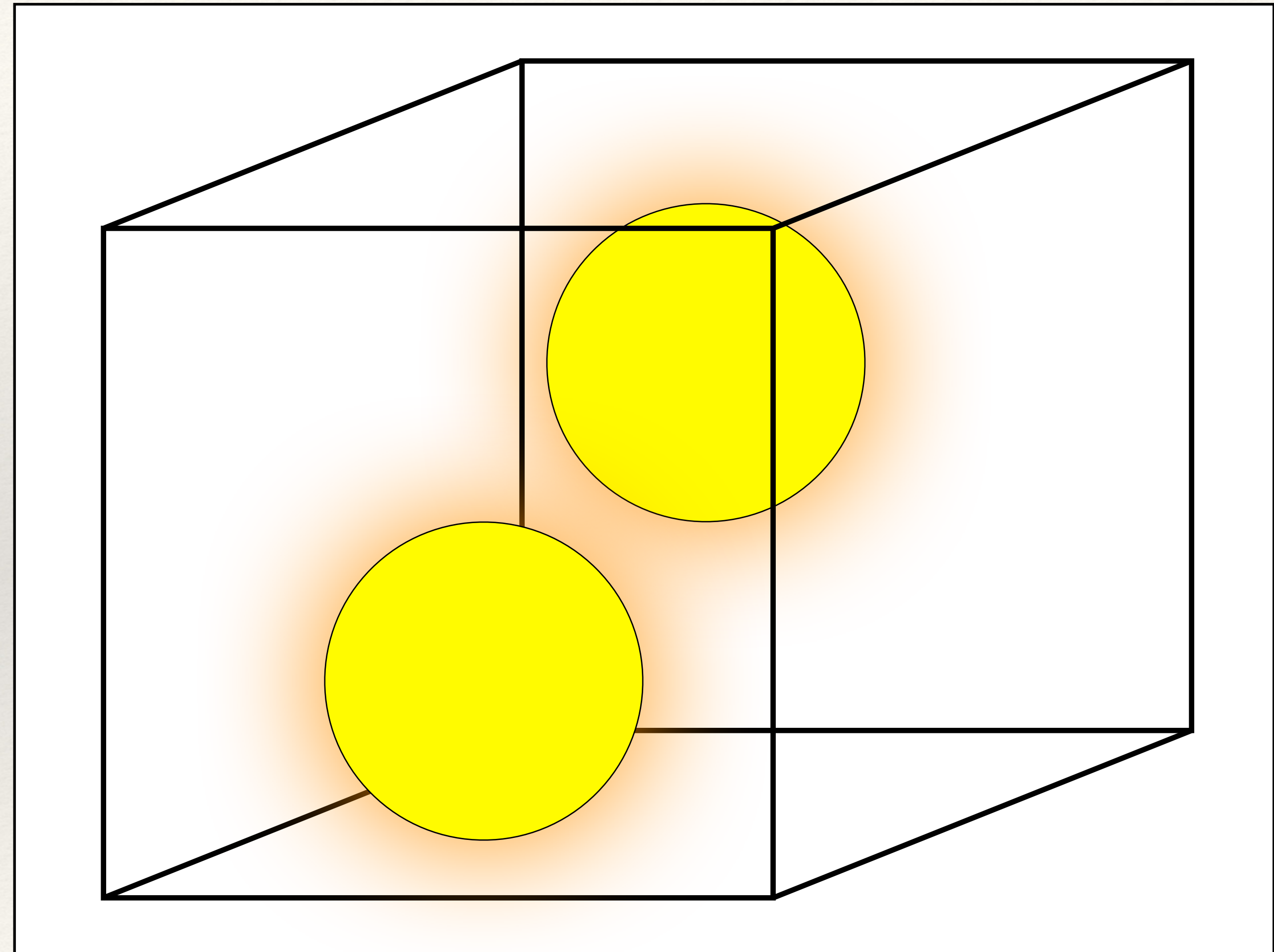


Lüscher et al.: Commun. Math. Phys. 104/105 (1986)

Briceño et al.: Rev.Mod.Phys. 90 (2018) 2, 025001

Scattering on the lattice

- ❖ Relate finite volume energy levels with infinite volume scattering properties
 - ❖ "Lüscher quantization condition"
- ❖ $\tan(\delta(\sqrt{s})) = f(E, \vec{P}, L) |_{E=E(L)}$
- ❖ Result: Energy-dependent phase-shift
 - ❖ Scattering length, cross-section, ...

