

# MUOGRAPHERS 2026



## Portable Glass-RPC Telescope for Muography and Future Applications at Laser Wakefield Accelerators

Sumaira Ikram<sup>1</sup>, A. Giammanco<sup>1</sup>, M. Lagrange<sup>1</sup>, E. Cortina Gil<sup>1</sup>, P. Demin<sup>1</sup>, R. Versaci<sup>3</sup>, A. Cimmino<sup>3</sup>, G.M. Grittani<sup>3</sup>, M. Nevrkla<sup>3</sup>, D. Ahmadi<sup>2</sup>, M. Tytgat<sup>2</sup>

1. Centre for Cosmology, Particle Physics and Phenomenology, UCLouvain, Louvain-la-Neuve, Belgium

2. Inter-University Institute for High Energies (IIHE), Vrije Universiteit Brussel, Brussels, Belgium

3. ELI Beamlines Facility, The Extreme Light Infrastructure ERIC, Dolni Brezany, Czechia



# Portable Particle Detector

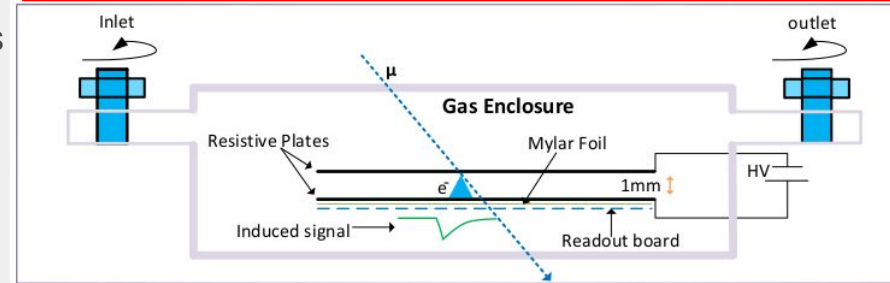
Portable muon telescopes are mainly used for experiments conducted in confined or restricted environments.

The design goal of the detector includes:

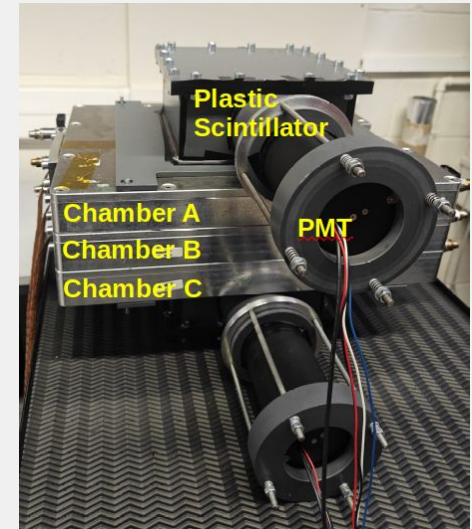
- Portability
  - Autonomous Operation
  - Gas Tightness (In gaseous Detectors)
  - Robustness
  - Modular Geometry
  - Safe and Cheap
- 
- **Glass electrodes with an active area of  $16 \times 16 \text{ cm}^2$**
  - **Detectors have been in operation for  $\sim 2$  years with good efficiency.**

For more details regarding the detector setup, presentation by Donya Ahmadi, [\*A Sealed, Light-weight GRPC-Based Muon Telescope for Muography Applications\*](#)

Schematic representation of the glass-electrode resistive plate chamber



S. Ikram *et al.*, "Development and performance analysis of glass-based gas-tight RPCs for muography applications, *J. Appl. Phys.*, vol. 138, no. 17, p. 174502, Nov. 2025.



Experimental setup for efficiency measurement

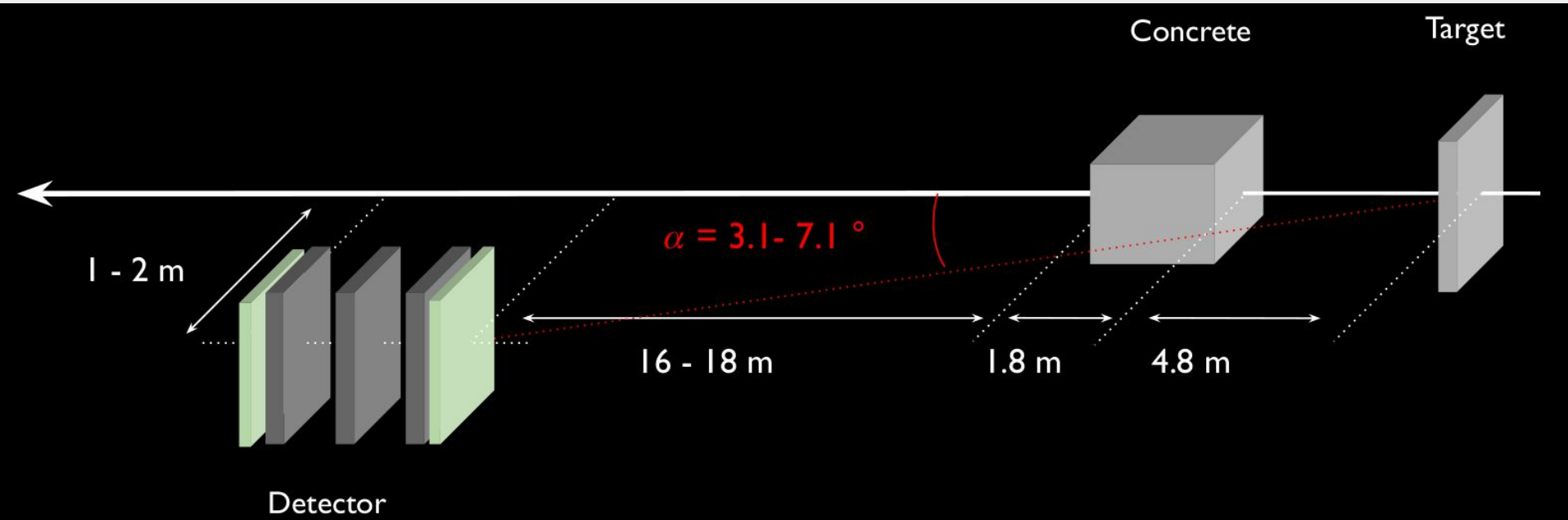
# Portable System Validated

- **ELI (The Extreme light infrastructure)** → One of the world's largest and most advanced high-power laser infrastructure and a global technology.
- **ELBA (Electron Beamline for Fundamental Science)** → a laser–electron collider and powered by the high-repetition-rate L3-HAPLS laser system.



- We successfully transported and operated our detector to the test site (Prague) in April/May (two weeks) and then in August (12 days)/September (10 days).

More details about ELI beams were presented in Roberto Versaci's talk: [Overview of Laser-Driven Multi-GeV Muon Sources and the ELI Beamlines Plan.](#)



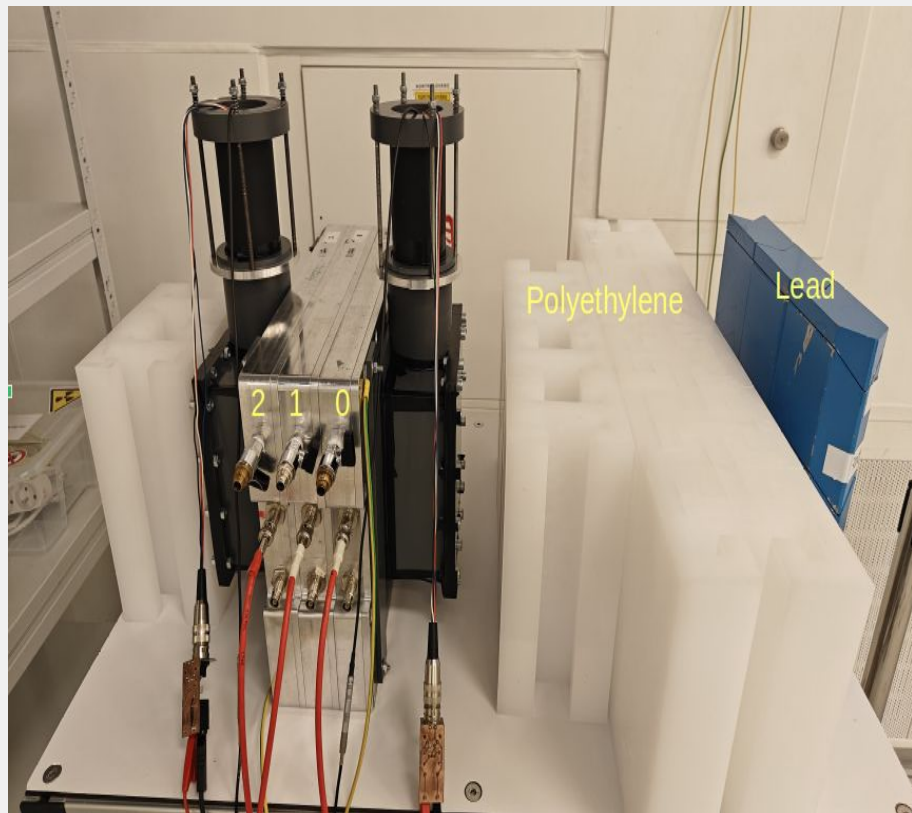
Schematic of the experimental configuration and geometry used to define the detector system in the experimental hall.

# Overview of Data-Taking

- **Cosmic Data (D1, D2, D3)** – baseline reference for detector performance
- **High-Energy Beam (D1, D2)** – ELBA-validation under intense particle flux
  - Multi-GeV electron beams
- **Low-Energy Beam (D3)** – GAMMATRON-study of detector response at
  - Energy  $\sim 1$  GeV.
  - Beam was present but muon production is negligible,
  - Characterizing the beam-induced background.

