

Exploring the role of initial α -clustered nuclear structure on final state flow coefficients via pC, pO and OO collisions at the LHC

Based on:

A. Menon Kavumpadikkal Radhakrishnan, S. Prasad, N. Mallick and R. Sahoo, Eur. Phys. J. A **61**, 134 (2025).

A. Menon Kavumpadikkal Radhakrishnan, S. Prasad, N. Mallick, R. Sahoo and G. G. Barnaföldi, [arXiv:2505.22367 [hep-ph]].



HUN
REN



Aswathy Menon K. R.

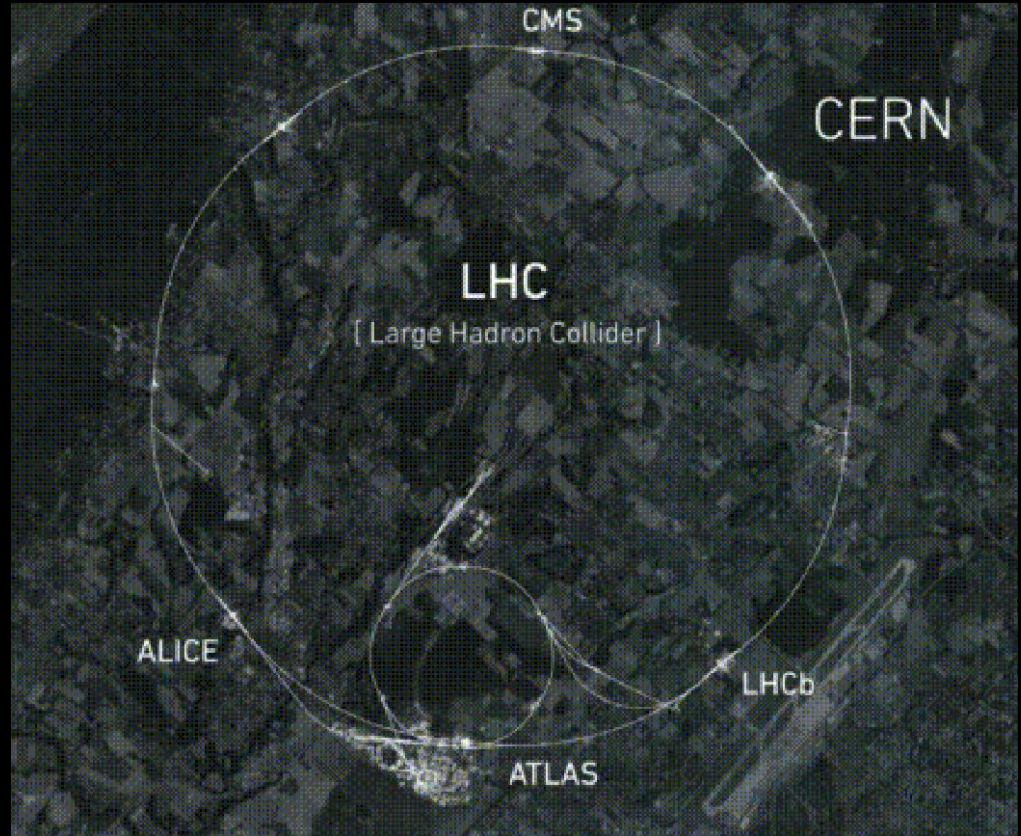
Department of Physics

Indian Institute of Technology Indore

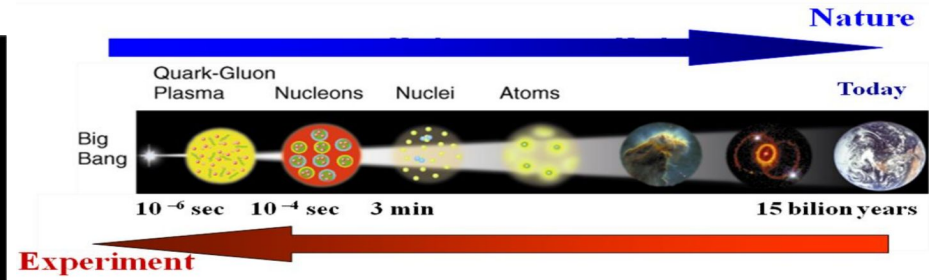
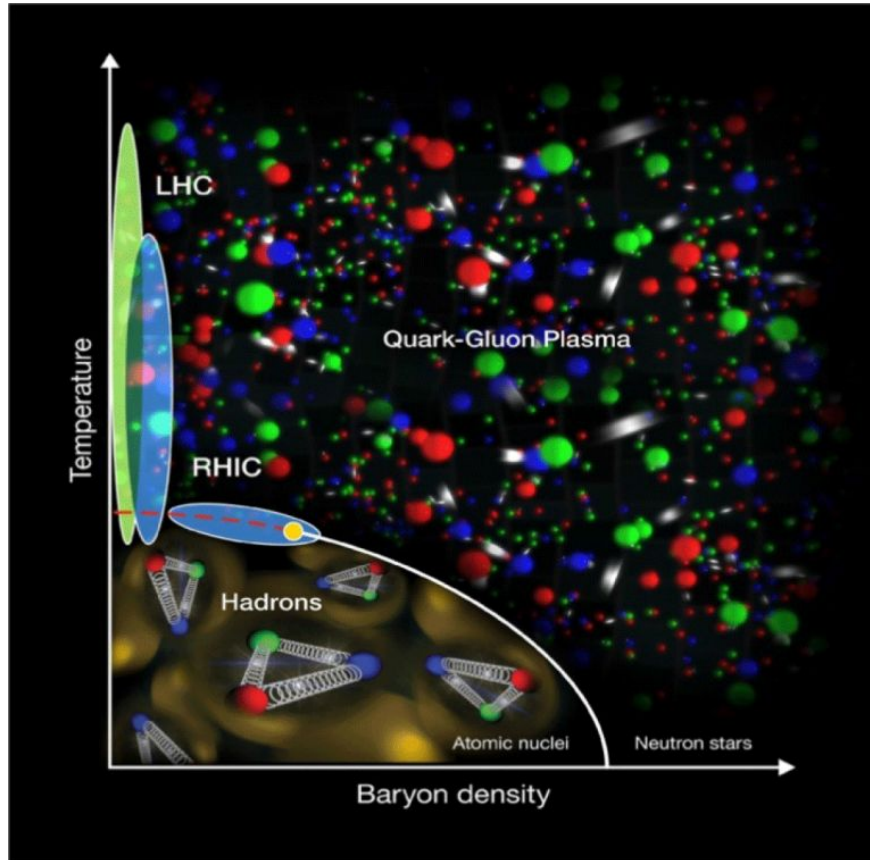
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Outline