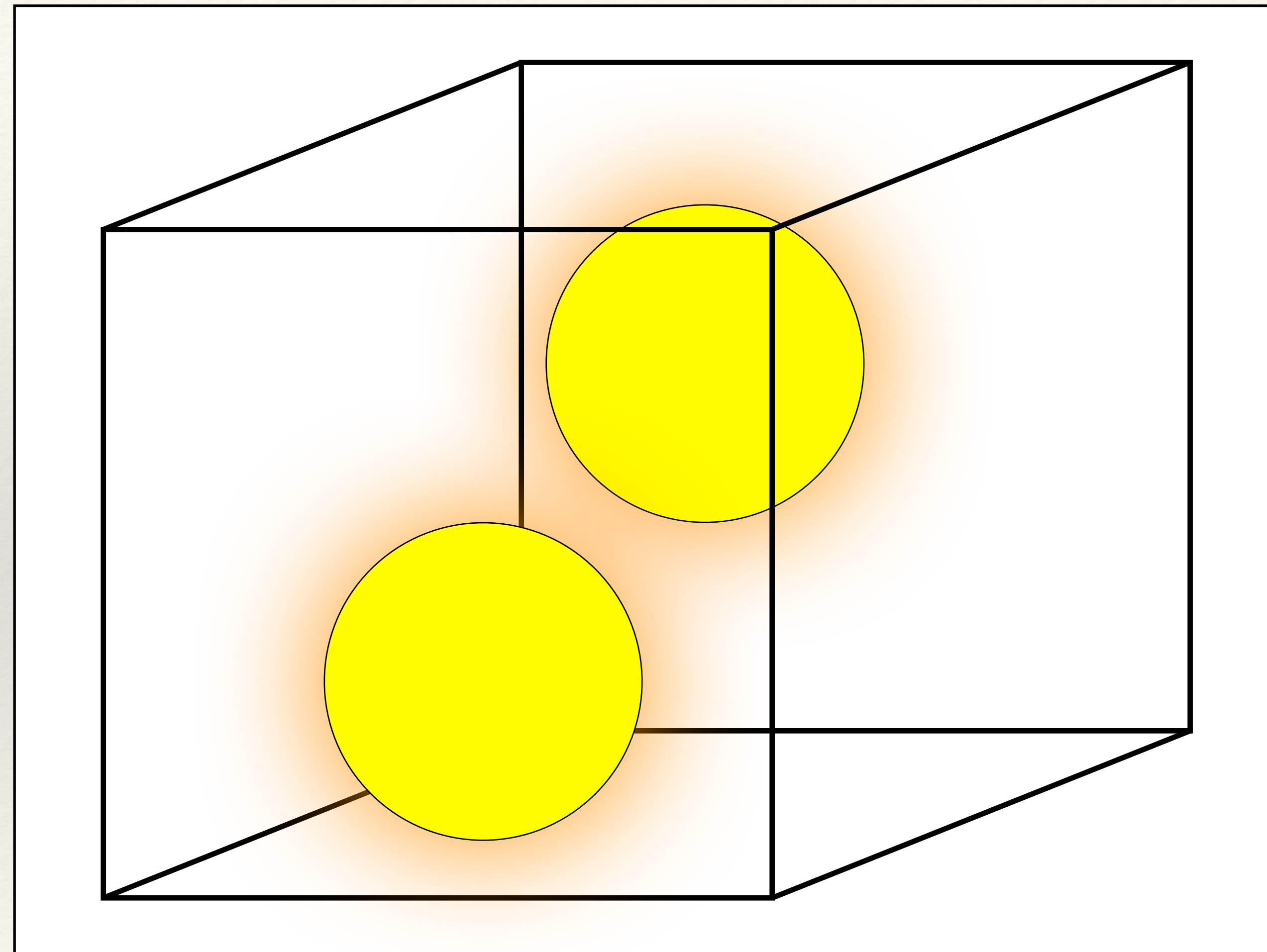


Lüscher et al.: Commun. Math. Phys. 104/105 (1986)

Briceño et al.: Rev.Mod.Phys. 90 (2018) 2, 025001

Scattering on the lattice

- ❖ 14-dim channel:
- ❖ Test ChPT prediction
- ❖ Comparison astro data

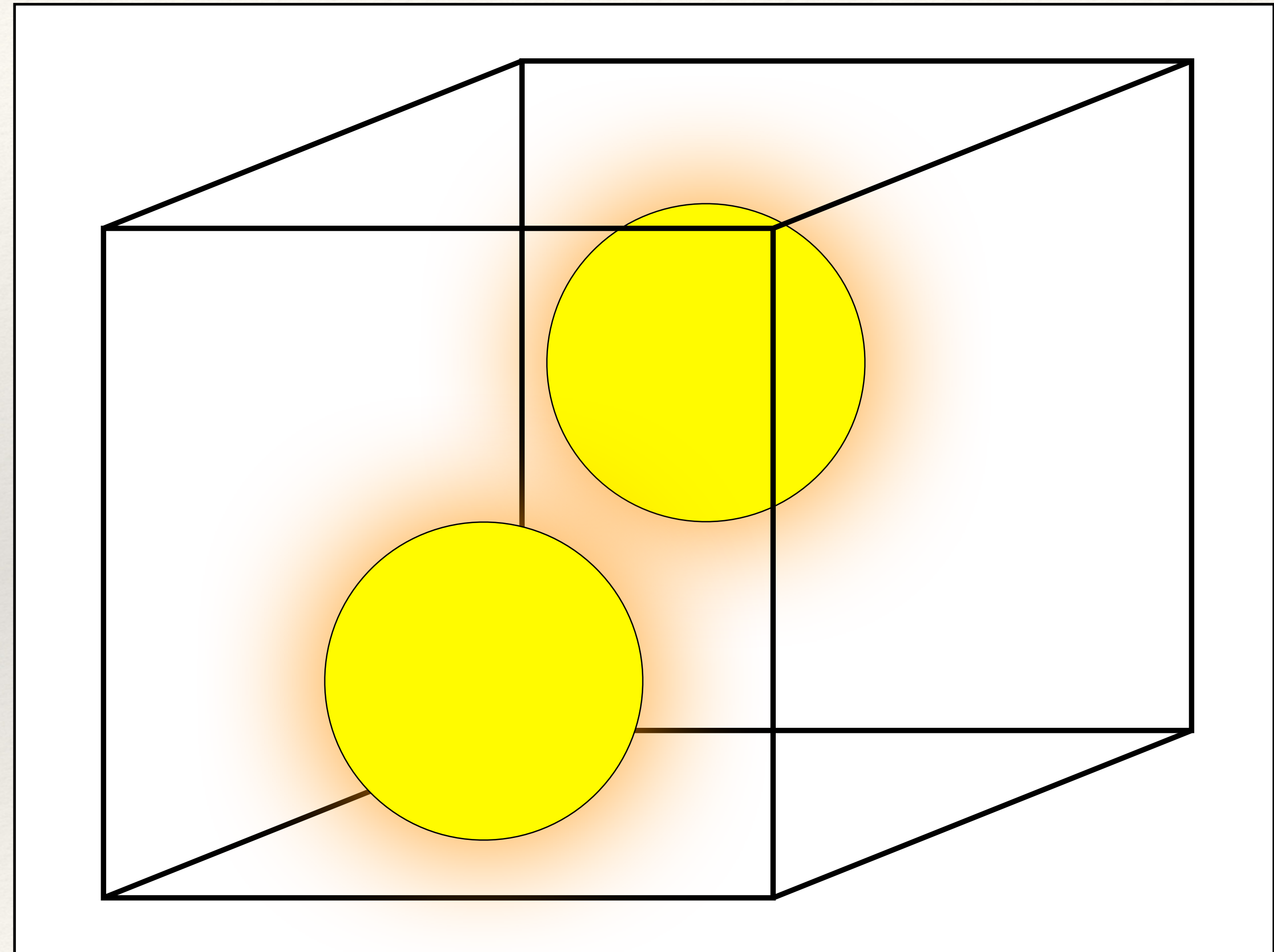


Lüscher et al.: Commun. Math. Phys. 104/105 (1986)

Briceño et al.: Rev.Mod.Phys. 90 (2018) 2, 025001

Scattering on the lattice

- ❖ 14-dim channel:
 - ❖ Test ChPT prediction
 - ❖ Comparison astro data
- ❖ 10-dim channel:
 - ❖ Work in progress
 - ❖ Preliminary comparison to 14