## Shielding Configuration

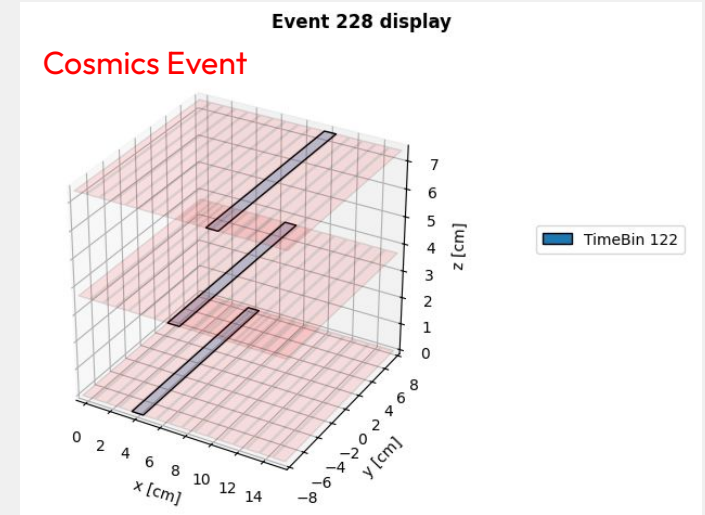
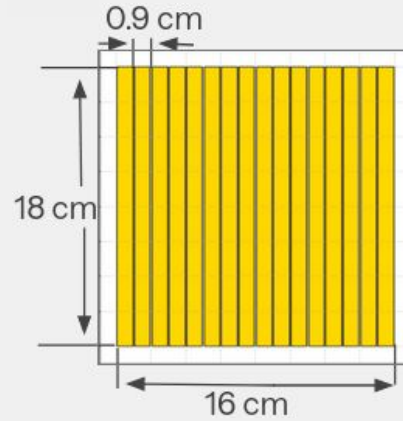
- No Shielding
- Lead
- Polyethylene
- Lead + polyethylene



**Detector setup during April-May (left) and August-September 2025 (right) Beam test**

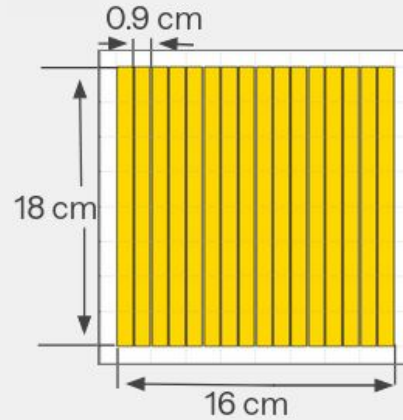
# Event Display Across Datasets

- Cosmic run: ~1 fired strip per layer per event



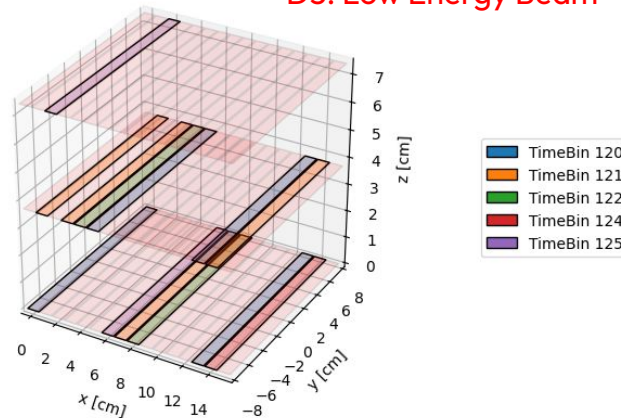
# Event Display Across Datasets

- Cosmic run: ~1 fired strip per layer per event
- D3 low energy beam data: Higher strip multiplicity



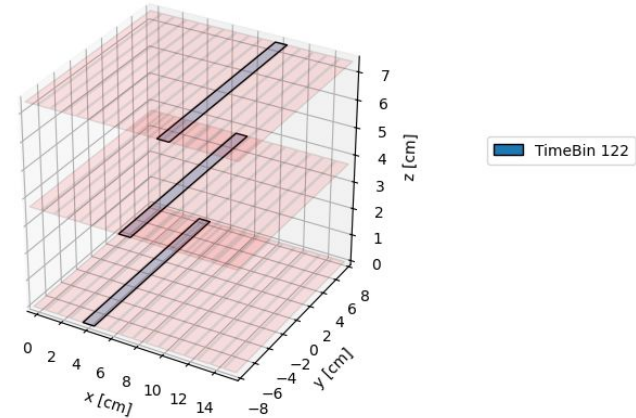
Event 331 display

D3: Low Energy Beam



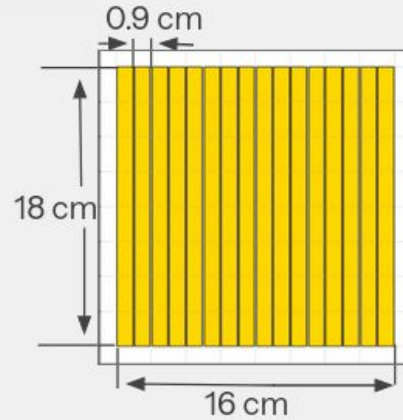
Event 228 display

Cosmics Event

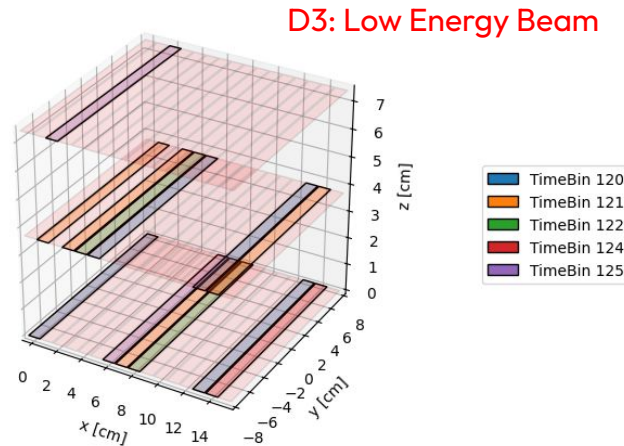


# Event Display Across Datasets

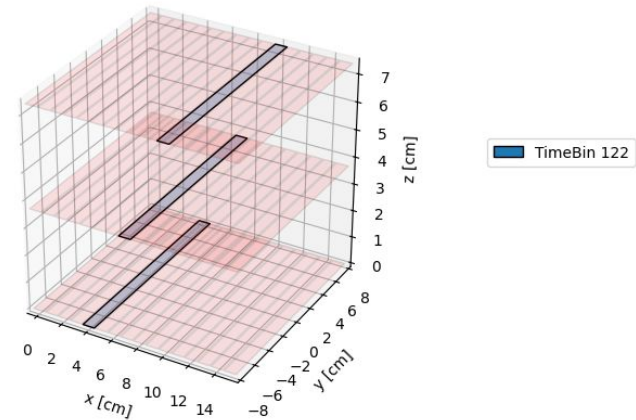
- Cosmic run: ~1 fired strip per layer per event.
- D3 low energy beam data: Higher strip multiplicity.
- High Energy Beam : Up to 9 fired strips observed in the first detector layer.