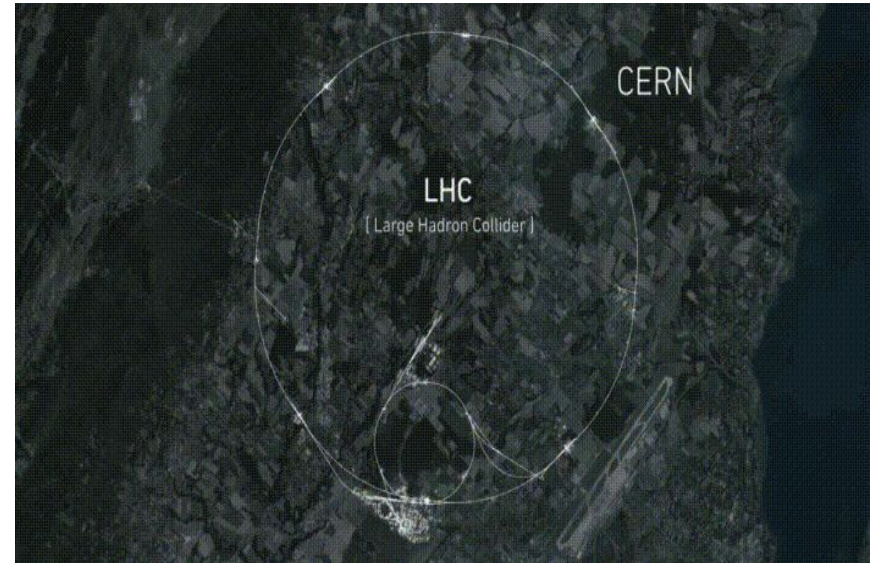
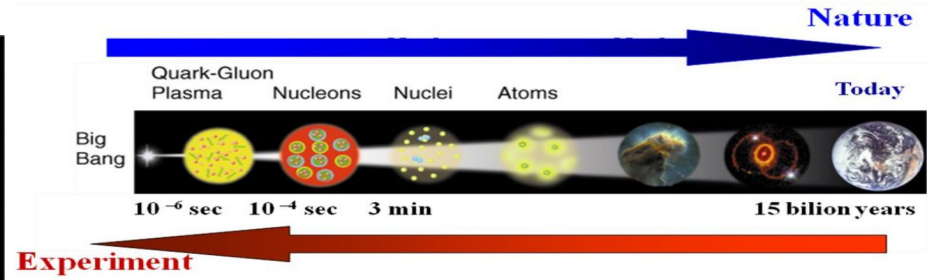
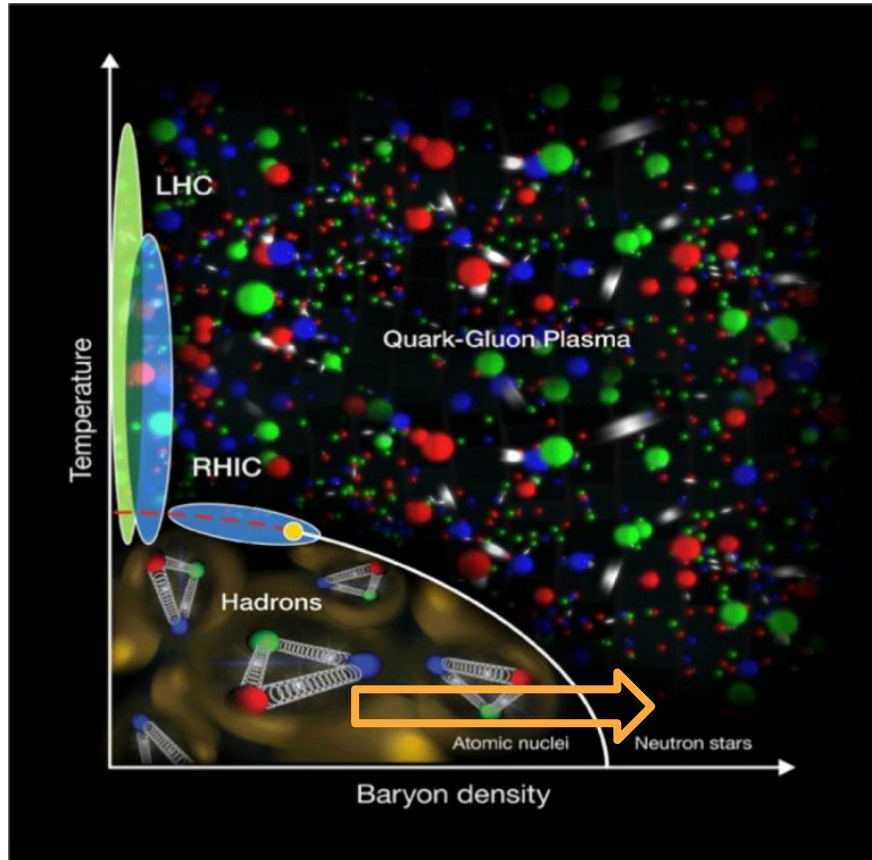
- ★ Relativistic heavy-ion collisions
- ★ Light-ion collisions
- ★ α -clustered nuclear structure
- ★ Anisotropic flow estimation
- ★ Results



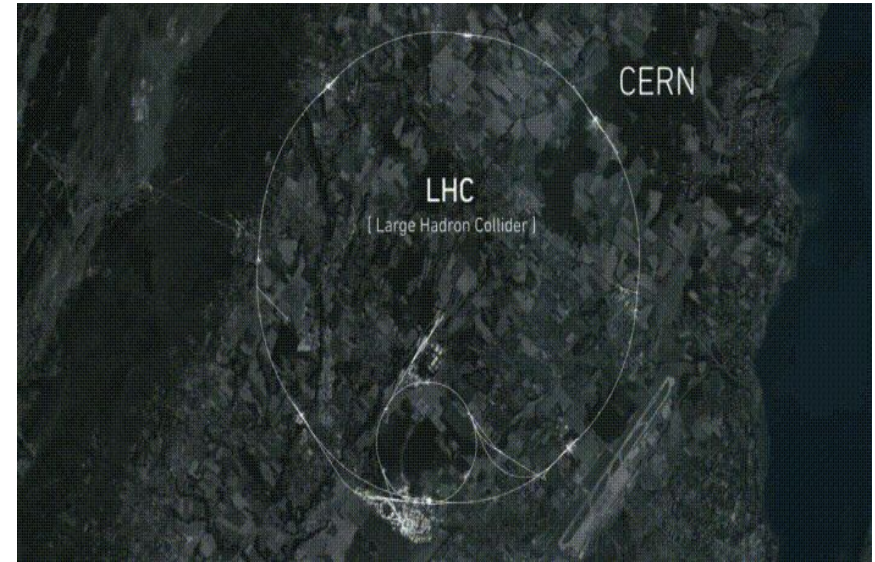
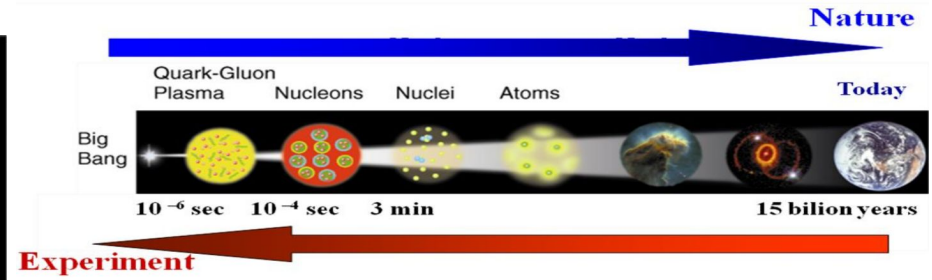
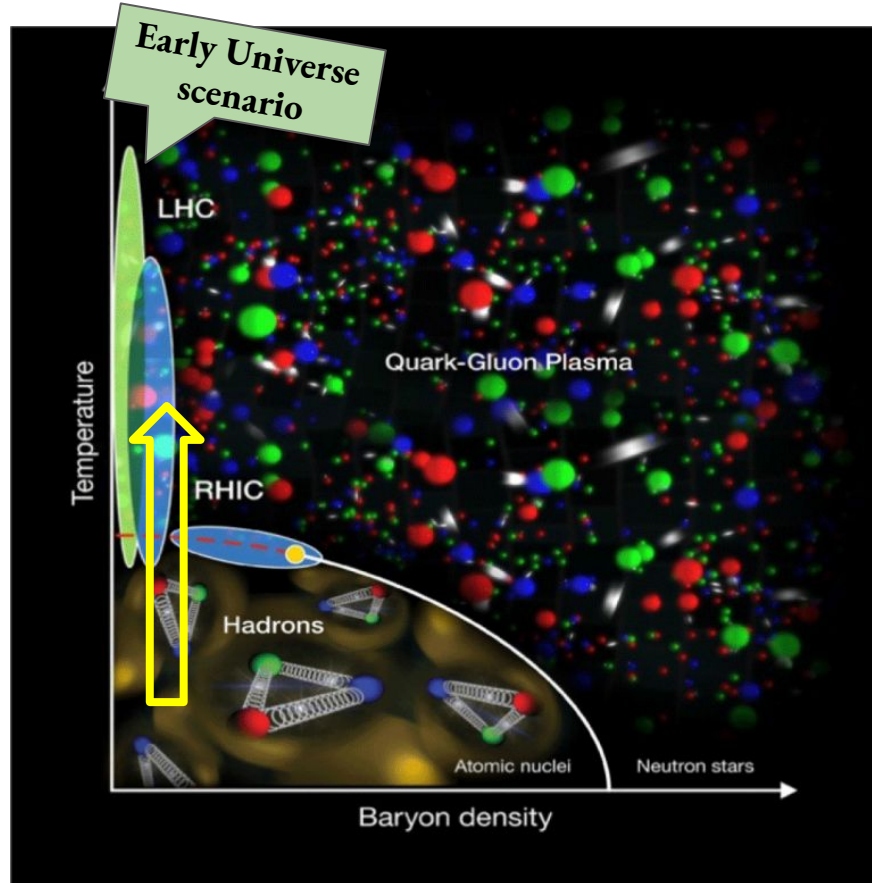
Why Relativistic collisions ?



Why Relativistic collisions ?