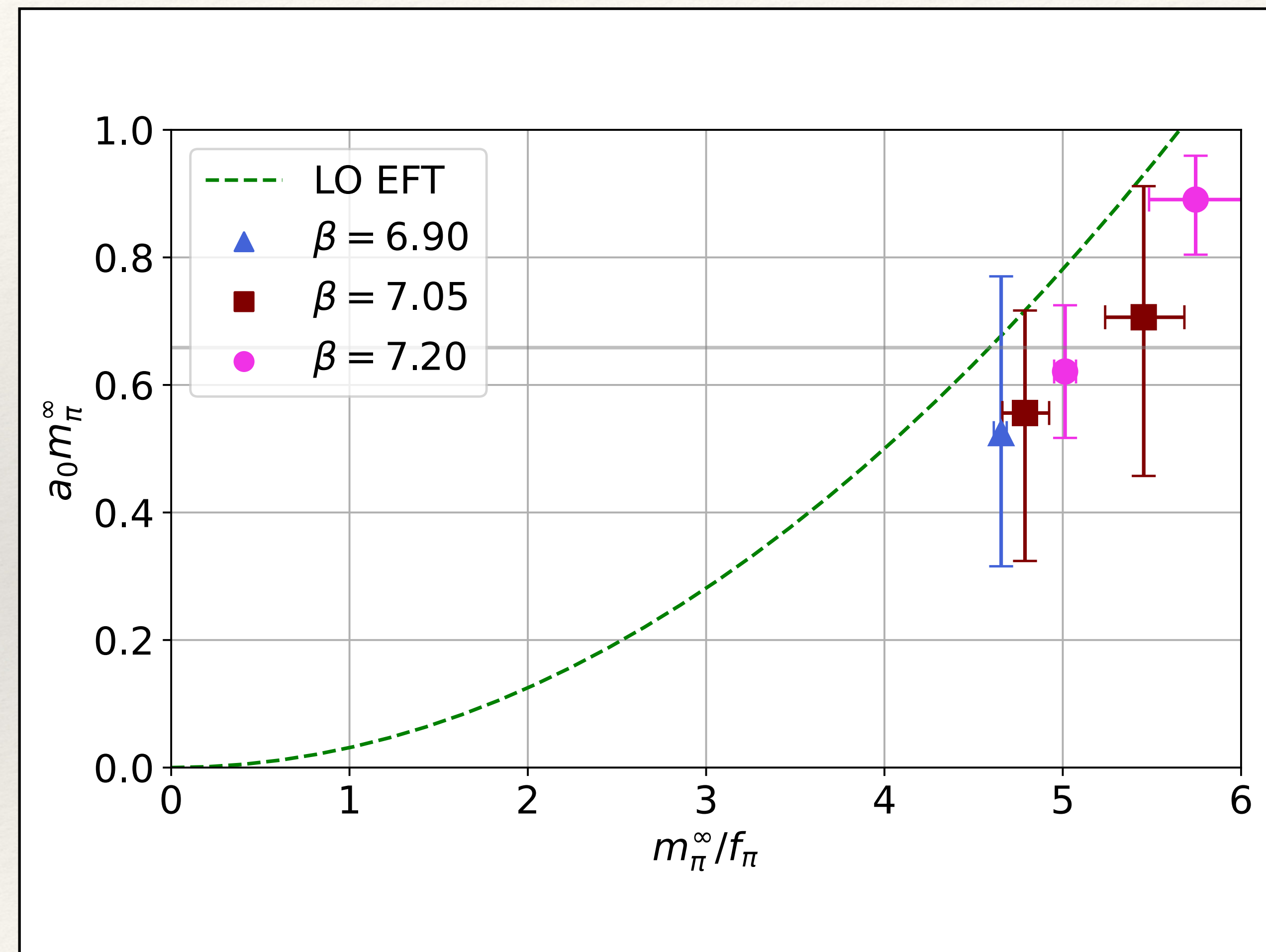


Lüscher et al.: Commun. Math. Phys. 104/105 (1986)