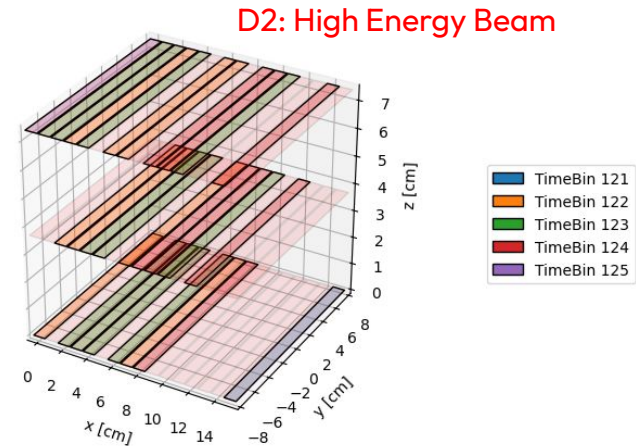
Event 331 display



Event 228 display  
**Cosmics Event**

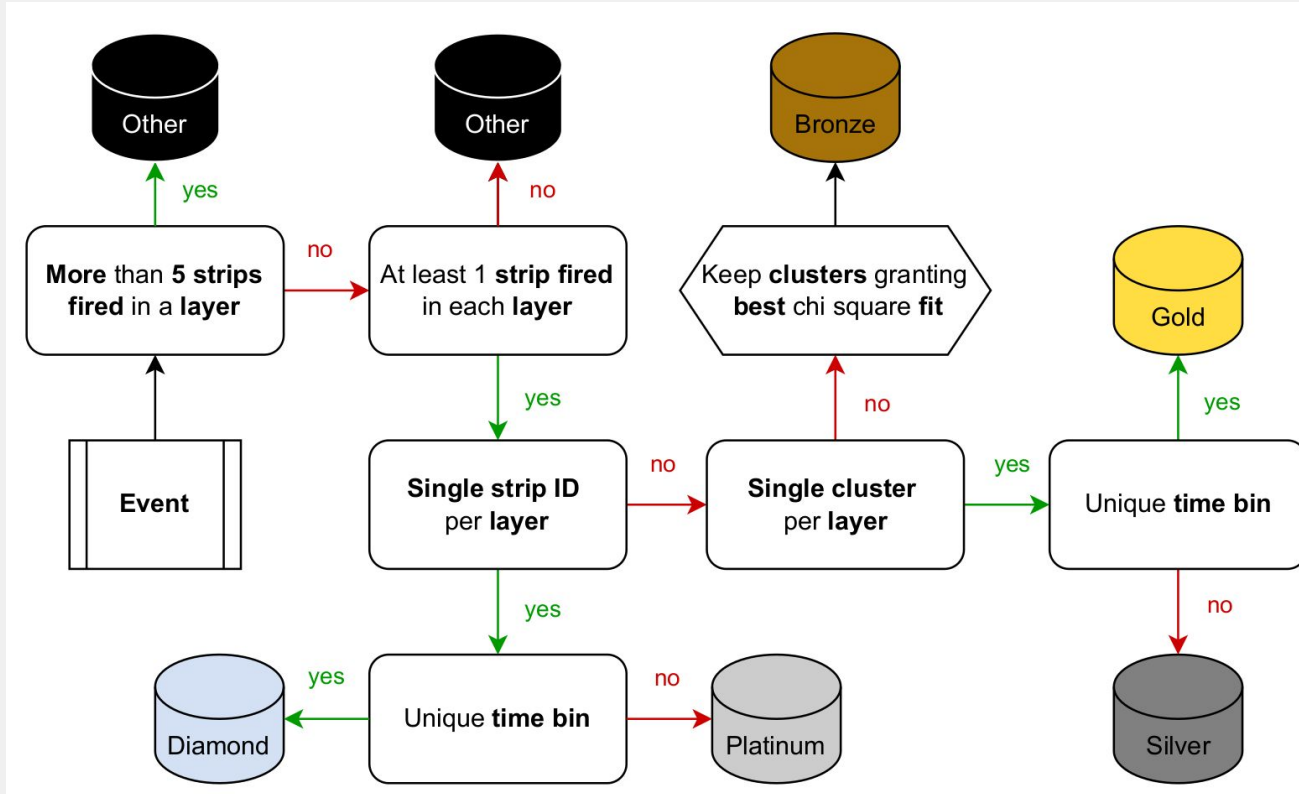


Event 228 display



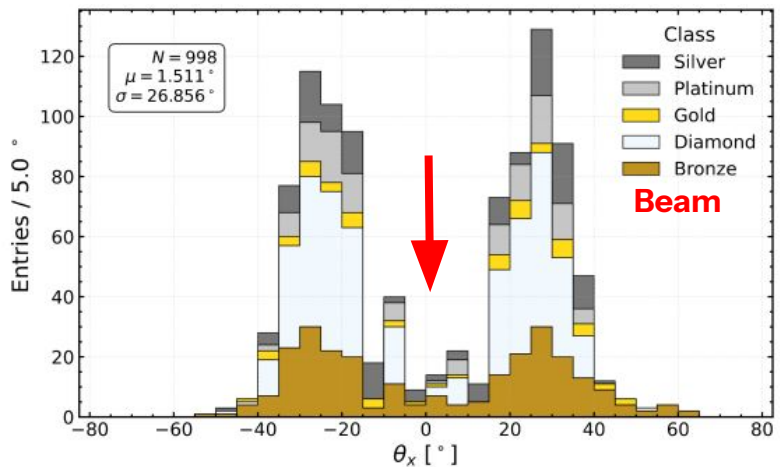
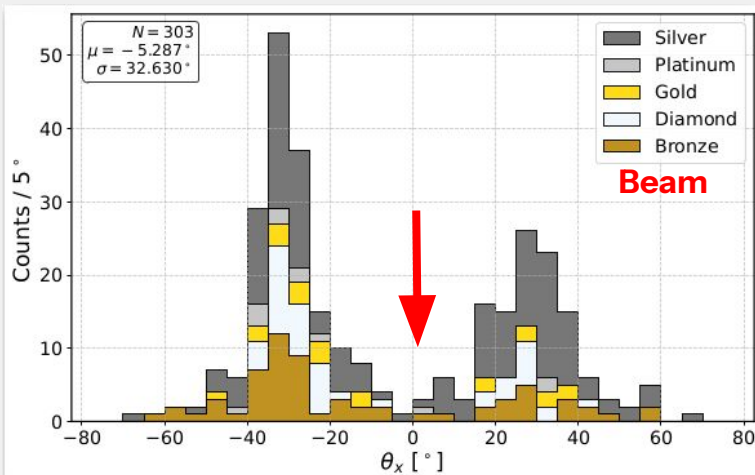
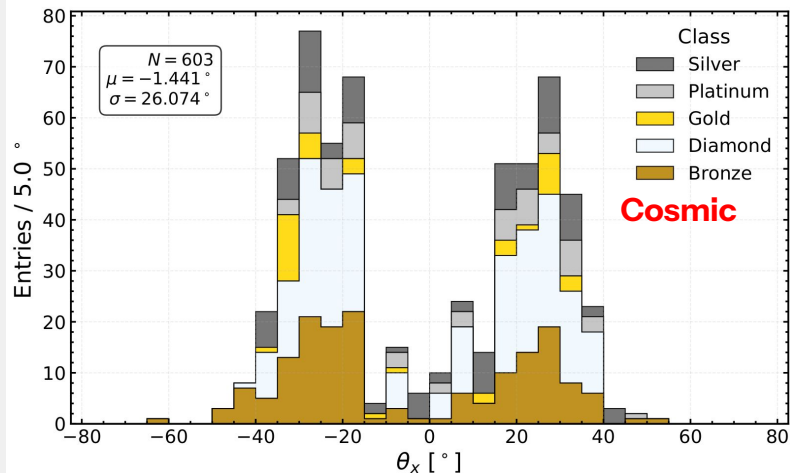
# Event Selection

- Events are classified into five categories based on the
  - Number of fired strips and/or clusters per layer
  - Temporal distribution, as well as a final exclusion class.



# Angular Distribution

- Angular distribution consistent with cosmic muon behavior.
- Beam-induced muons are expected to show predominantly horizontal trajectories  
Expected angular range:  $\theta_x \in [-5^\circ, 5^\circ]$ .

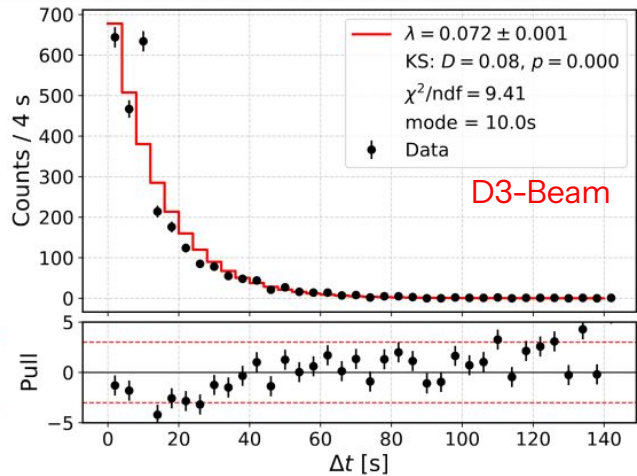
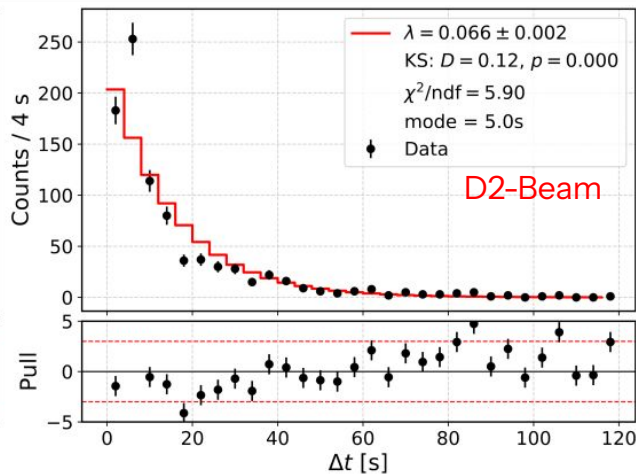
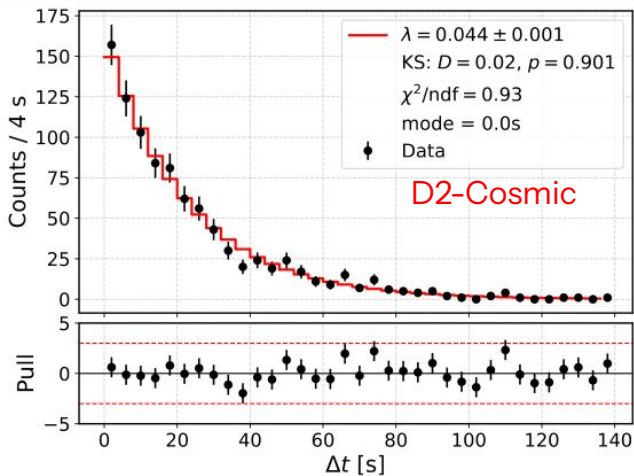


- Overall tracking results do not conclusively confirm beam-induced muons

# Inter-arrival time

Exponential fit applied to  $\Delta t$  distributions

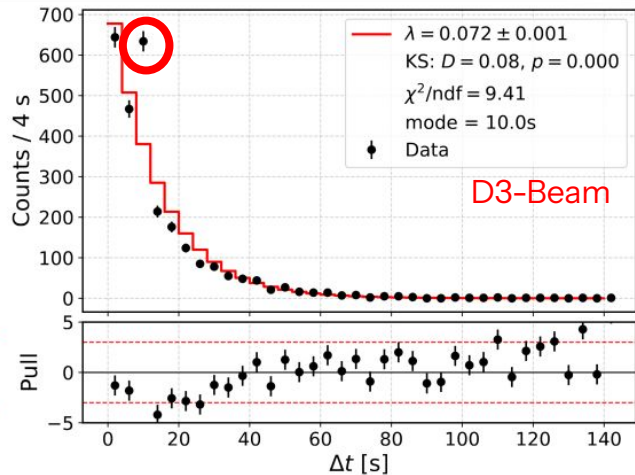
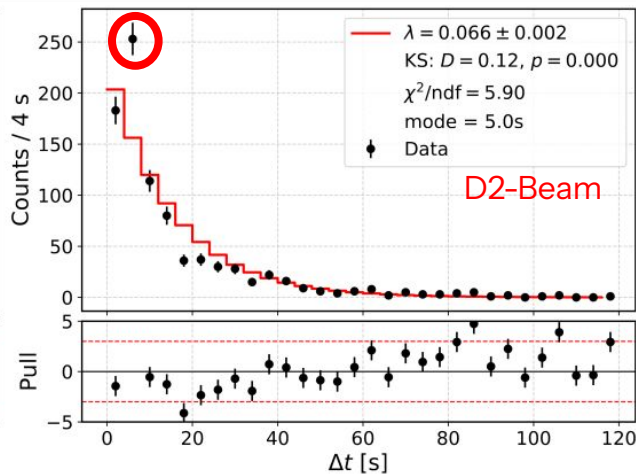
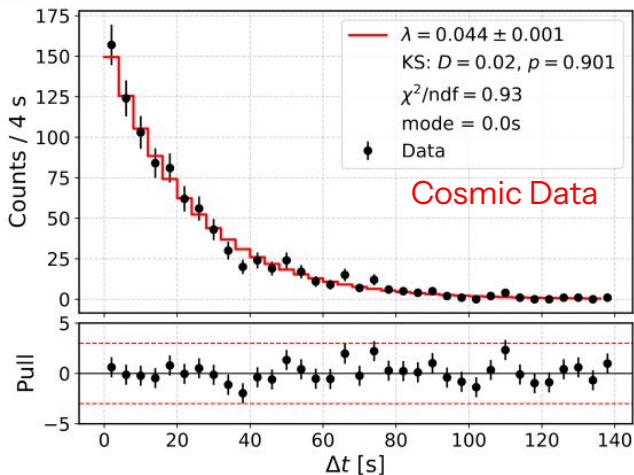
- D2 cosmic  $\rightarrow$  consistent with random arrival time (Poisson)
- D2 beam  $\rightarrow >5\sigma$  deviation
- D3 beam  $\rightarrow >5\sigma$  deviation