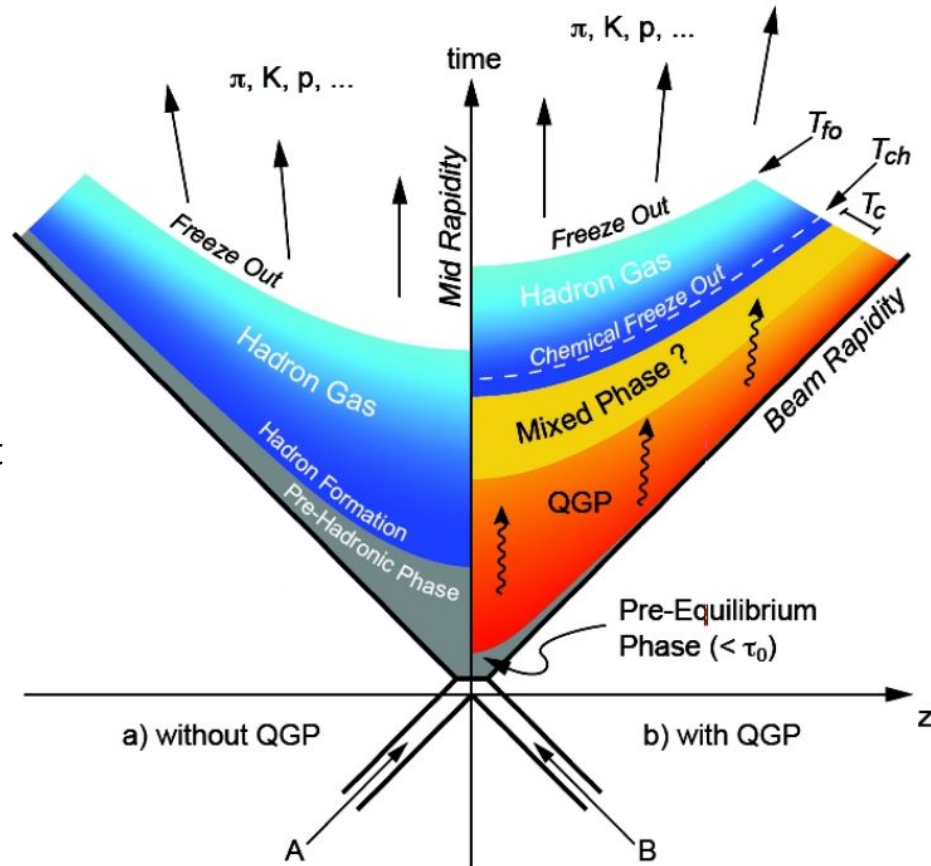
Why Relativistic collisions ?



Space-time evolution of Relativistic collisions

Without QGP :

- Pre-hadronic phase
- Hadron Gas Phase
- Hadronic Freezeout



With QGP :

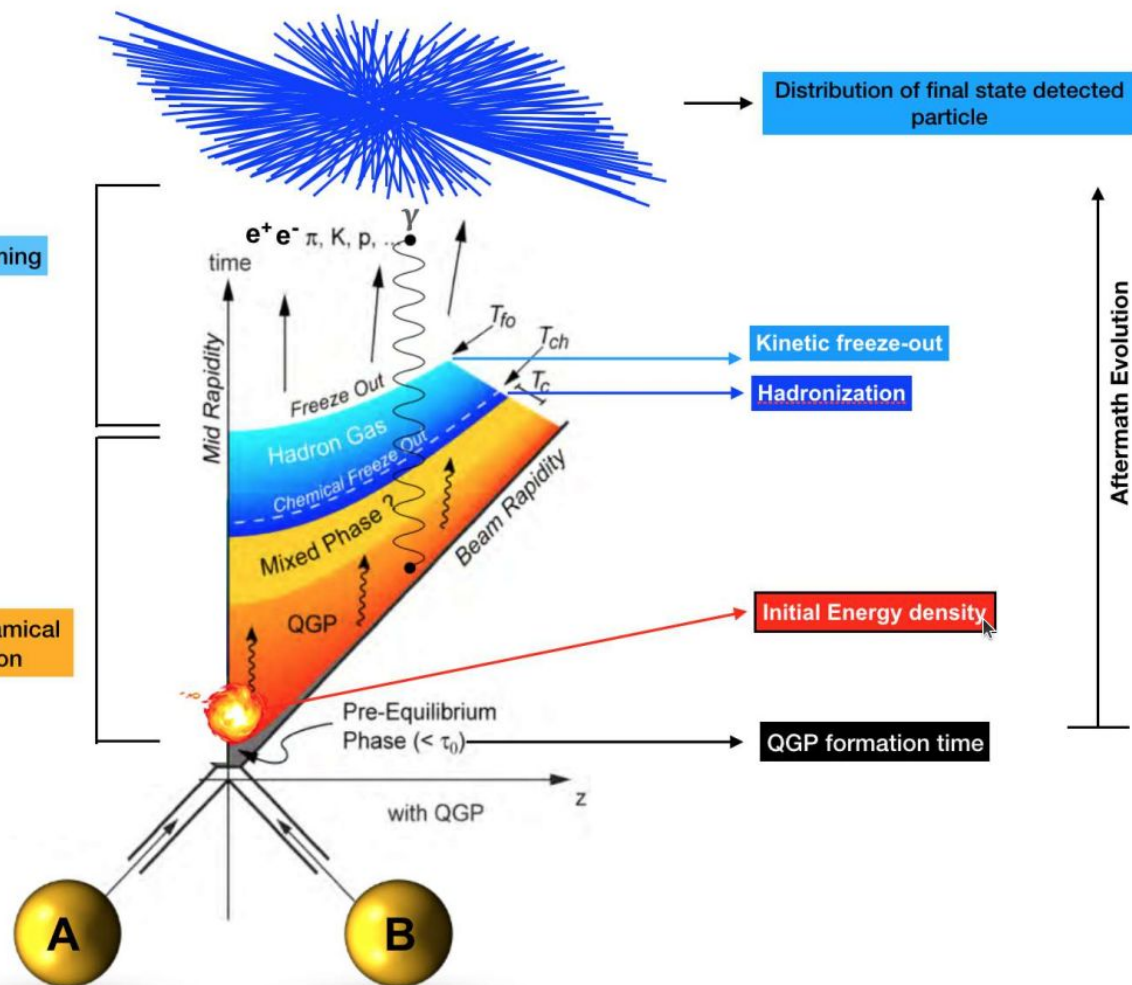
- Pre-Equilibrium phase
- QGP
- Mixed Phase
- Chemical Freezeout
- Hadronic Phase
- Kinetic Freezeout

With QGP :

- Pre-Equilibrium phase
- Quark Gluon Plasma (QGP)
- Mixed Phase
- Chemical Freezeout
- Hadronic Phase
- Kinetic Freezeout

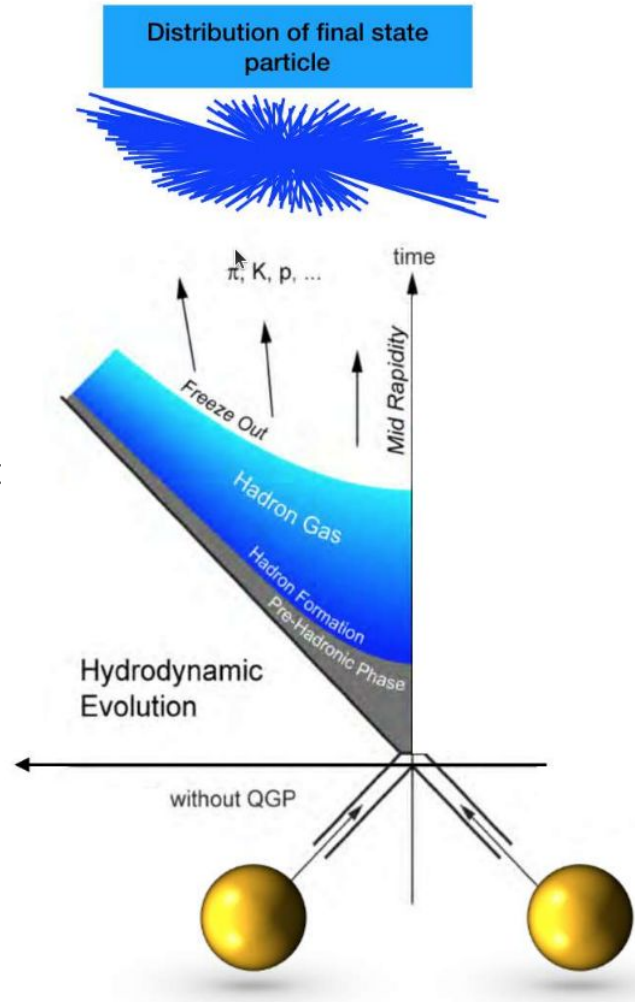
Free streaming

Hydrodynamical Evolution



Without QGP :

- Pre-hadronic phase
- Hadron Gas Phase
- Hadronic Freezeout

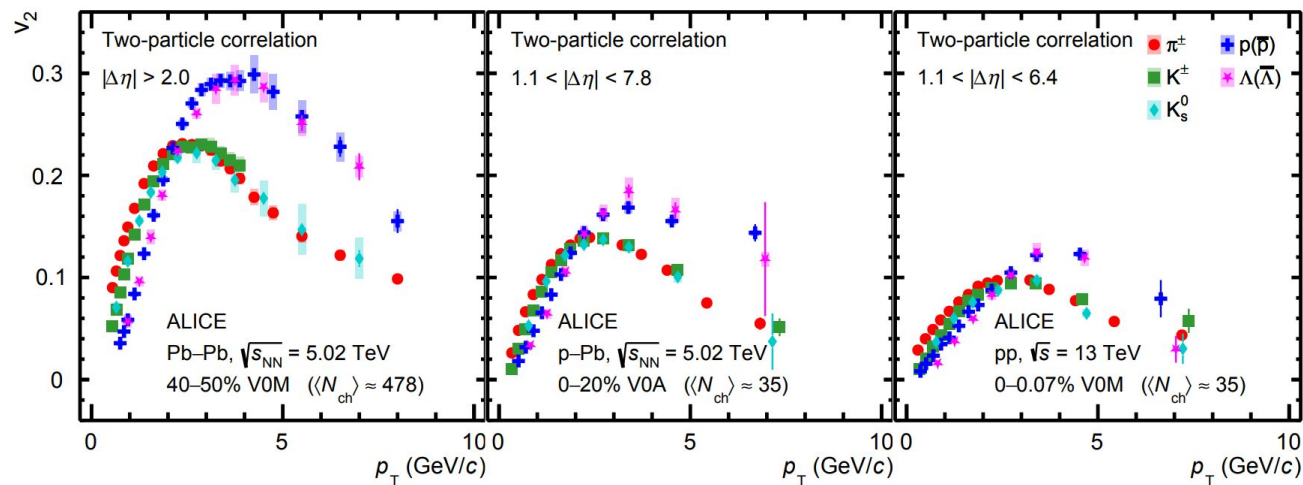


This picture of space-time evolution is expected to occur in **hadronic collisions**, where formation of QGP is least anticipated