Briceño et al.: Rev.Mod.Phys. 90 (2018) 2, 025001

χ -pT comparison

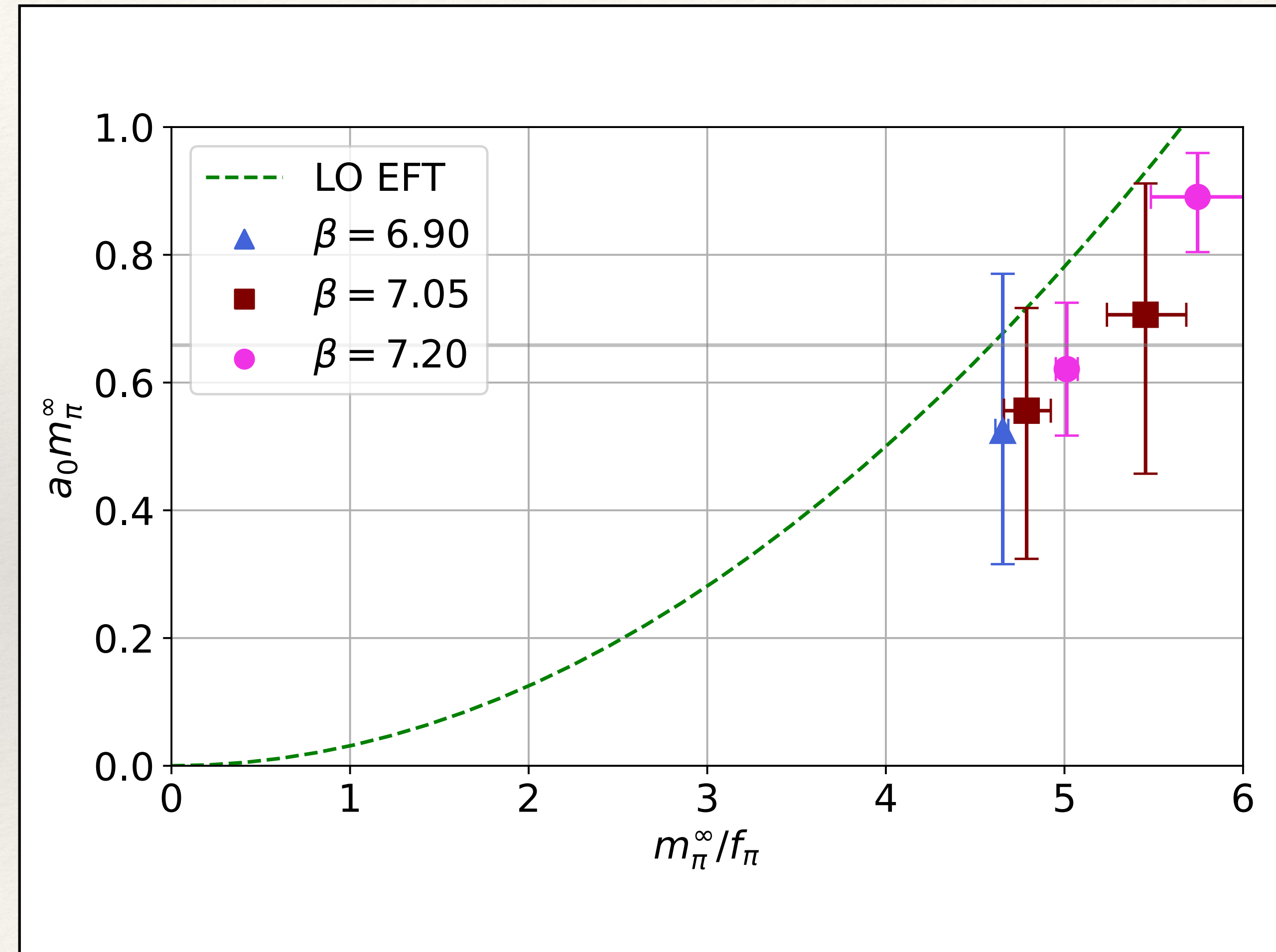
❖ Maximal channel



χ -pT comparison

❖ Maximal channel

❖ Prediction: $a_0 m_\pi = \frac{1}{32} \left(\frac{m_\pi}{f_\pi} \right)^2$



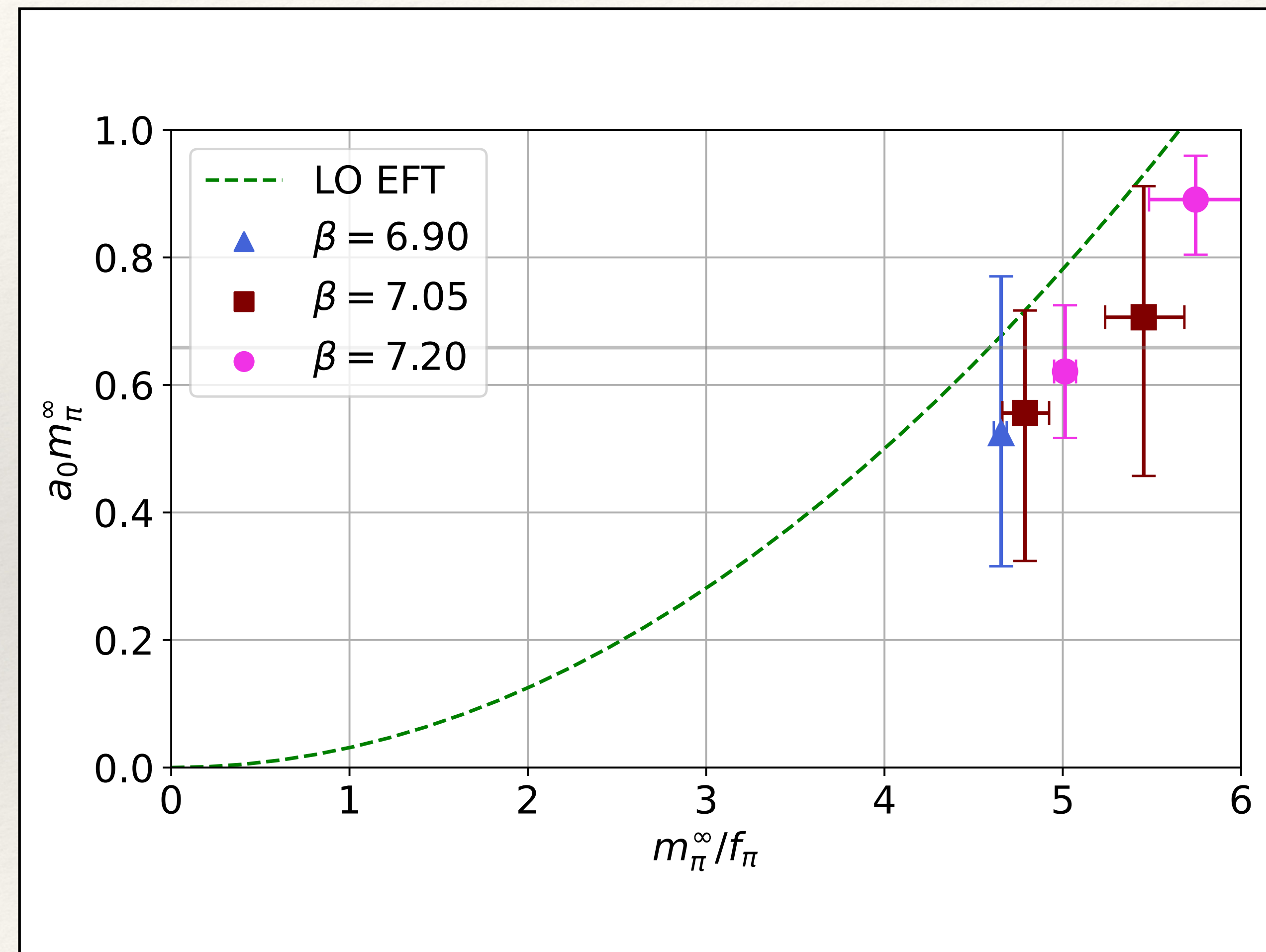
χ -pT comparison

❖ Maximal channel

❖ Prediction: $a_0 m_\pi = \frac{1}{32} \left(\frac{m_\pi}{f_\pi} \right)^2$

❖ ChPT works well

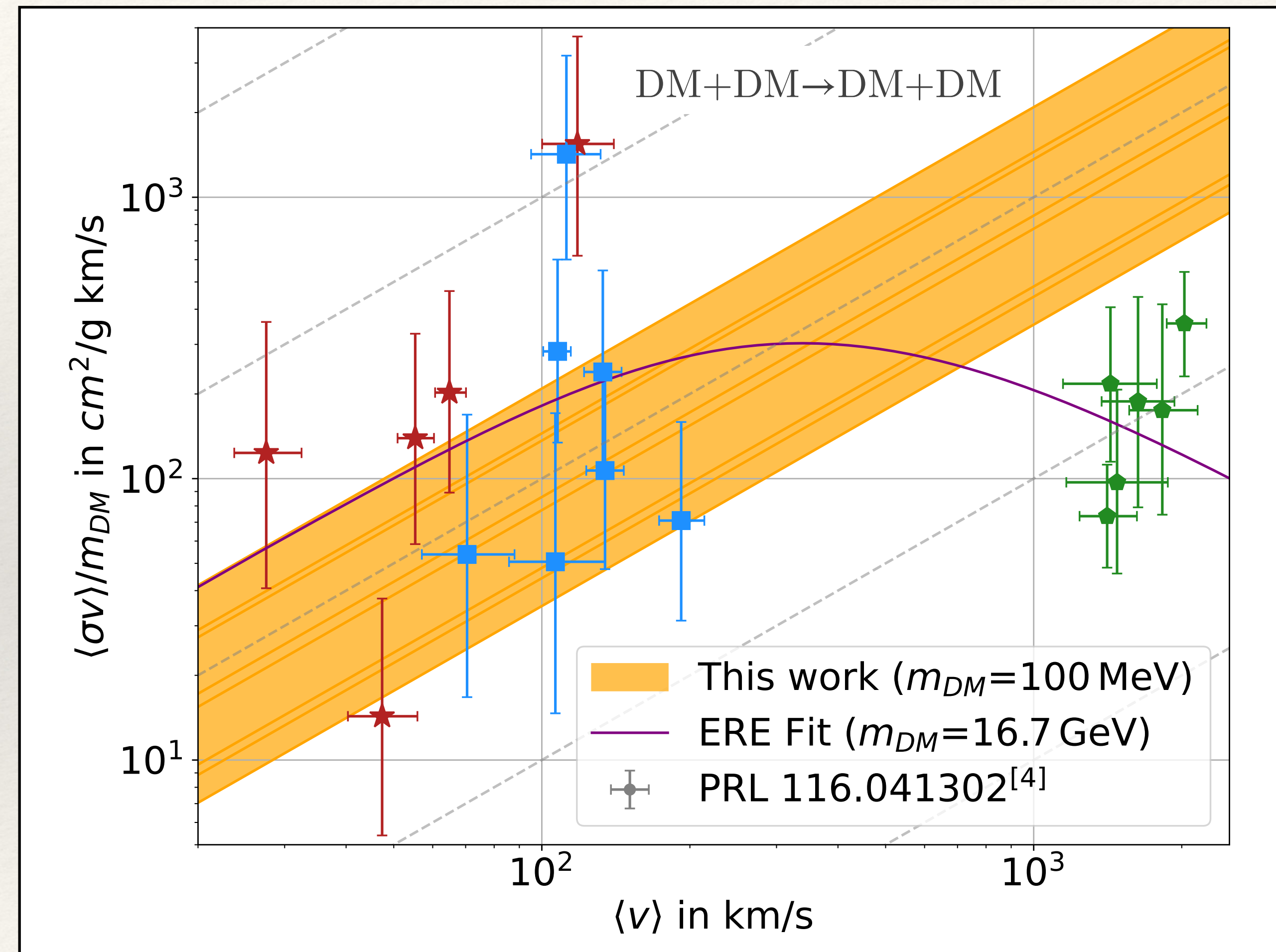
❖ NLO?