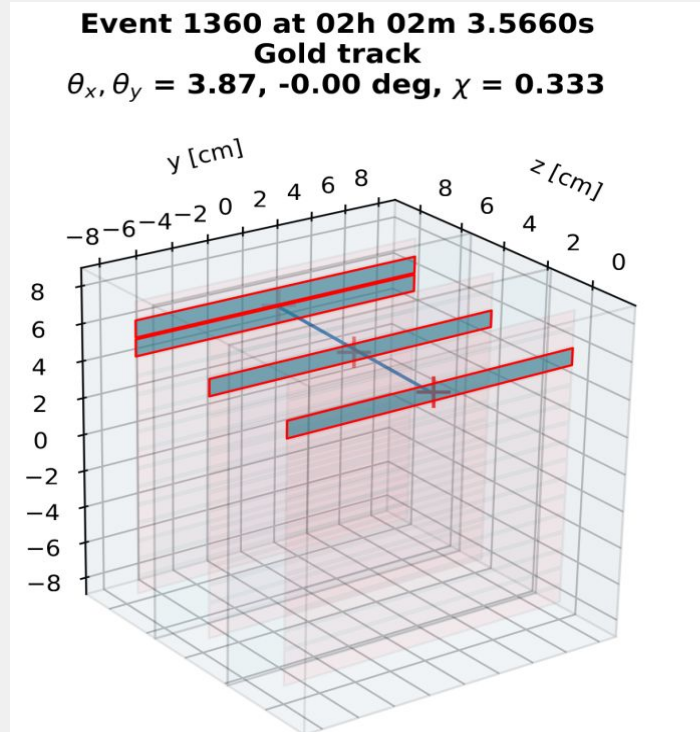
# Inter-arrival time

Exponential fit applied to  $\Delta t$  distributions

- D2 cosmic  $\rightarrow$  consistent with random arrival time (Poisson)
- D2 beam  $\rightarrow >5\sigma$  deviation
- D3 beam  $\rightarrow >5\sigma$  deviation
- Both spikes correspond exactly to the repetition rates of the lasers in those runs  $\rightarrow$  cause-effect relationship



# Beam muon candidates



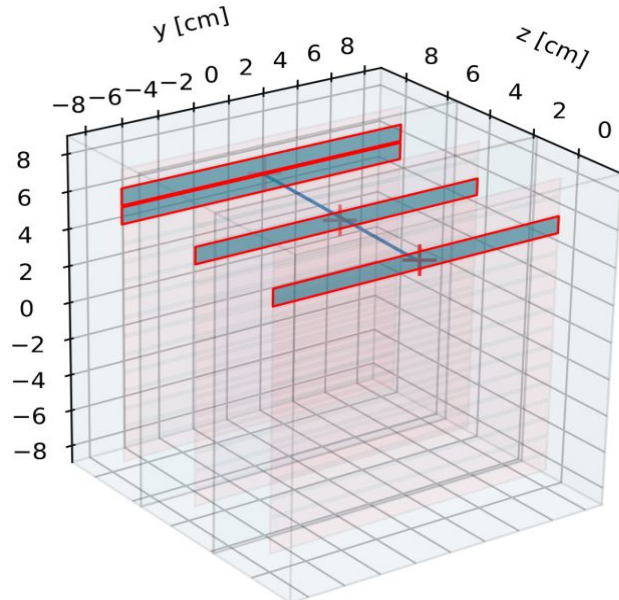
- Beam induced muon candidate from the D2 beam data-taking.

# Beam muon candidates

**Event 1360 at 02h 02m 3.5660s**

**Gold track**

$\theta_x, \theta_y = 3.87, -0.00$  deg,  $\chi = 0.333$

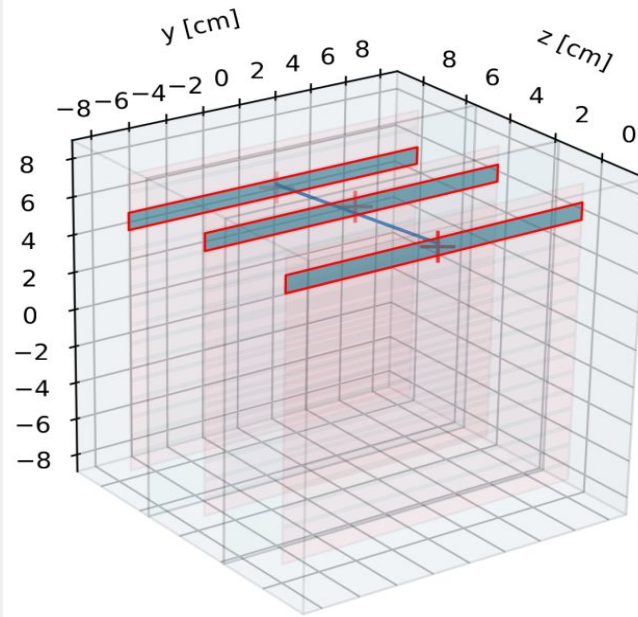


- Beam induced muon candidate from the D2 beam data-taking.

**Event 2086 at 07h 23m 57.0228s**

**Diamond track**

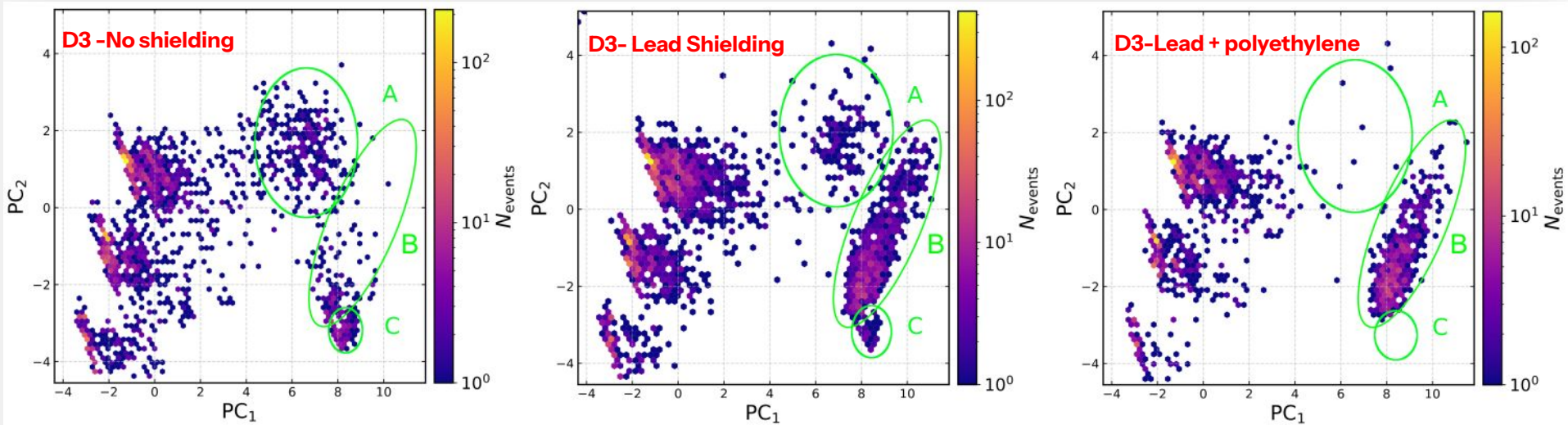
$\theta_x, \theta_y = -7.74, -0.00$  deg,  $\chi = 0.661$



- Beam-induced candidate observed in the D3 beam dataset, potentially corresponding to either a cosmic muon or an electron.

# Principal Component Analysis (PCA) for data categorization

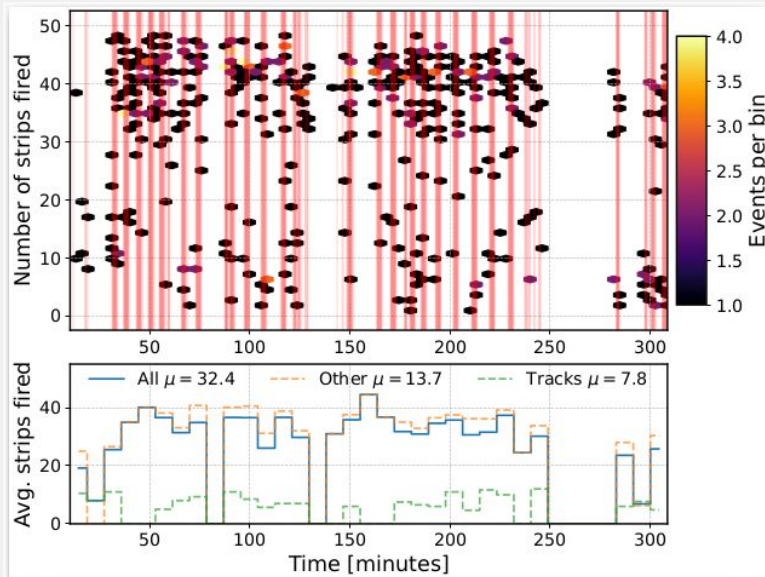
- Identifies dominant patterns and correlations while reducing data complexity for easier visualization and interpretation.
- Inputs: multiplicity, position and time information at strip and cluster levels.



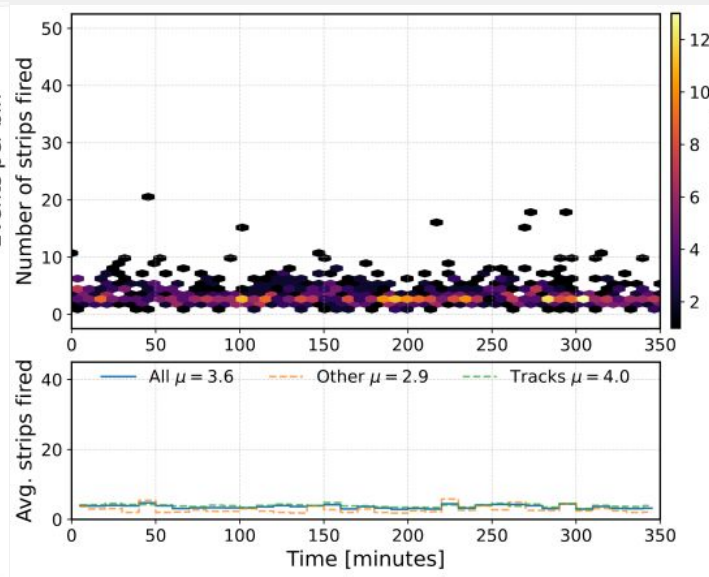
- **Region A** → moderate cluster multiplicity and relatively low strip counts,
- **Region B** → large strip multiplicity with fewer clusters
- **Region C** → very high strip multiplicity.

# Laser Tag-D2 Beam Data

- 5.5 h additional D2 beam run with recorded laser trigger
- Laser tag directly indicates beam status
- Laser ON: muon beam present
- No laser tag: beam absent



**Laser Tagged**



**Non tagged**

# Summary

- **Successful Beam Test at ELI in 2025**

- BeamTest at ELI validated the **portability and robustness** of the prototype in real operational conditions.

- **Data Analysis**

- Beam events are **high-multiplicity events**, making tracking difficult.
- **No statistically significant** excess of muons with horizontal direction was found.
- **Muons candidates were identified** in each data set.