Motivation

- Through relativistic heavy-ion collisions, we search for indirect **signatures of QGP**
- Small collisions like **pp collisions provide baseline measurements** as medium formation is not expected here
- But presence of heavy-ion like signatures are now observed in small collision systems too!
 - **High-multiplicity pp and p-Pb collisions** show signatures of collective flow and strangeness enhancement

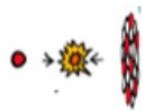


S. Acharya et al. [ALICE], [arXiv:2411.09323]

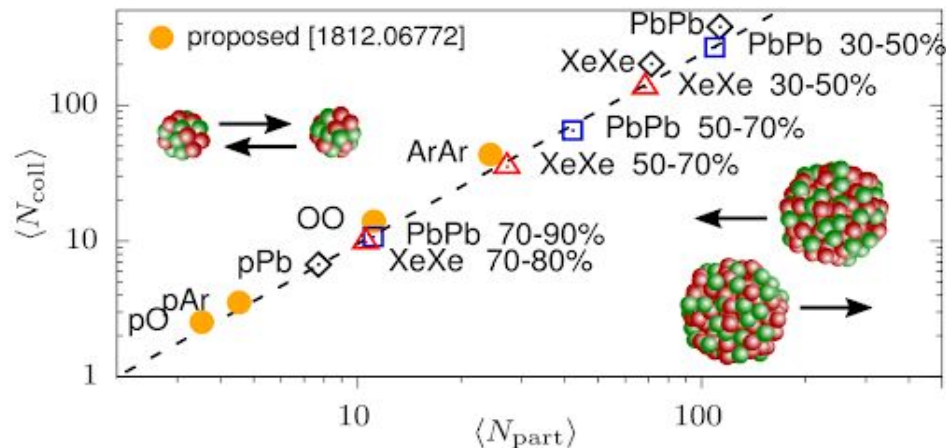
**Focus :
SMALL
COLLISION
SYSTEMS**



**p-O and O-O
collisions
took place in
the Run 3 at
the LHC in
2025**



pO and OO collisions at the LHC



1

PERFECT SYSTEM SIZE to fill multiplicity gap between pp, p-Pb and Pb-Pb

3

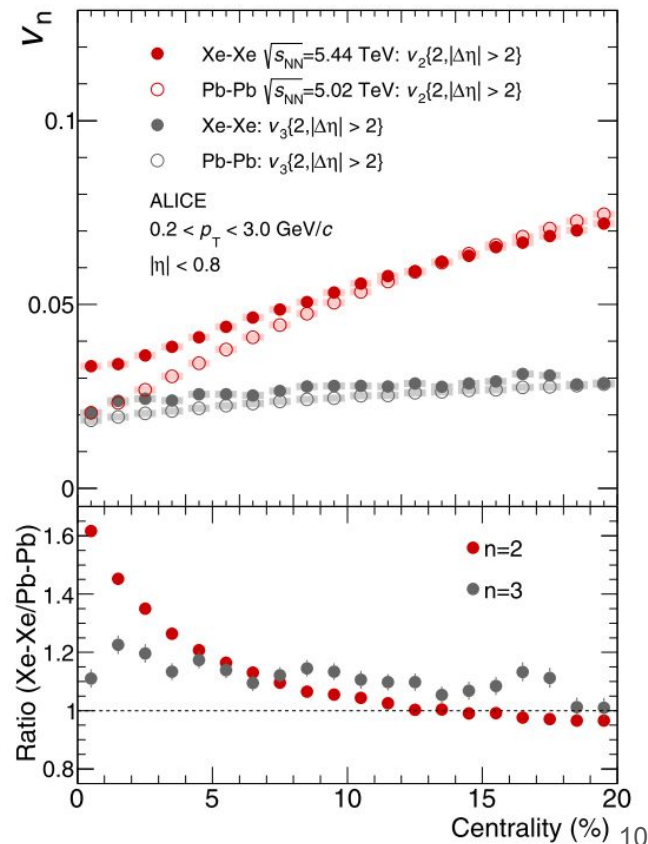
p-O studies help COSMIC AIR SHOWER MODELLING

2

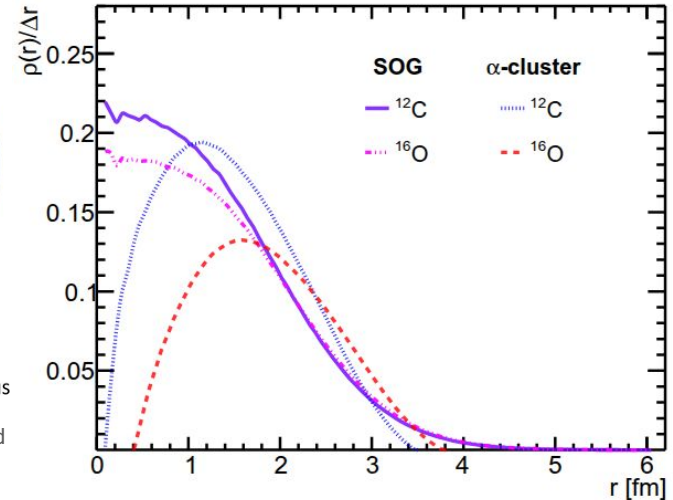
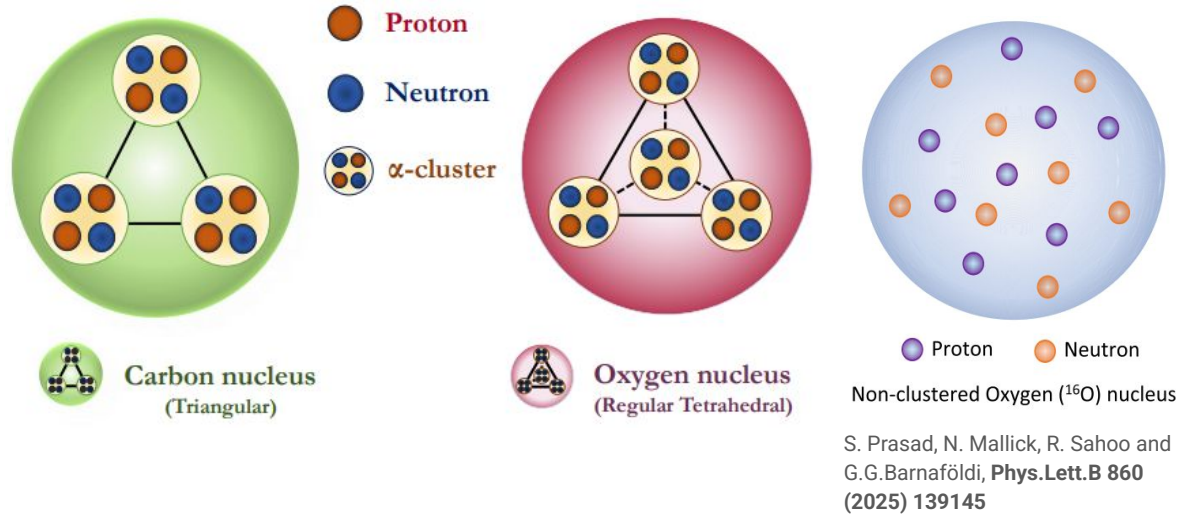
To investigate jet quenching effects & comprehend COLLECTIVE PHENOMENA in p-Pb

4

Effect of initial CLUSTERED geometry on final-state azimuthal correlations



α -clusters in O and C nuclei



α - cluster density profile

- ❖ ^4He nuclei with two protons and two neutrons is called an α -particle
- ❖ Light nuclei having $4n$ nucleons can possess **α -clustered** nuclear structure \rightarrow Ex.: ^8Be , ^{12}C , ^{16}O etc.
- ❖ **α -clustering** provides additional stability to nucleus