Velocity-weighted cross-section

$$\langle \sigma v \rangle = \int_0^{v_{esc}} dv \sigma(v) v f(v)$$

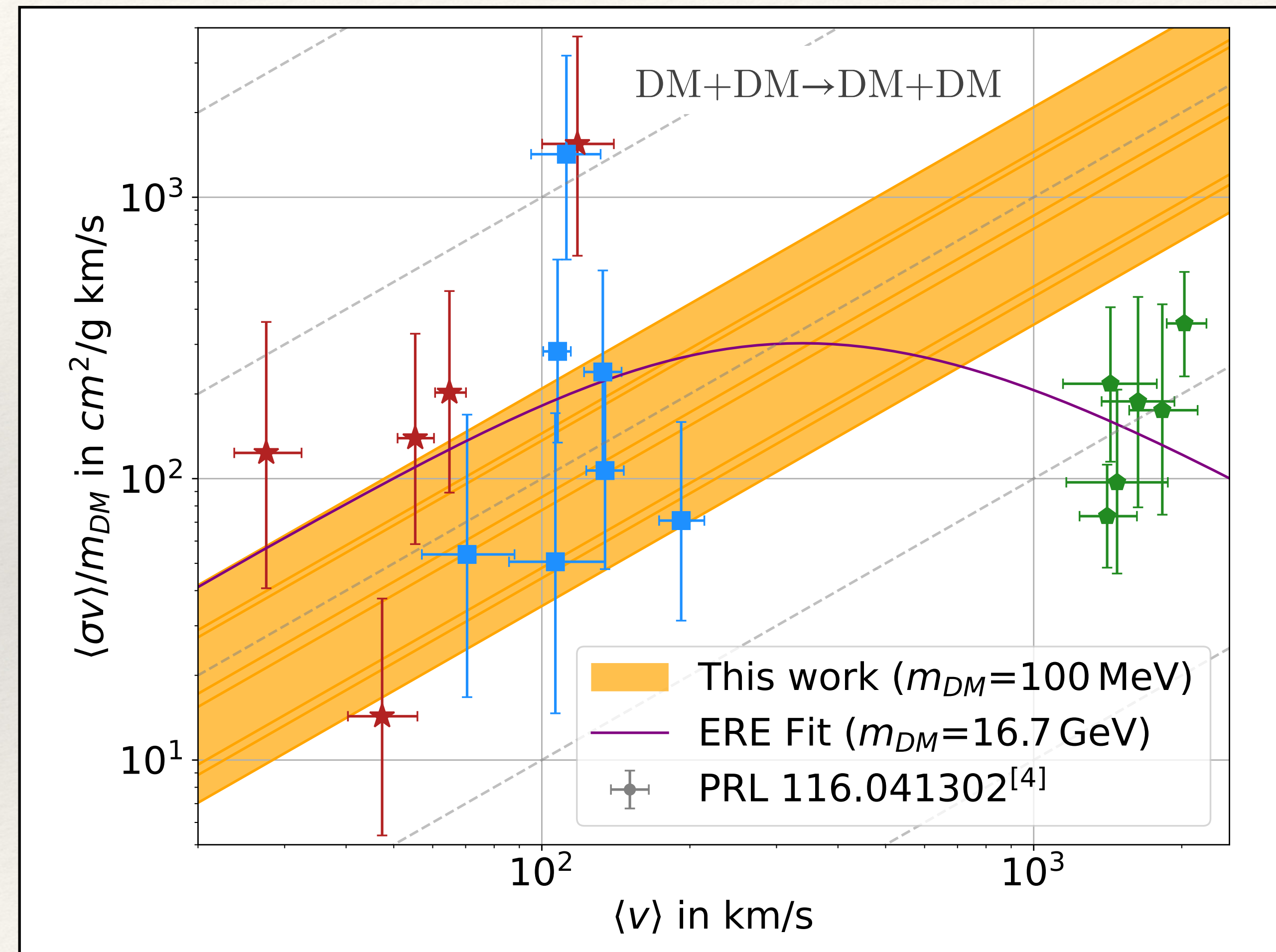
- ❖ Assumption: s-wave and maximal scattering channel