- **Beam Test at ELI in 2026**

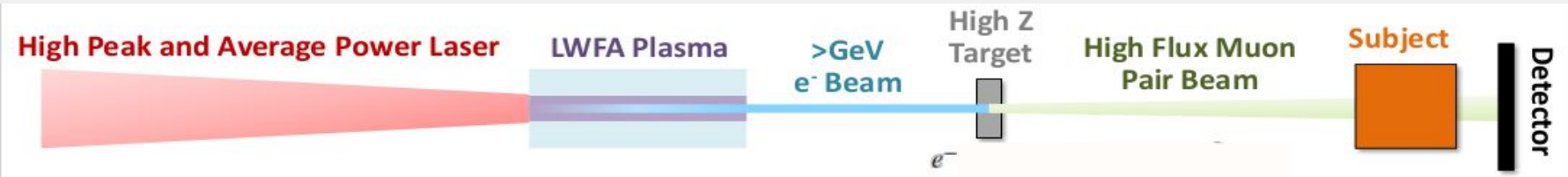
- Aiming at first observation of muon production at ELI, for long beam time as **main users**.

**Thank you..**

# Backup slides

# Muon Production via Laser Accelerated Electrons

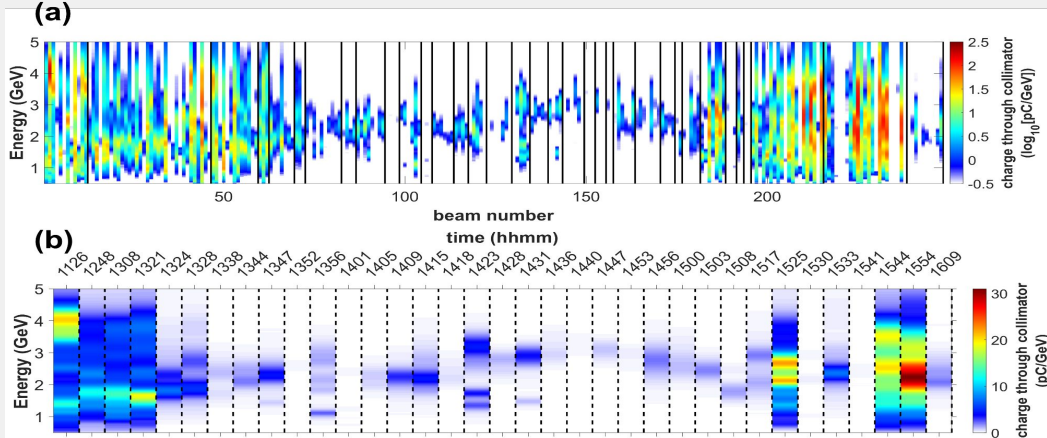
- Muons could be a game changing technology if there were a compact active source of muons available.
- Muons produced via photo-production or electro-production are characterized by their short pulse duration and high flux.



Conceptual layout of the muon source.

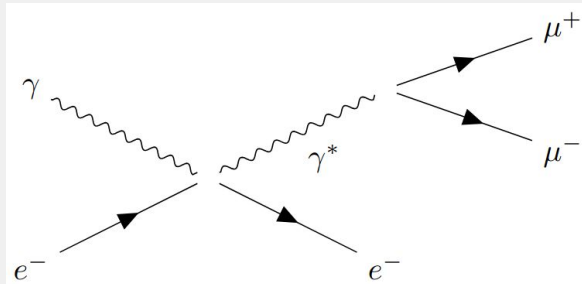
Image Source: "Reagan ELI-BL Muon Source Workshop".

Example of beams produced during a typical day of parasitic operation. (a) All angle-integrated spectra collected during a day of operation at 0.2 Hz repetition rate. The solid lines denote periods of operation when the operating conditions were fixed. (b) Mean angle-integrated collected for fixed operating conditions noted in panel (a). The top axis labels show the time each data set was collected at a 0.2 Hz repetition rate.

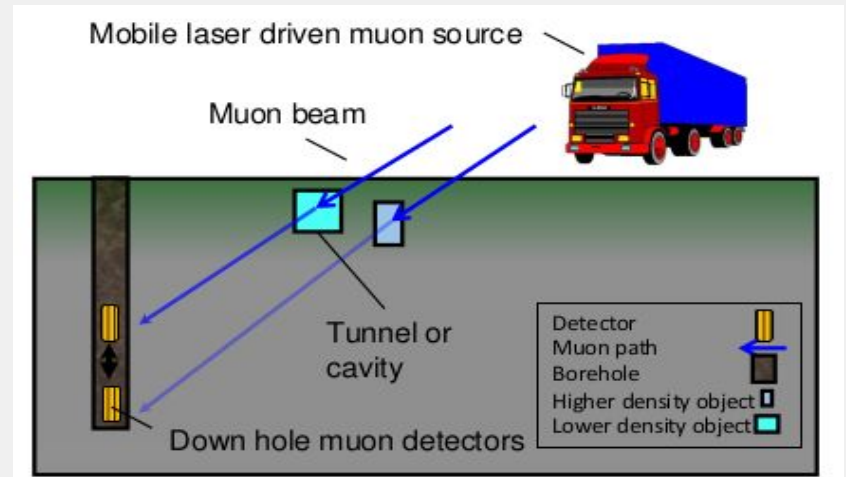


# Potential Benefits of Artificial Muons

- **Higher Particle Flux:** The flux is significantly higher than that of cosmic muons.
- **Controllable Energy:** Unlike cosmic muons, whose energies vary unpredictably, beamline allows for control over the energy of the beam.
- **Known Directionality:** The beam direction is well defined, unlike cosmic muons which arrive from random directions.
- Muons are not hazardous to life forms or the environment, provided the flux is not extremely high.
- Muons can be used to detect many different materials






Feynman Diagram for Muon pair production



Source: Todd Ditmire "Laser Induced Muon Production for Detecting Special Nuclear Materials and WMD"

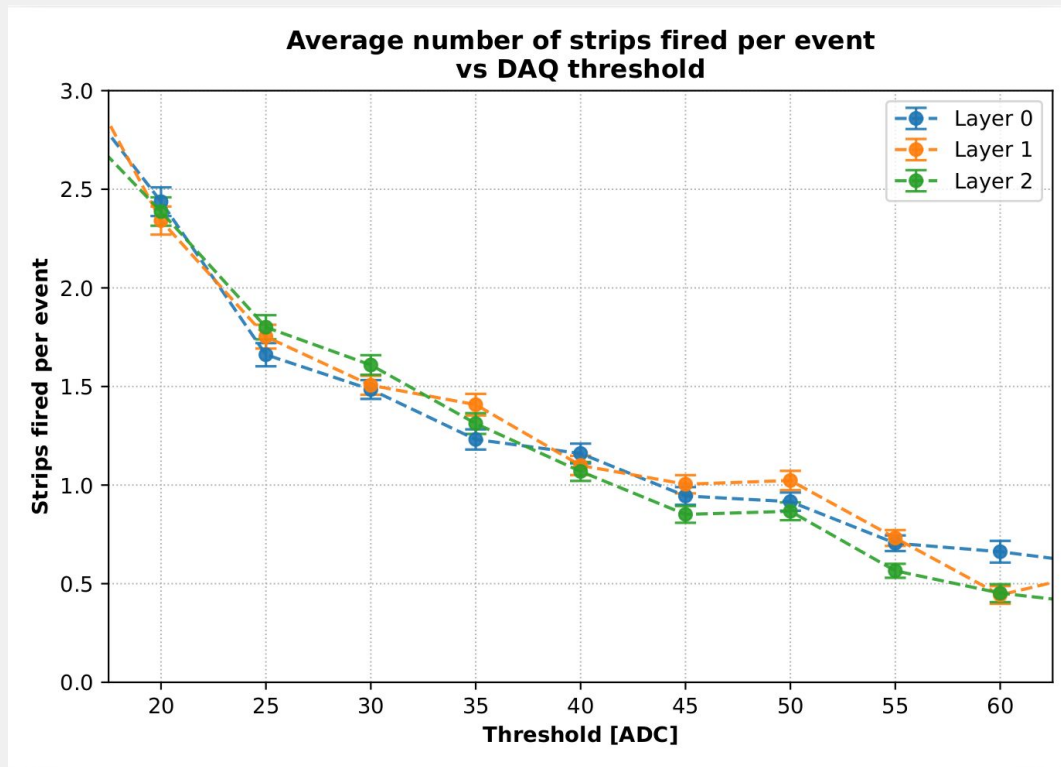
# Overview of Data-Taking

## Shielding Configurations and Durations

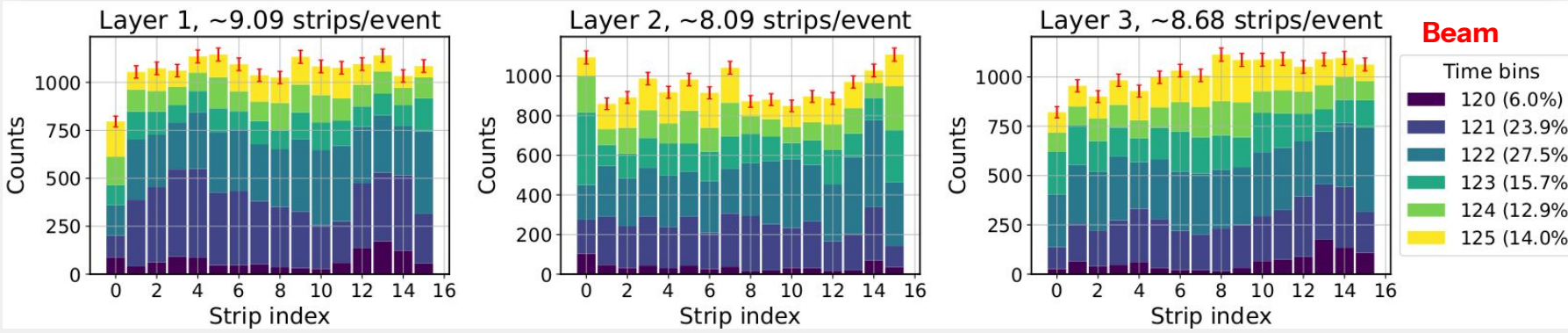
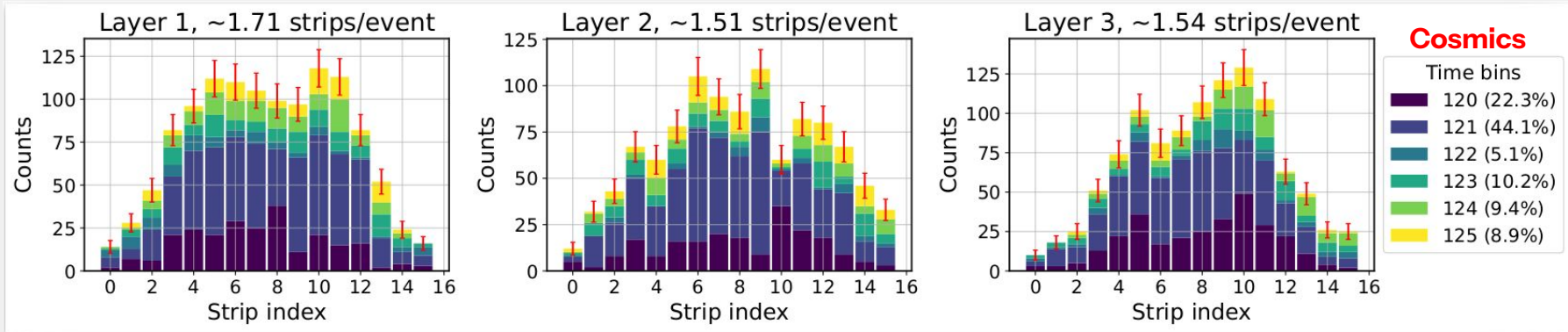
 Configuration	 Shielding	 Duration
D1 Cosmic	Polyethylene	4 h
D1 Beam	Polyethylene	28 h
D2 Cosmic	None	6 h
D2 Beam	None	3 h 45 min
D2 Beam	Lead + Polyethylene	5 h 30 min
D3 Beam	None	10 h 30 min
D3 Beam	Lead	15 h
D3 Beam	Lead + Polyethylene	7 h 30 min