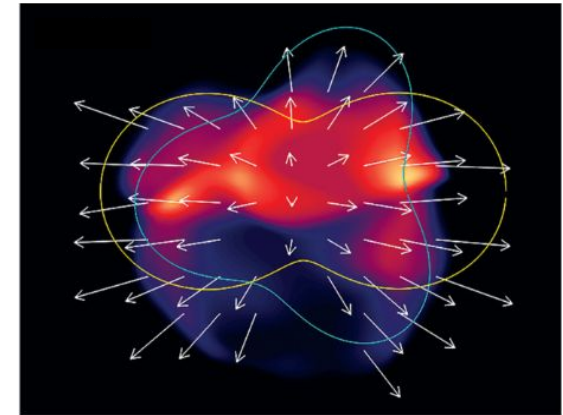
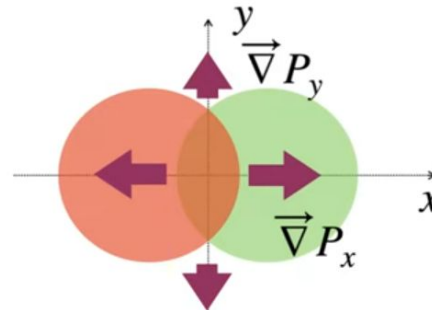
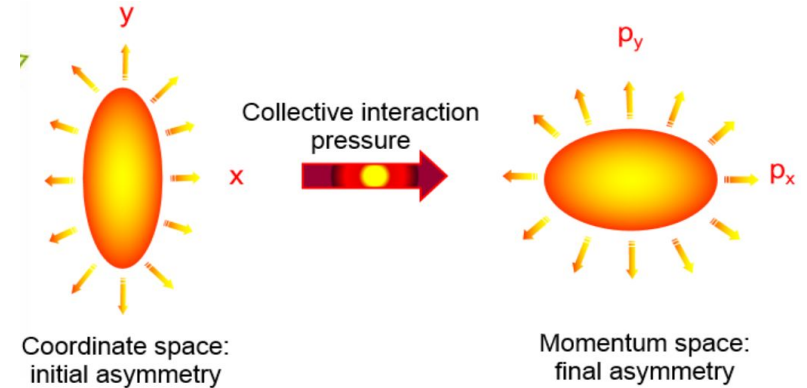
Anisotropic flow

- In non-central collisions, **spatial asymmetry** along different directions leads to **hierarchy of pressure gradients**: $\vec{\nabla} P_x \gg \vec{\nabla} P_y$
- Strong pressure gradients convert **initial spatial anisotropy** to final-state azimuthal **momentum space anisotropy**, via the collective expansion of the medium
- Anisotropic transverse expansion/**anisotropic flow** is quantified via coefficients of Fourier expansion of the azimuthal distribution of final state particles:

$$\frac{dN}{d\phi} = \frac{1}{2\pi} \left(1 + \sum_{n=1}^{\infty} 2v_n \cos[n(\phi - \psi_n)] \right)$$

where $v_n = \langle \cos[n(\phi - \psi_n)] \rangle$

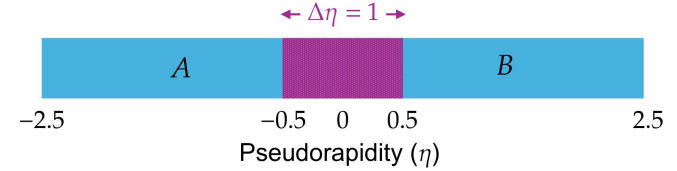
is the **n^{th} order anisotropic flow coefficient**



$v_2 \Rightarrow$ Elliptic Flow
 $v_3 \Rightarrow$ Triangular Flow

Anisotropic Flow Estimation

- In this study, estimation of v_n is done by **two-particle Q-cumulant method**
- Pseudorapidity gap in the sub-events helps in suppressing non-flow contributions



S. Prasad, N. Mallick, R. Sahoo and G.G. Barnaföldi, *Phys.Lett.B* 860 (2025) 139145

$$Q_n = \sum_{j=1}^M e^{in\phi_j}$$

$$\langle 2 \rangle = \frac{|Q_n|^2 - M}{M(M-1)}$$

$$c_n\{2\} = \langle \langle 2 \rangle \rangle = \frac{\sum_{i=1}^{N_{\text{ev}}} (W_{\langle 2 \rangle})_i \langle 2 \rangle_i}{\sum_{i=1}^{N_{\text{ev}}} (W_{\langle 2 \rangle})_i}$$

$$v_n\{2\} = \sqrt{c_n\{2\}}$$

$$p_n = \sum_{j=1}^{m_p} e^{in\phi_j} \quad q_n = \sum_{j=1}^{m_q} e^{in\phi_j}$$

$$\langle 2' \rangle = \frac{p_n Q_n^* - m_q}{m_p M - m_q}$$

$$d_n\{2\} = \langle \langle 2' \rangle \rangle = \frac{\sum_{i=1}^{N_{\text{ev}}} (w_{\langle 2' \rangle})_i \langle 2' \rangle_i}{\sum_{i=1}^{N_{\text{ev}}} (w_{\langle 2' \rangle})_i}$$

$$v_n\{2\}(p_T) = \frac{d_n\{2\}}{\sqrt{c_n\{2\}}}$$

$$\langle 2 \rangle_{\Delta\eta} = \frac{Q_n^A \cdot Q_n^{B*}}{M_A \cdot M_B}$$

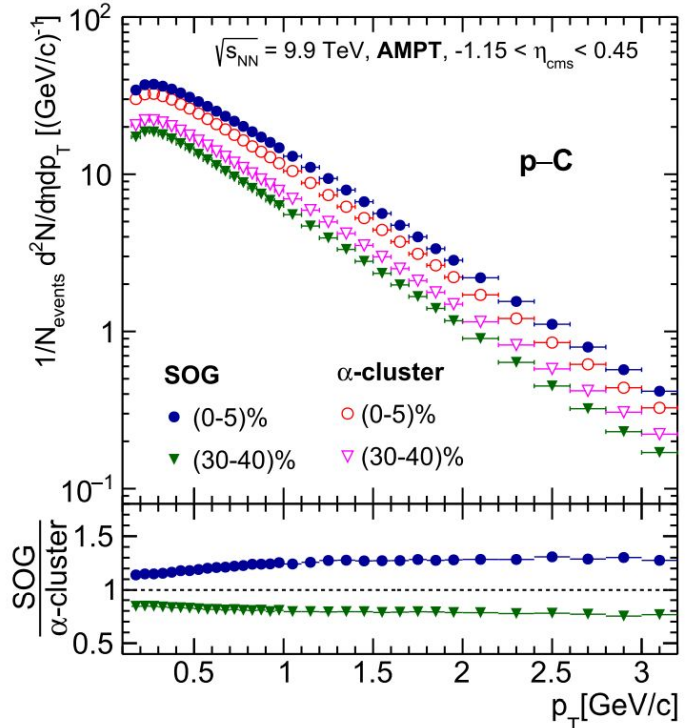
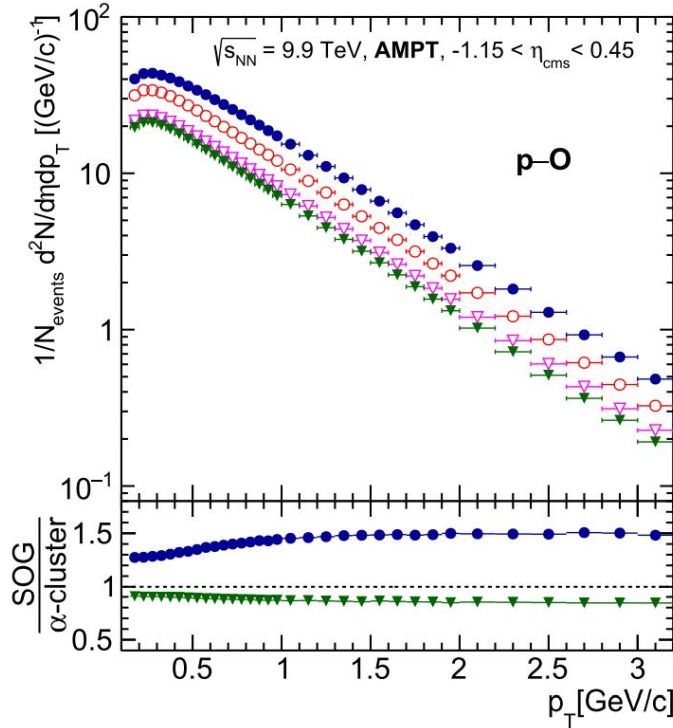
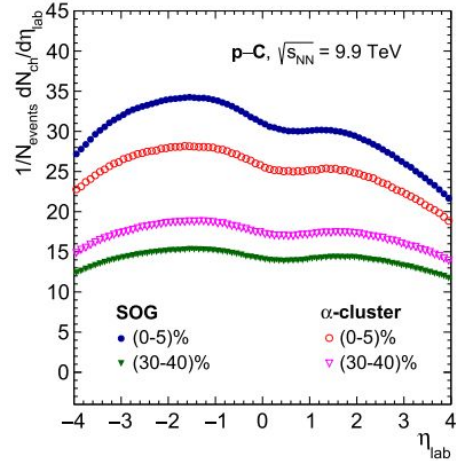
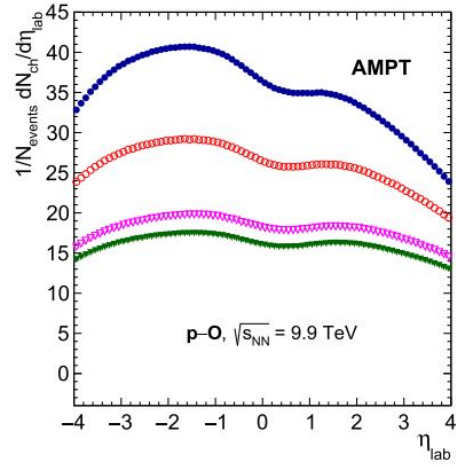
$$c_n\{2, |\Delta\eta|\} = \langle \langle 2 \rangle \rangle_{\Delta\eta}$$