Velocity-weighted cross-section

$$\langle \sigma v \rangle = \int_0^{v_{esc}} dv \sigma(v) v f(v)$$

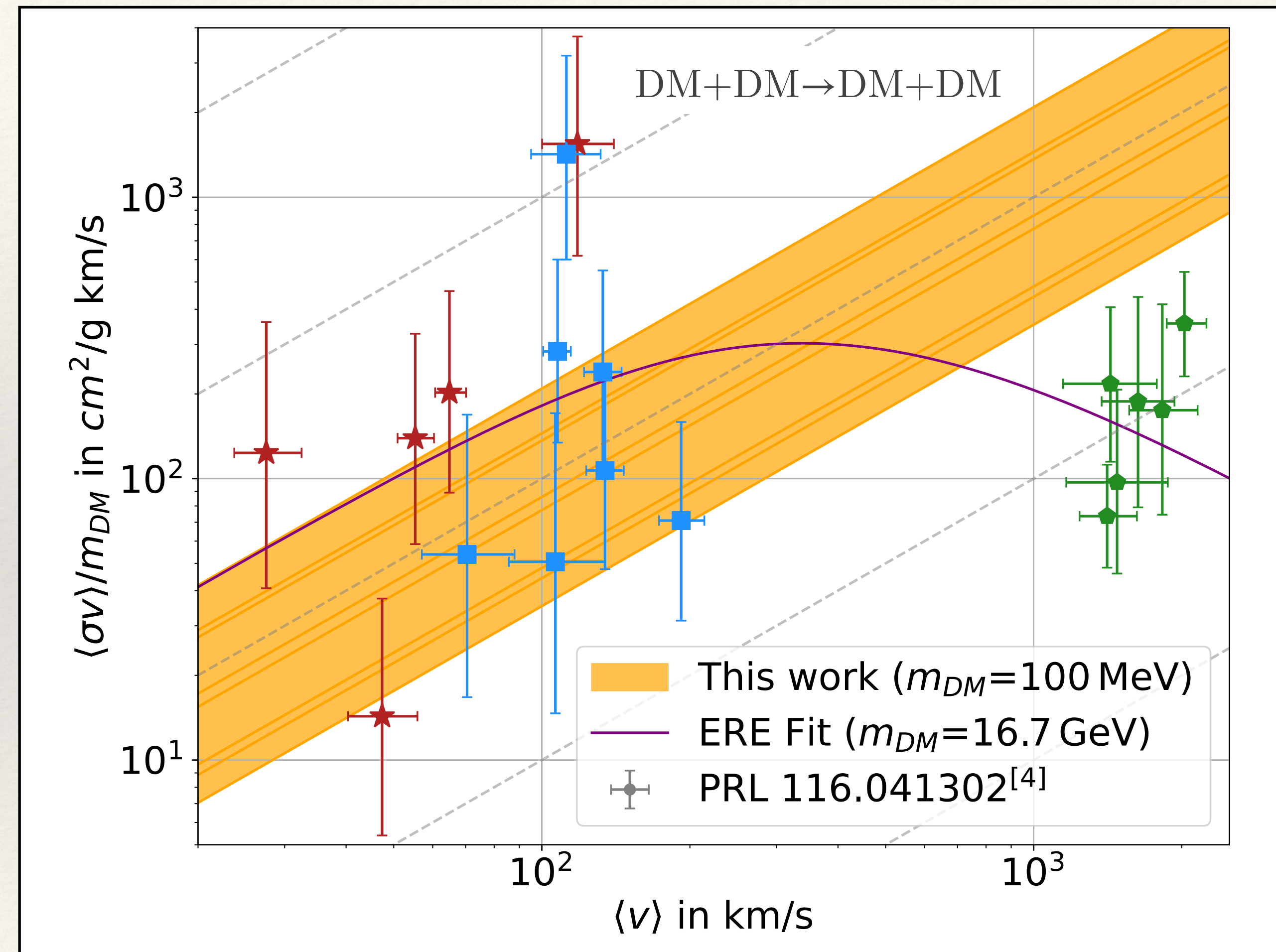
- ❖ Assumption: s-wave and maximal scattering channel
- ❖ No sign for a velocity dependence
- ❖ Discrepancy in $a_0 m_{DM}$

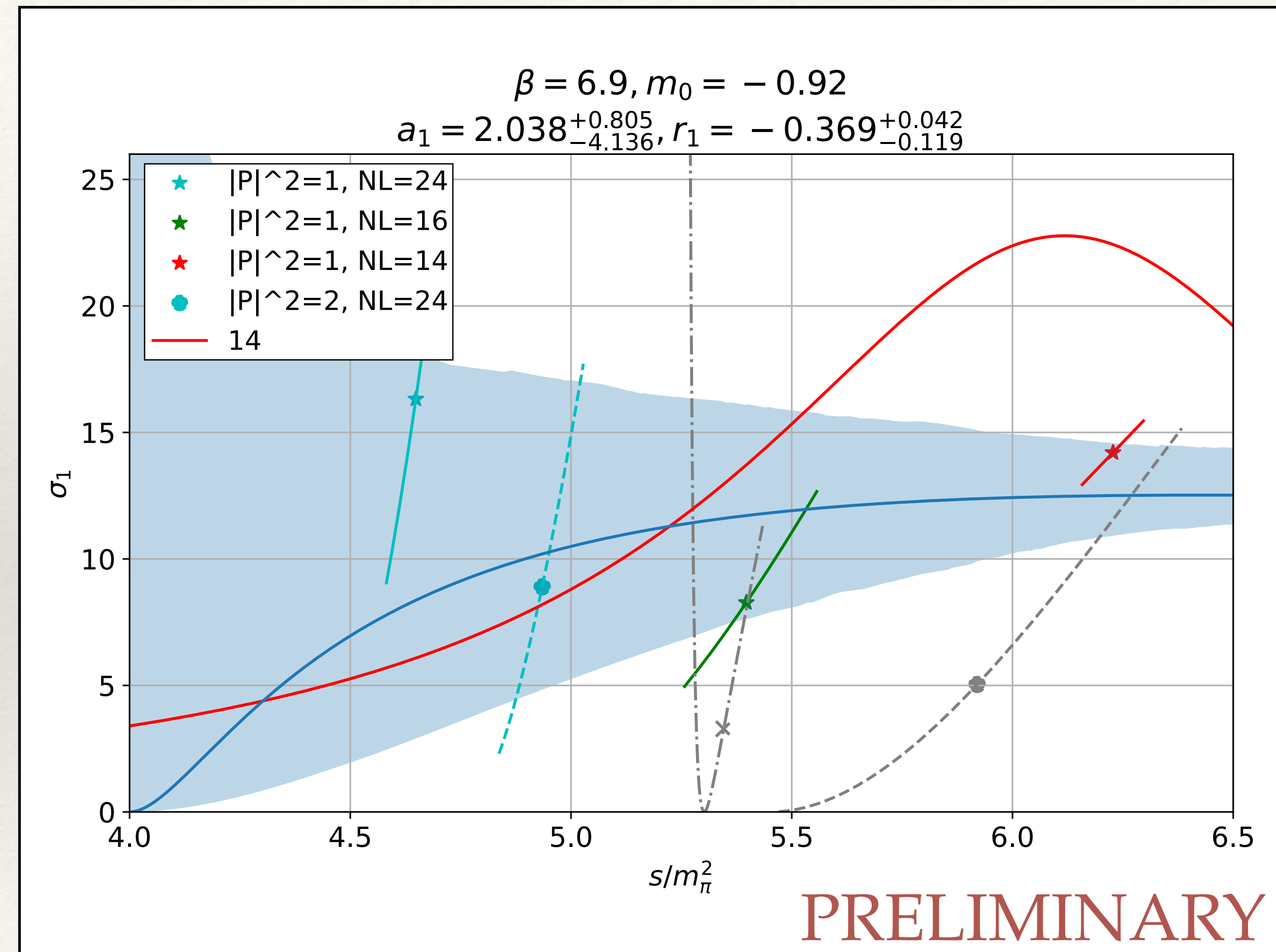


Velocity-weighted cross-section

$$\langle \sigma v \rangle = \int_0^{v_{esc}} dv \sigma(v) v f(v)$$

- ❖ Assumption: s-wave and maximal scattering channel
- ❖ No sign for a velocity dependence
- ❖ Discrepancy in $a_0 m_{DM}$
- ❖ $m_{DM} \sim 100$ MeV predicted by SIMP