# Threshold Scan

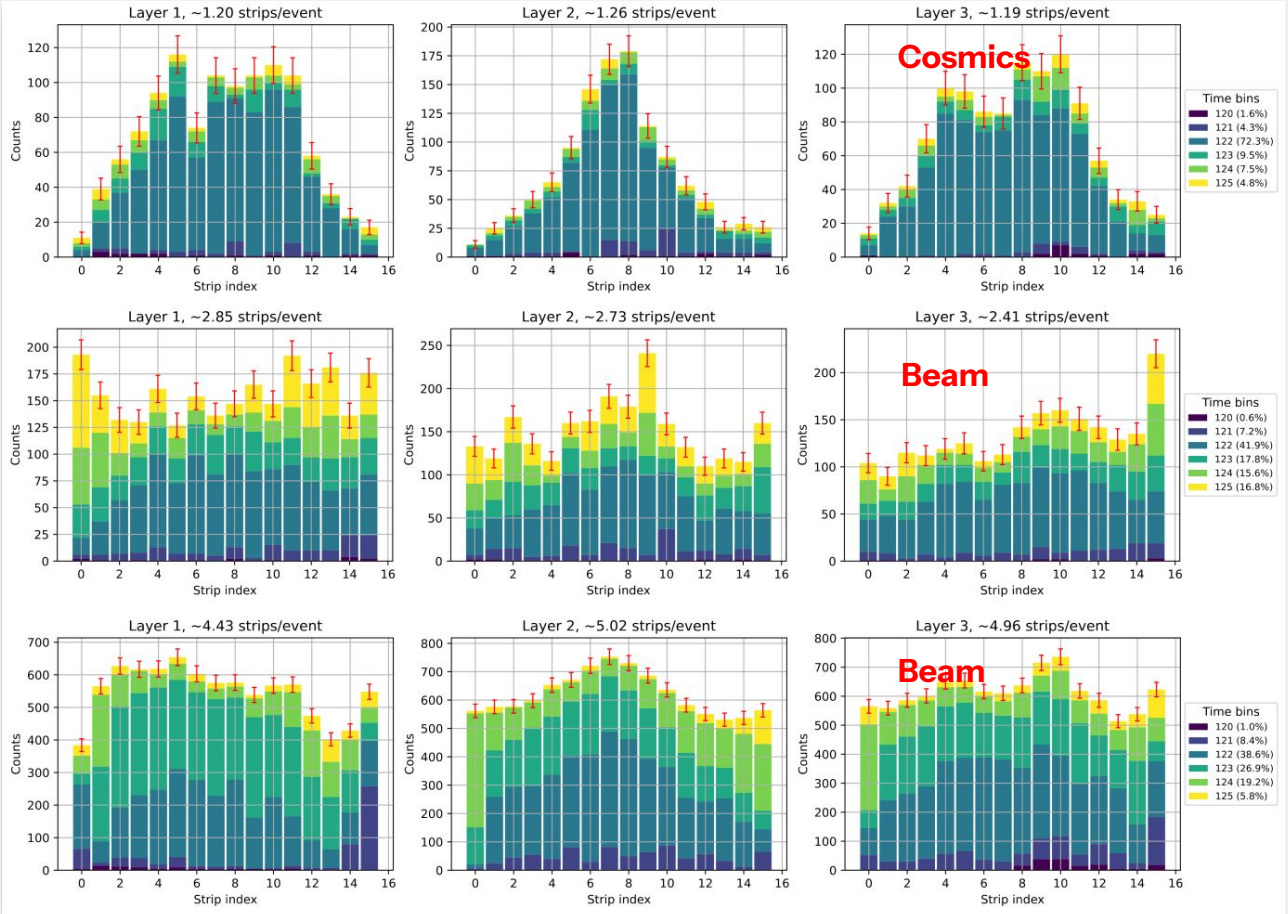
- Mean strip multiplicity  $> 1$  for all chambers
- Matches expectations for cosmic muons
- Validates selected threshold values [35,40].



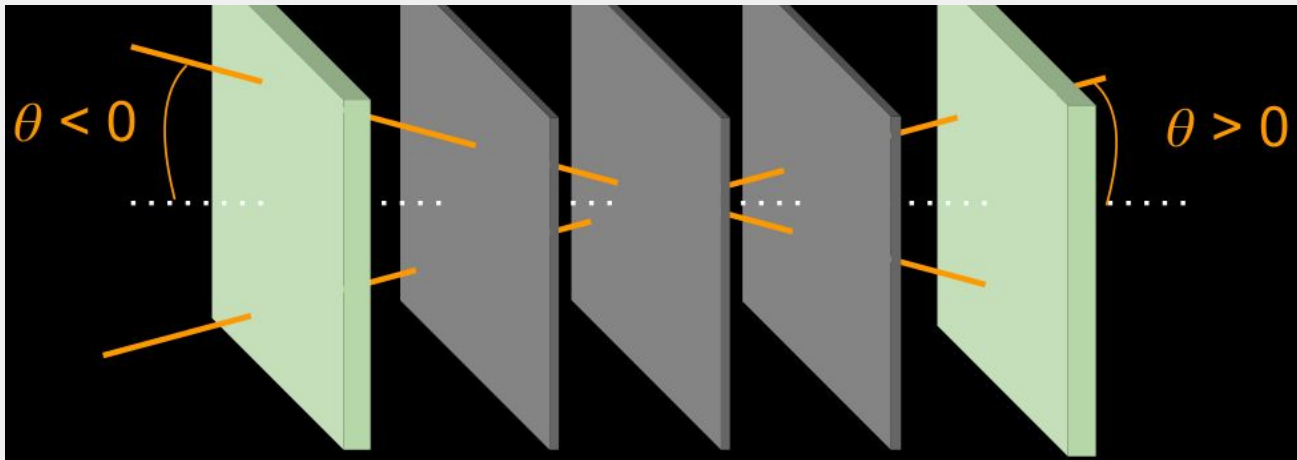
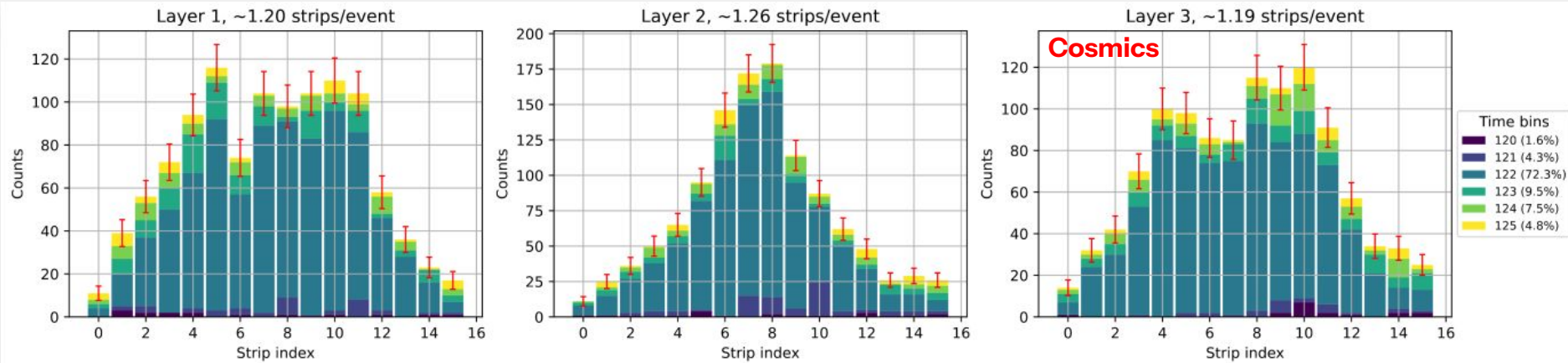
# Occupancy D1