$$\langle 2' \rangle_{\Delta\eta} = \frac{p_{n,A} Q_{n,B}^*}{m_{p,A} M_B}$$

$$d_n\{2, |\Delta\eta|\} = \langle \langle 2' \rangle \rangle_{\Delta\eta}$$

$$v_n\{2, |\Delta\eta|\}(p_T) = \frac{d_n\{2, |\Delta\eta|\}}{\sqrt{c_n\{2, |\Delta\eta|\}}}$$

Results

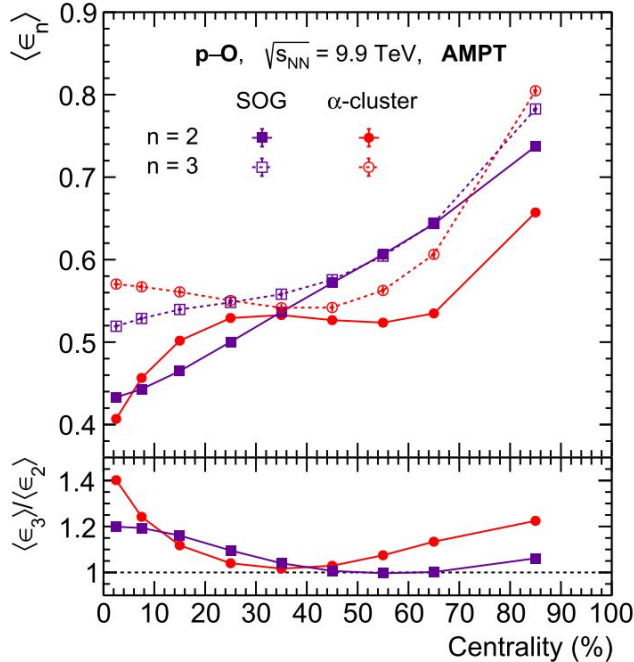


Results

Initial spatial anisotropies, such as, **eccentricity** (ϵ_2), **triangularity** (ϵ_3), etc., are quantified as:

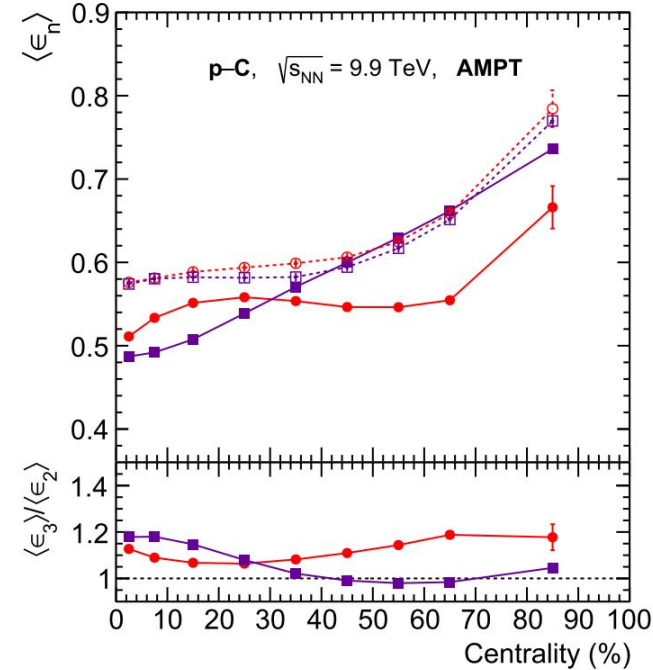
D. Behera, S. Prasad, N. Mallick and R. Sahoo,
Phys. Rev. D **108**, 054022 (2023)

$$\epsilon_n = \frac{\sqrt{\langle r^n \cos(n\phi_{\text{part}}) \rangle^2 + \langle r^n \sin(n\phi_{\text{part}}) \rangle^2}}{\langle r^n \rangle}$$

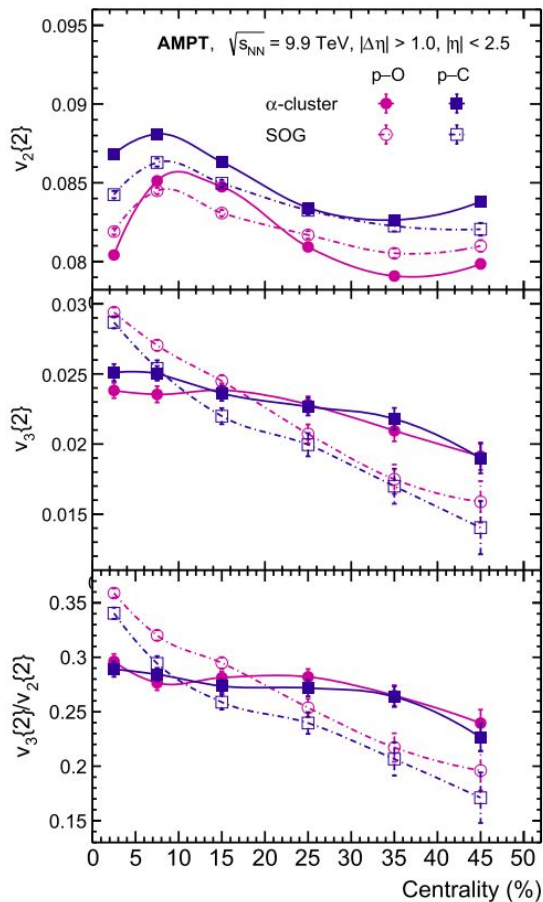


★ Trends of $\langle \epsilon_n \rangle$ and $\langle \epsilon_3 \rangle / \langle \epsilon_2 \rangle$ for α -cluster profiles are closely similar to that in O-O collisions at $\sqrt{s_{\text{NN}}} = 7$ TeV

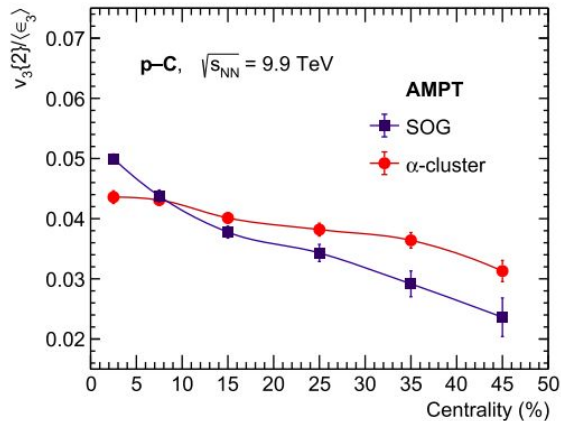
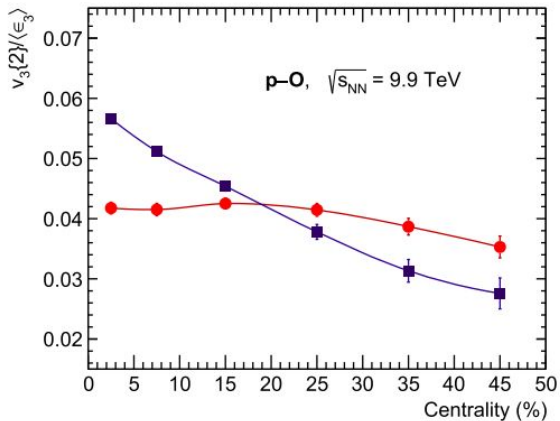
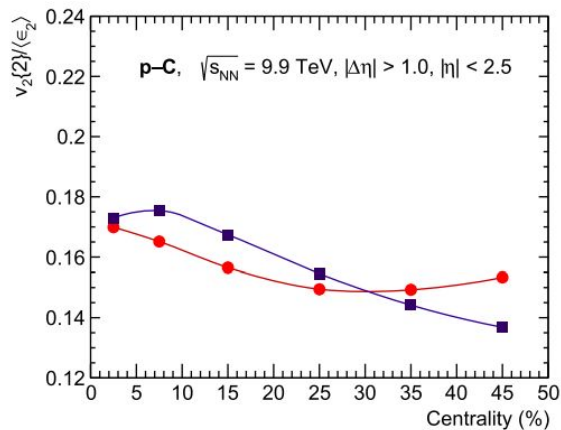
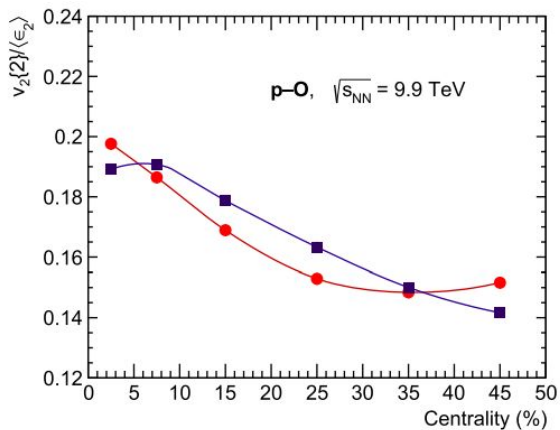
★ $\langle \epsilon_3 \rangle / \langle \epsilon_2 \rangle$ shows a hike at most central p-O collisions (absent in p-C)