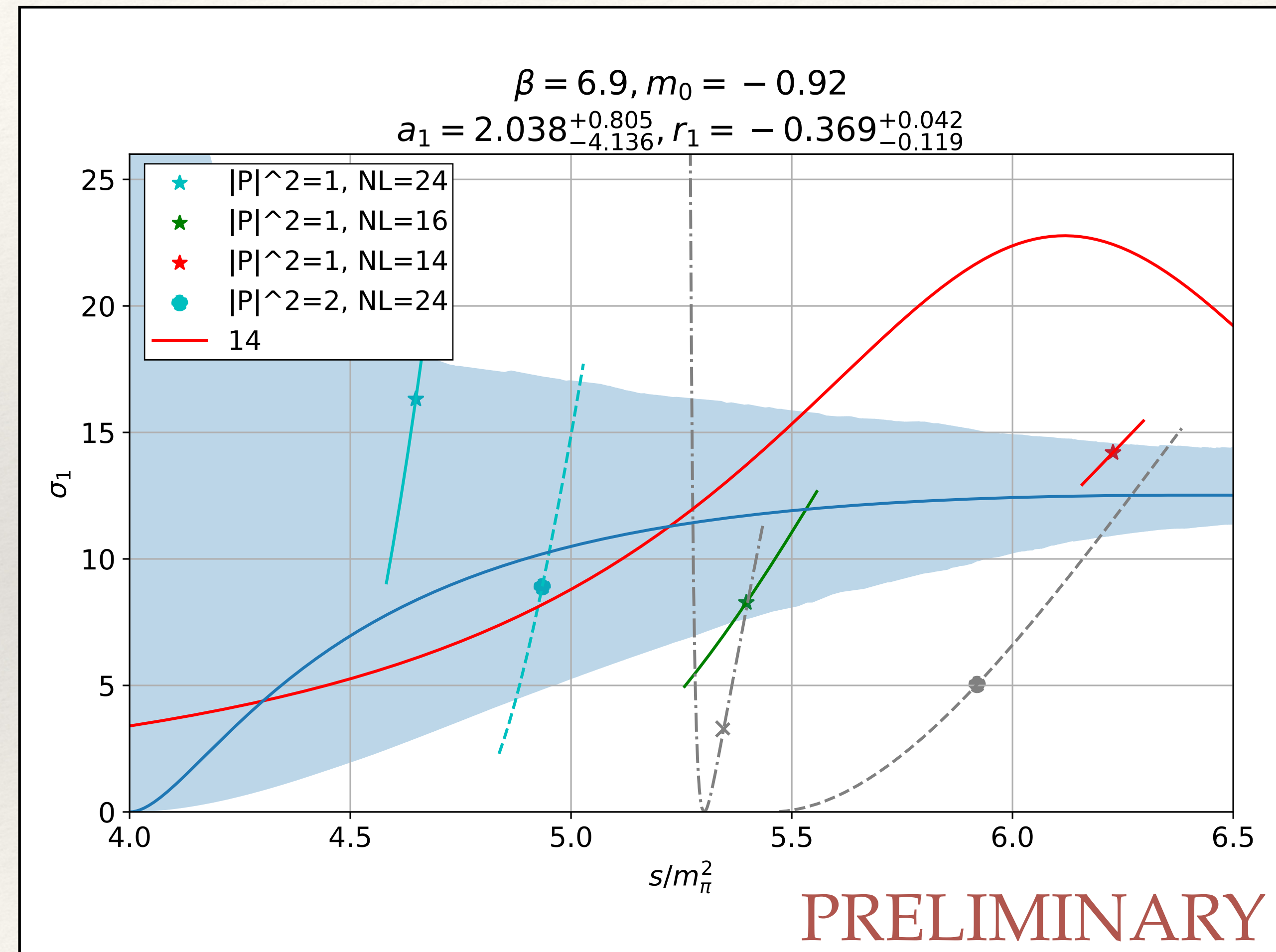
$$\diamond \frac{m_\rho}{m_\pi} = \{1.6, 1.9, 2.3\}$$


Work in Progress

- ❖ 10-dim scattering channel: 3 ensembles:

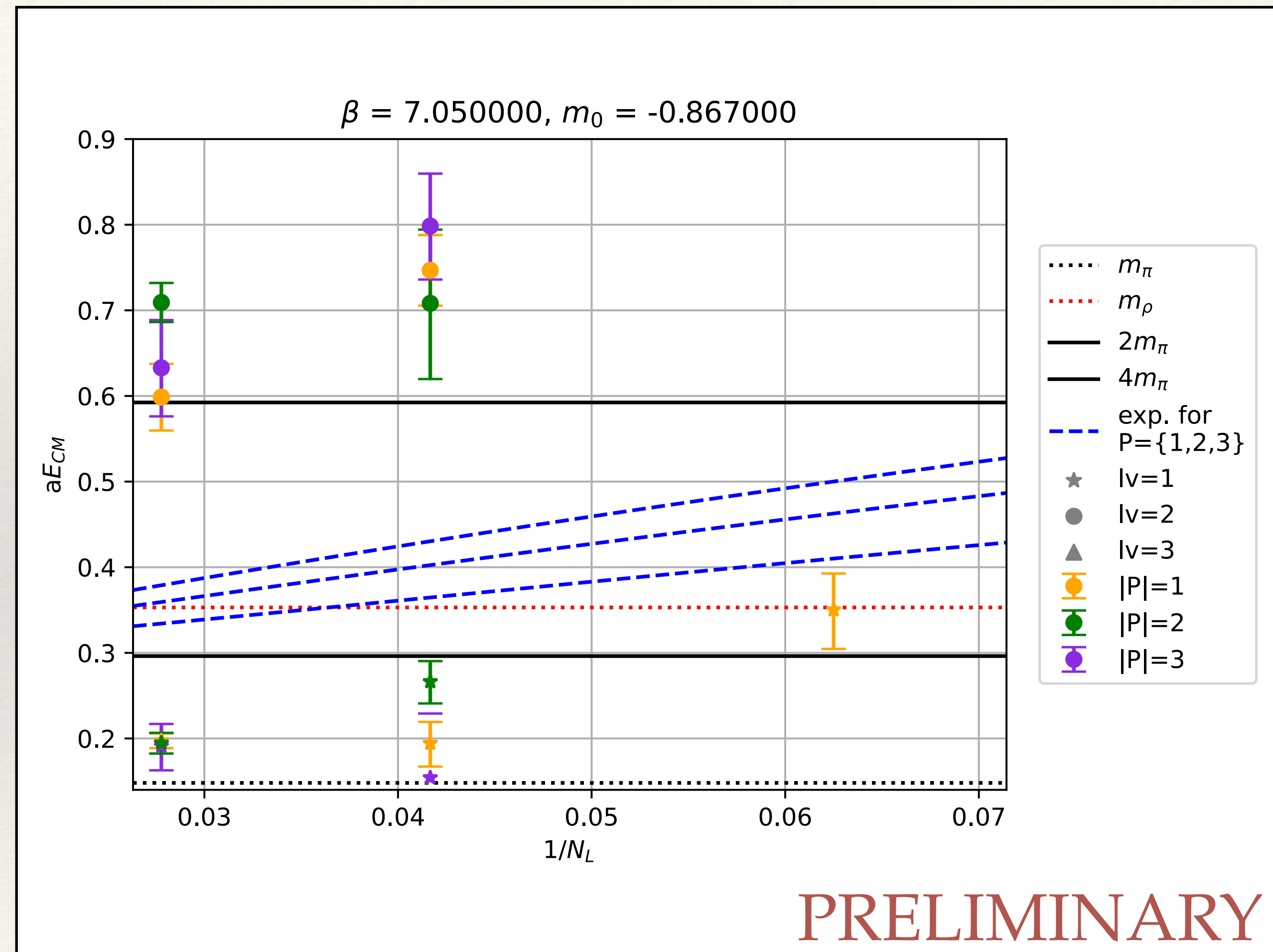
- ❖ $\frac{m_\rho}{m_\pi} = \{1.6, 1.9, 2.3\}$