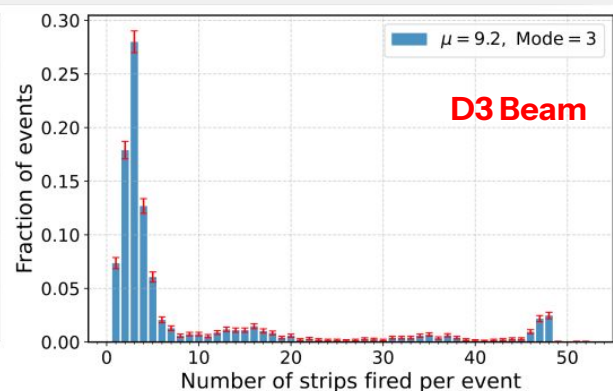
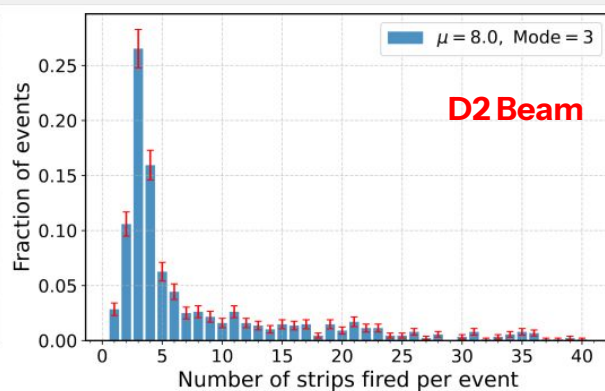
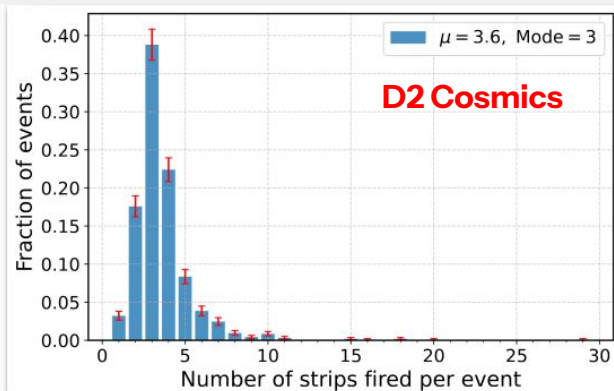
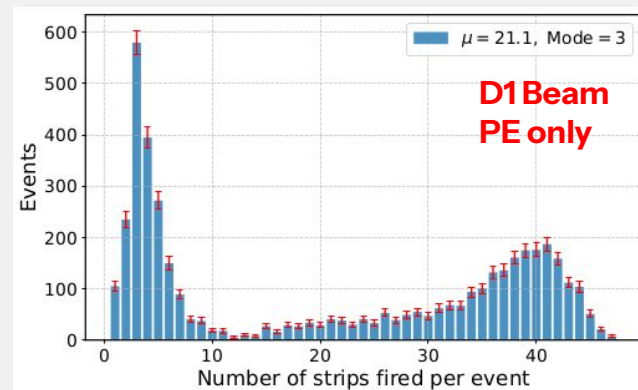
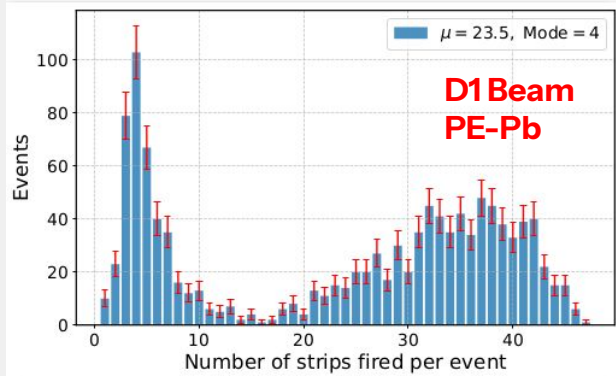
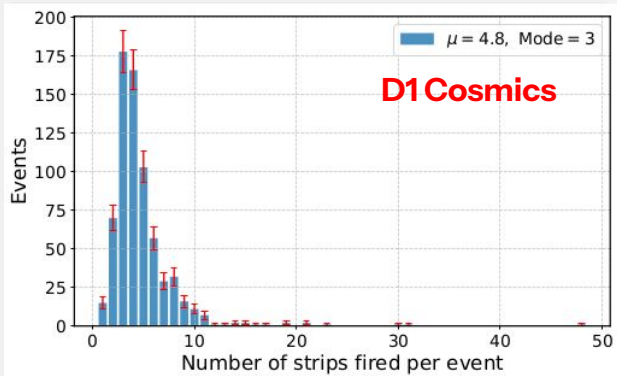
# Occupancy D2 and D3



# Occupancy

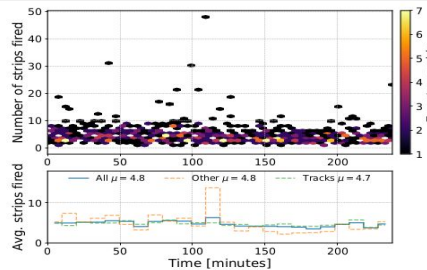


# Strip Multiplicity

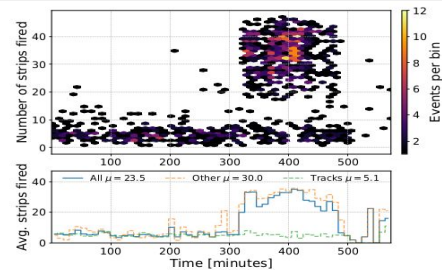


# Strip Multiplicity VS time

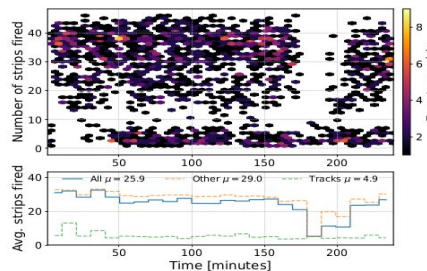
- External scintillator trigger timestamps recorded for all configurations Allows study of temporal structure of detected events
- cosmic run shows stable behavior over time



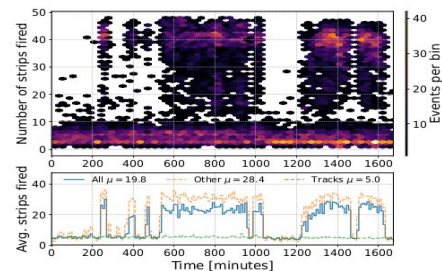
(a) D1 cosmic



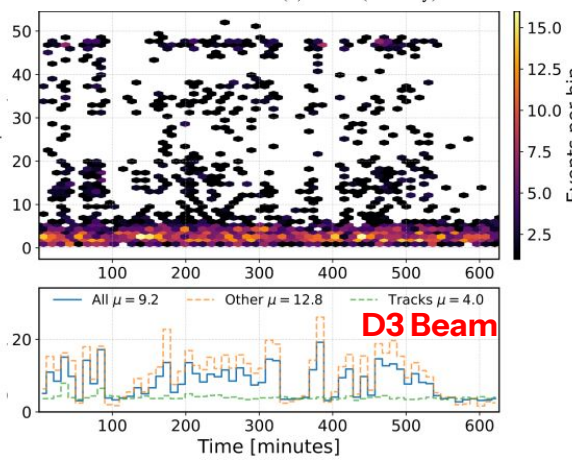
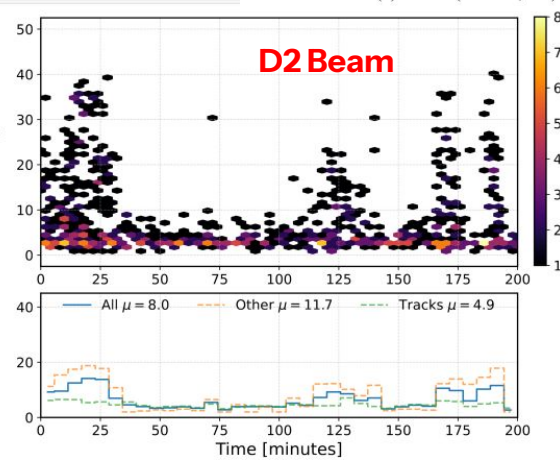
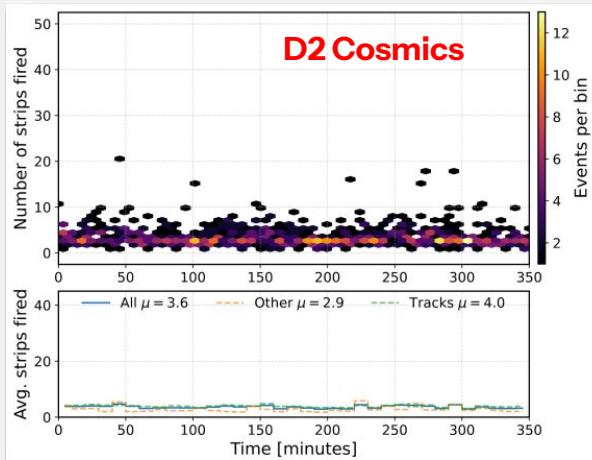
(b) Beam (PE first, Pb)

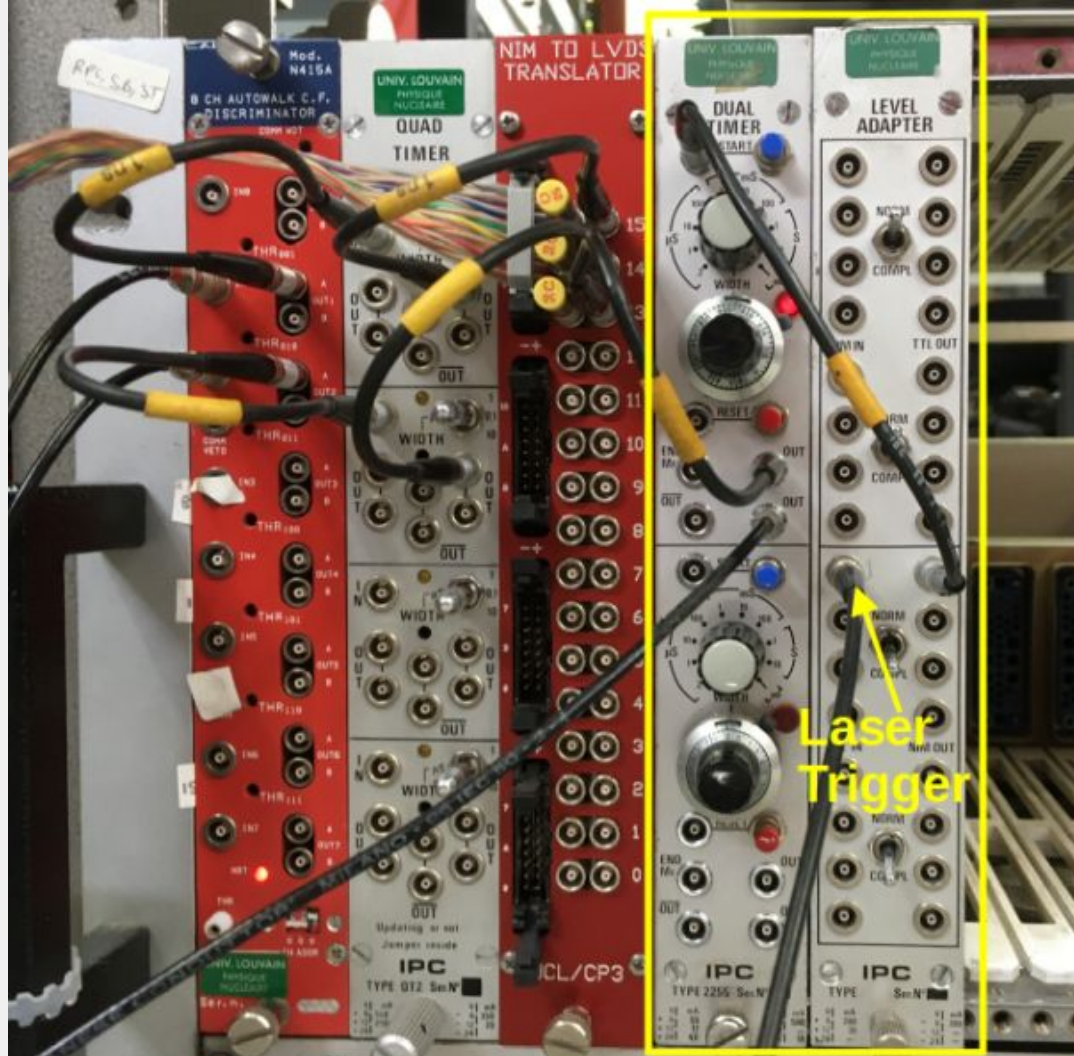


(c) Beam (Pb first, PE)