Results

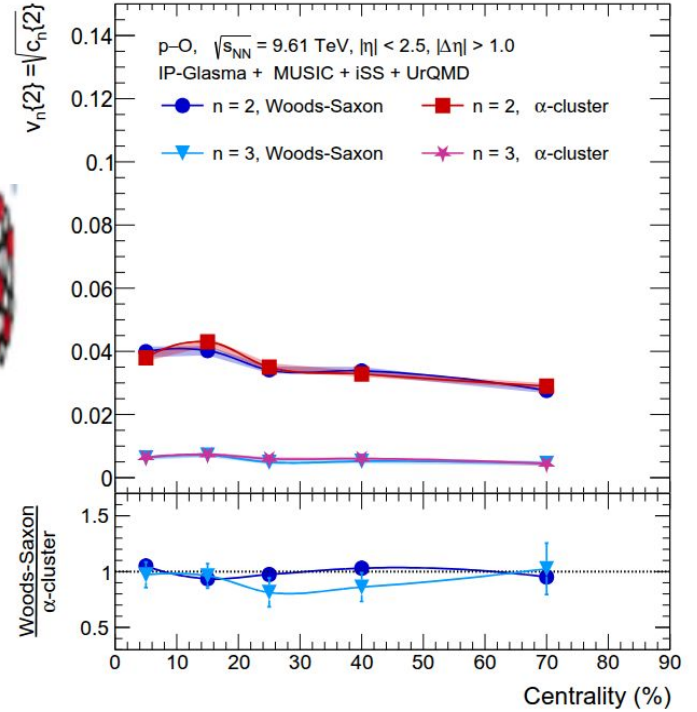
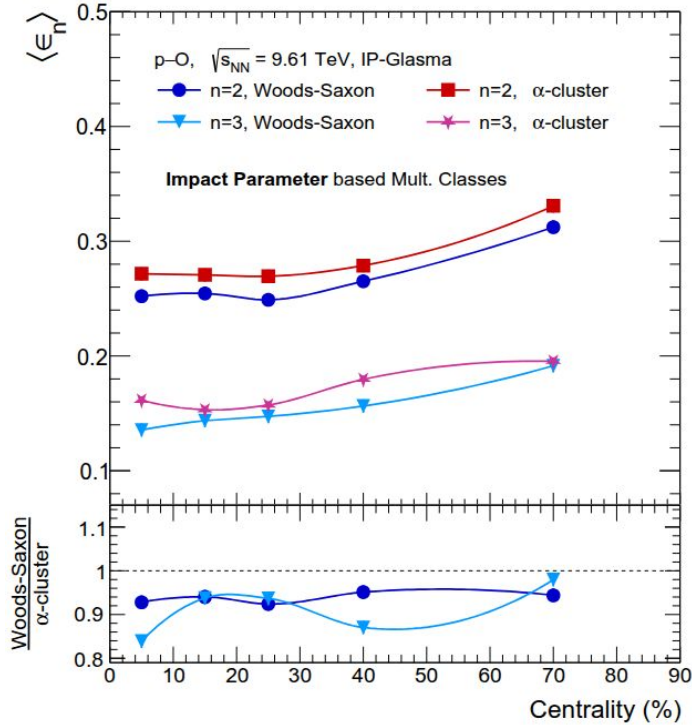


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Eur. Phys. J. A 61, 134 (2025)



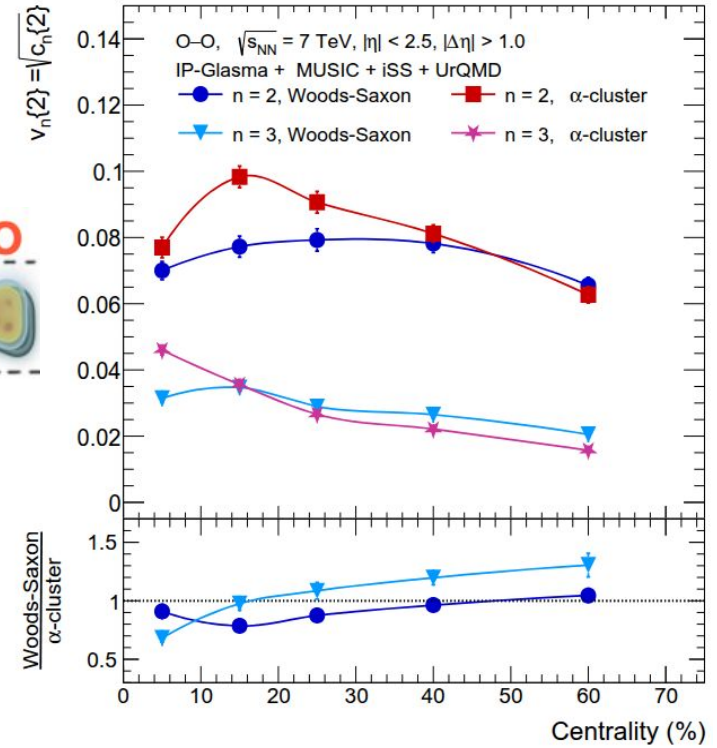
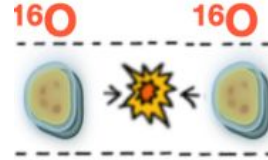
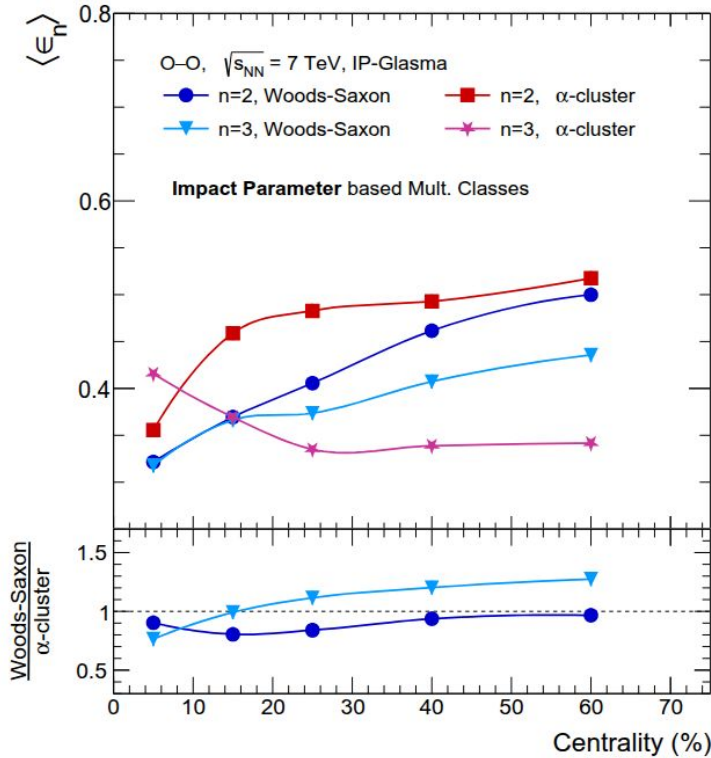
Results

$$\epsilon_n = \frac{\sqrt{(\iint_A r^n \epsilon(x, y) \cos(n\phi_{\text{part}}) dx dy)^2 + (\iint_A r^n \epsilon(x, y) \sin(n\phi_{\text{part}}) dx dy)^2}}{\iint_A r^n \epsilon(x, y) dx dy}$$