- ❖ Non-resonant: Cross-sections comparable



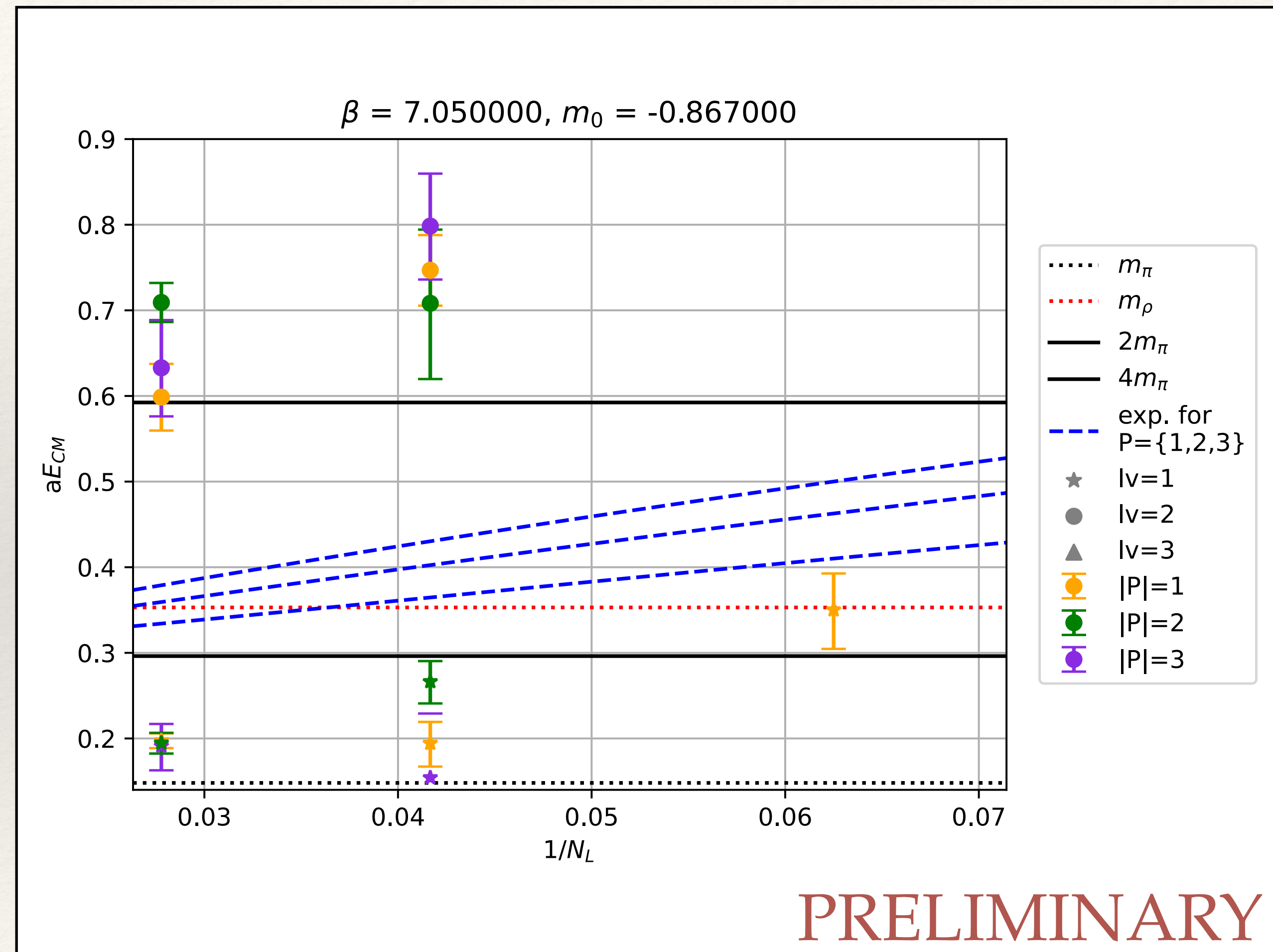
Work in Progress

- ❖ 10-dim scattering channel: 3 ensembles:
 - ❖ $\frac{m_\rho}{m_\pi} = \{1.6, 1.9, 2.3\}$
- ❖ Non-resonant: Cross-sections comparable
- ❖ What happens if the vector meson becomes resonant (like in QCD)?
 - ❖ Work in progress



Work in Progress

- ❖ 10-dim scattering channel: 3 ensembles:
 - ❖ $\frac{m_\rho}{m_\pi} = \{1.6, 1.9, 2.3\}$
- ❖ Non-resonant: Cross-sections comparable
- ❖ What happens if the vector meson becomes resonant (like in QCD)?
 - ❖ Work in progress
- ❖ Energy levels can be reused for $3\pi \rightarrow 2\pi$



Summary & Outlook

- ❖ Non-perturbative results for $\pi\pi$ -scattering
- ❖ First principle verification for low energy constants
- ❖ With $\pi\pi \rightarrow \rho$ & $\pi\pi\pi \rightarrow \pi\pi$ we will obtain a good understanding of the model

Summary & Outlook

- ❖ Non-perturbative results for $\pi\pi$ -scattering
- ❖ First principle verification for low energy constants
- ❖ With $\pi\pi \rightarrow \rho$ & $\pi\pi\pi \rightarrow \pi\pi$ we will obtain a good understanding of the model

Thank you!