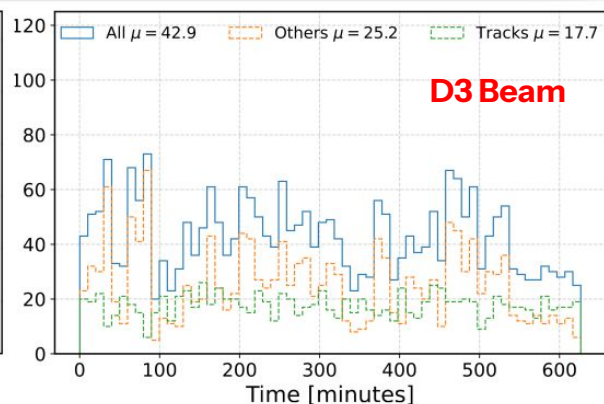
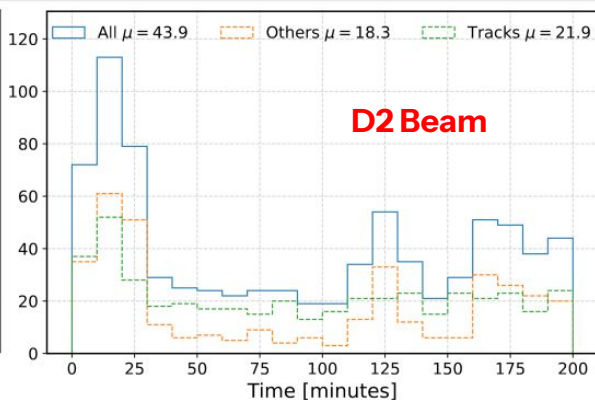
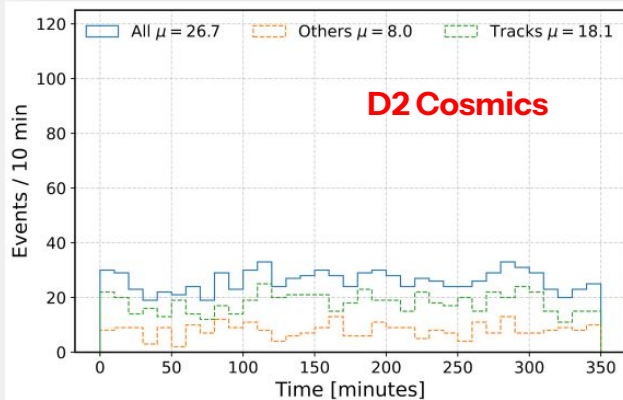
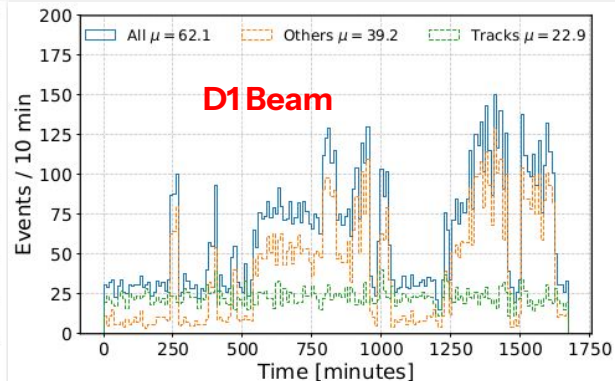
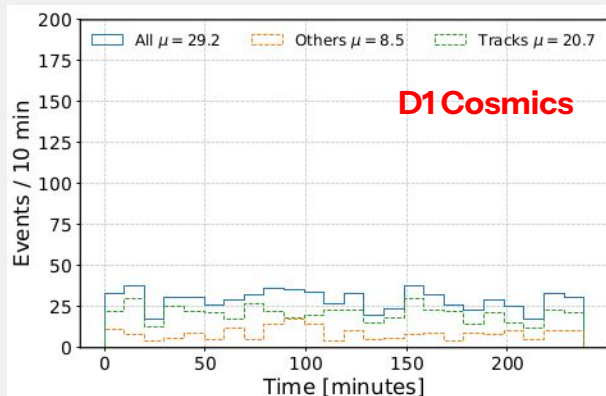
(d) Beam (PE only)



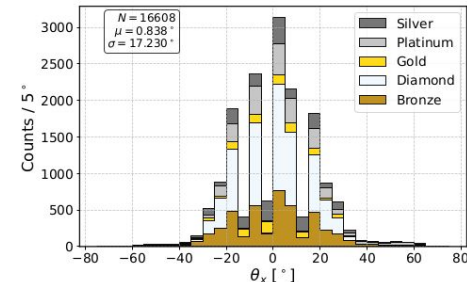
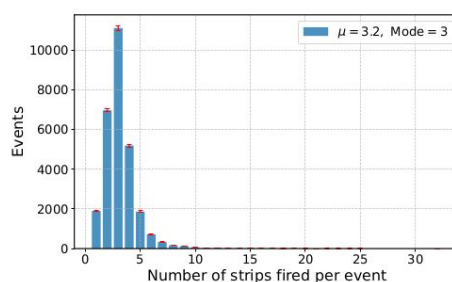
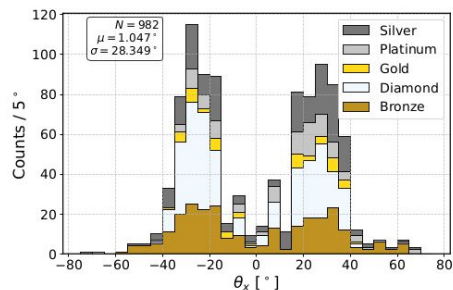
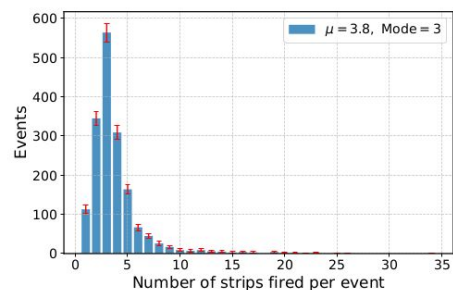


# Event Rate

- Cosmic run shows stable behavior over time
- Track rate  $\approx 18$  events / 10 minutes
- Beam dataset shows strong temporal fluctuations



# Validation with cosmic Data at CP3



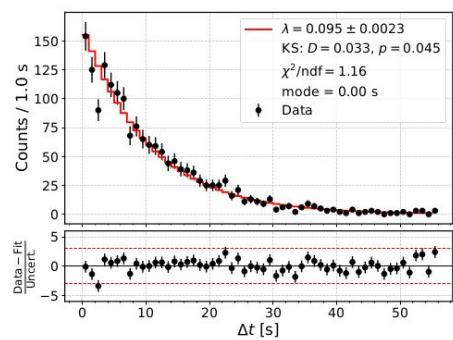
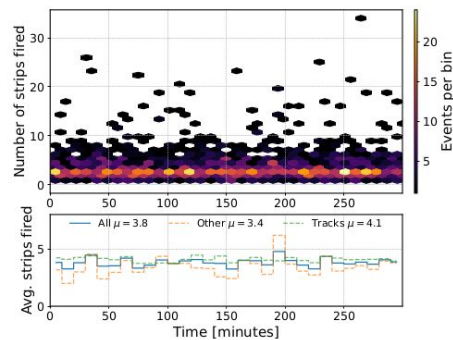
(a) Strip multiplicity for the D1 cosmic-ray dataset.

(b) Zenith-angle distribution for laboratory cosmic rays.

(a) Strip multiplicity for the D1 cosmic-ray dataset.

(b) Zenith-angle distribution for laboratory cosmic rays.

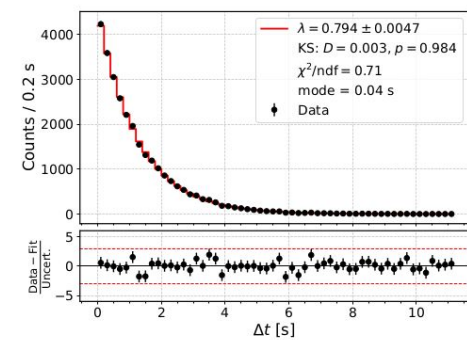
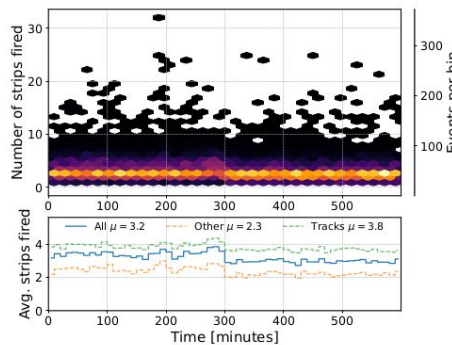
## ELI-like Setup Data



(c) Strip multiplicity as a function of time.

(d) Inter-arrival time distribution.

## Vertical Flux Setup data



(c) Strip multiplicity as a function of time.

(d) Inter-arrival time distribution.

# Principal Component Analysis (PCA)

```
['dt_next',  
 'dt_prev',  
 'hit_10',  
 'hit_11',  
 'hit_12',  
 'n_clusters_10',  
 'n_clusters_11',  
 'n_clusters_12',  
 'n_clusters_total',  
 'n_layers_hit',  
 'n_strips_10',  
 'n_strips_11',  
 'n_strips_12',  
 'n_strips_total',  
 'tb_mean_10',  
 'tb_mean_11',  
 'tb_mean_12',  
 'tb_std_10',  
 'tb_std_11',  
 'tb_std_12',  
 'x_mean_10',  
 'x_mean_11',  
 'x_mean_12',  
 'x_std_10',  
 'x_std_11',  
 'x_std_12']
```

**26 dimensions**

**2 dimensions**

[PCA1, PCA2]