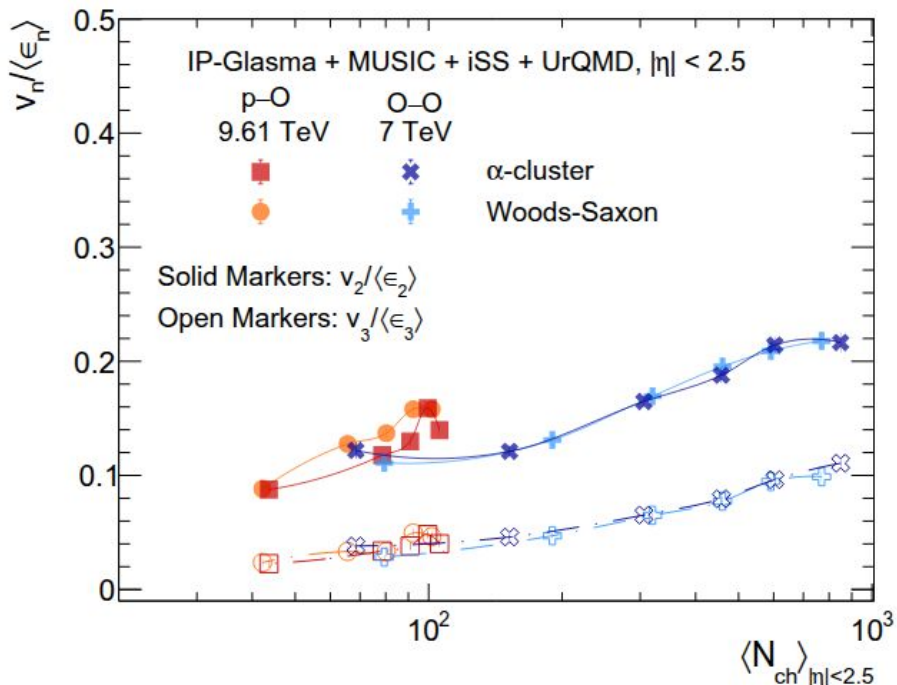
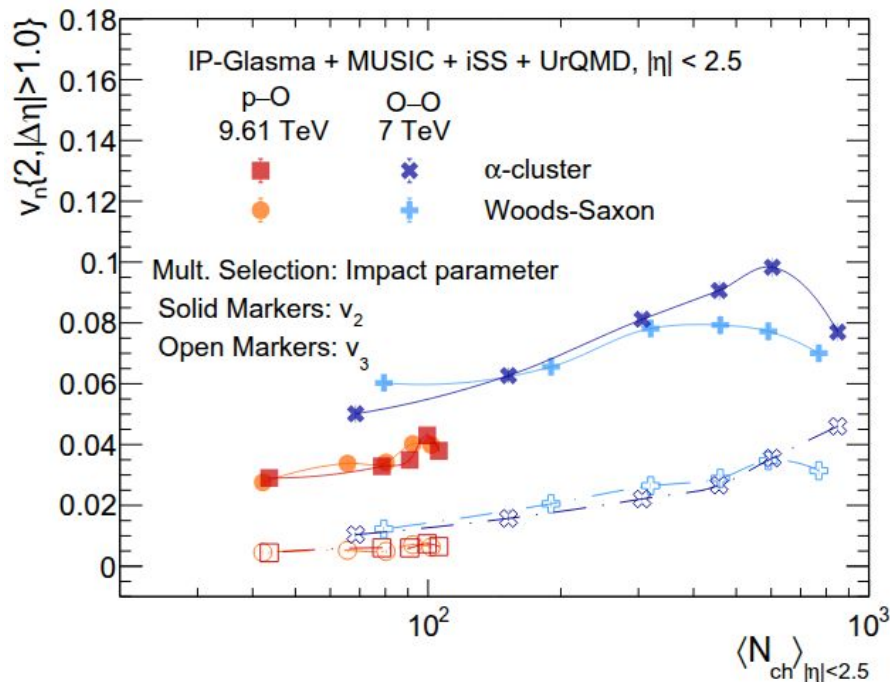
Results

A. Menon Kavumpadikkal Radhakrishnan, S. Prasad, N. Mallick, R. Sahoo and G. G. Barnaföldi, arXiv:2505.22367 [hep-ph]



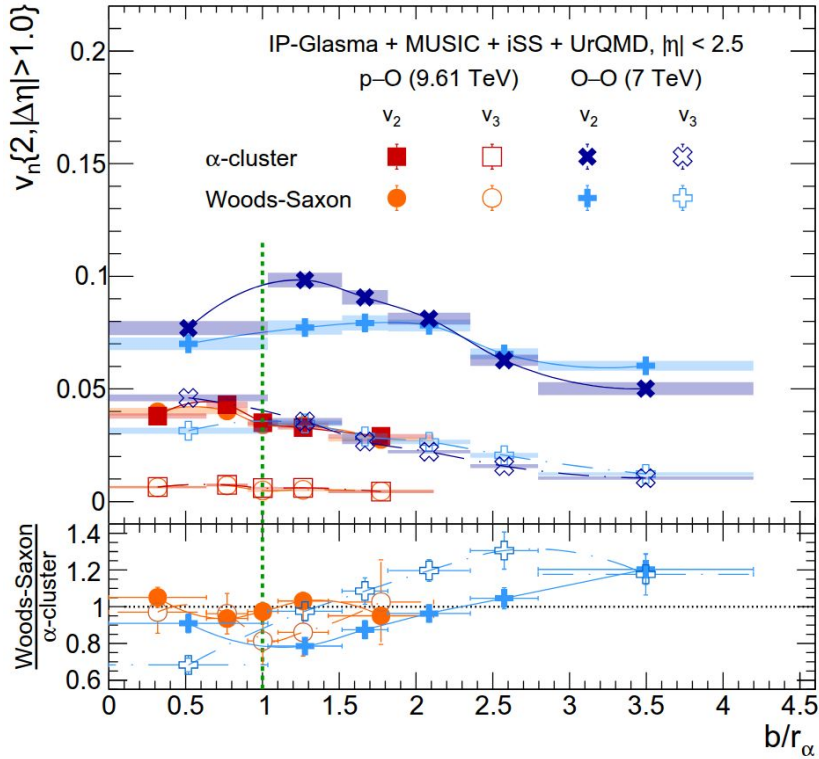
★ Distinction due to nuclear density profiles is better evident in OO than from pO collisions.

Results

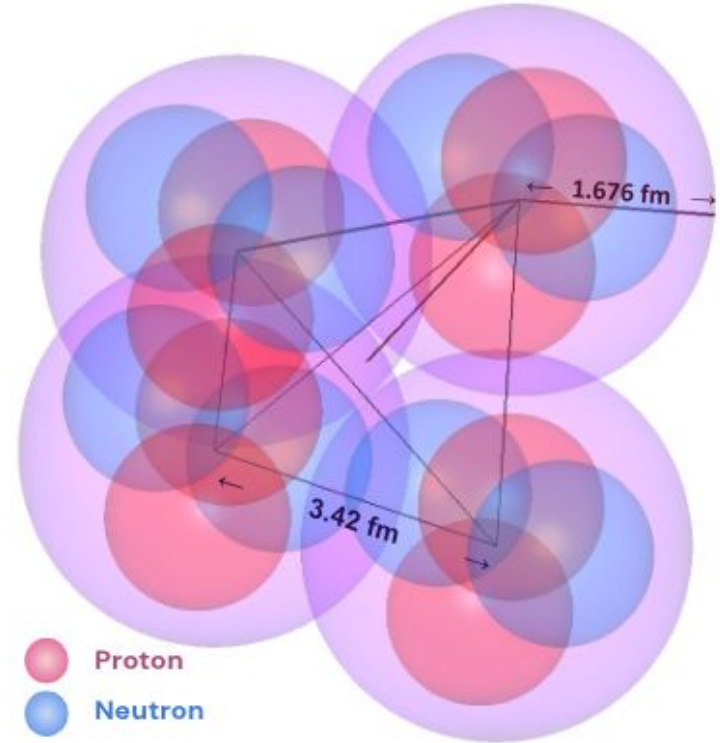


- ★ Effects from α -clustering gets pronounced at higher multiplicities (10-20% centrality)
- ★ Response coefficient is multiplicity-dependent than being collision-system-dependent?

Results



A. Menon Kavumpadikkal Radhakrishnan, S. Prasad, N. Mallick,
R. Sahoo and G. G. Barnaföldi, arXiv:2505.22367 [hep-ph]



- ★ The effects of α -clustering in O–O collisions are observed to manifest well in the region $b/r_\alpha \lesssim 1$; therefore it is comparable with the size of the ^4He !

Summary and Outlook

- For the first time, a systematic study on the effects of initial nuclear density profiles (α -cluster nuclear geometry especially) on final state flow coefficients is reported for **pO** and **pC** collisions using AMPT model and for **pO** and **OO** collisions (from impact parameter based centrality classification) using IP-Glasma+MUSIC+iSS+UrQMD at LHC energies
- **α -cluster profile maintains a similar but unique qualitative behavior** throughout the collision systems, pC, pO and OO (AMPT)
- Effects of α -clustering are more prominent in OO collisions than pO and they are **pronounced in (10-20%)** centrality class. Proton might not be a good probe to study α -cluster structure
- α -clustering effects might **manifest well near $b/r_\alpha \lesssim 1$** \rightarrow needs further confirmation from other realistic nuclear collisions (C-C, Ne-Ne etc.)

Work is in progress to look into the effects of the compactness of α -clustering on the final state v_n and to extend our study to Ne-Ne collisions...



Thank



You :)

