# Hyperonic Three-Body Forces \& Consequences for Neutron Stars 

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## In this talk I will ...

* Present NN $\Lambda$ and NN $\Sigma$ forces based on a two-meson exchange model
* Analyze the role of these forces in the solution of the hyperon puzzle

This study was part of the Ph.D. Thesis of Domenico Logoteta (University of Coimbra, September 2013)


## Hyperons in Neutron Stars

Hyperons in NS considered by many authors since the pioneering work of Ambartsumyan \& Saakyan (1960)

## Phenomenological approaches

$\diamond$ Relativistic Mean Field Models: Glendenning 1985; Knorren et al. 1995; Shaffner-Bielich \& Mishustin 1996, Bonano \& Sedrakian 2012, ...
« Non-realtivistic potential model: Balberg \& Gal 1997
$\diamond$ Quark-meson coupling model: Pal et al. 1999, ...
$\diamond$ Chiral Effective Lagrangians: Hanauske et al., 2000
$\triangleleft$ Density dependent hadron field models: Hofmann, Keil \& Lenske 2001

## Microscopic approaches

$\diamond$ Brueckner-Hartree-Fock theory: Baldo et al. 2000; I. V. et al. 2000, Schulze et al. 2006, I.V. et al. 2011, Burgio et al. 2011, Schulze \& Rijken 2011
¿ DBHF: Sammarruca (2009), Katayama \& Saito (2014)
$\triangleleft \mathrm{V}_{\text {low k }}$ : Djapo, Schaefer \& Wambach, 2010


Hyperons are expected to appear in the core of neutron stars at $\rho \sim$ $(2-3) \rho_{0}$ when $\mu_{\mathrm{N}}$ is large enough to make the conversion of N into Y energetically favorable.

$$
\begin{aligned}
& n+n \rightarrow n+\Lambda \\
& p+e^{-} \rightarrow \Lambda+v_{e^{-}} \\
& n+n \rightarrow p+\Sigma^{-} \\
& n+e^{-} \rightarrow \Sigma^{-}+v_{e^{-}}
\end{aligned}
$$

$$
\begin{gathered}
\mu_{\Sigma^{-}}=\mu_{n}+\mu_{e^{-}}-\mu_{v_{e^{-}}} \\
\mu_{\Lambda}=\mu_{n}
\end{gathered}
$$




## Effect of Hyperons in the EoS and Mass of Neutron Stars




Relieve of Fermi pressure due to the appearance of hyperons $\rightarrow$
EoS softer $\rightarrow$ reduction of the mass

## Measured Neutron Star Masses (up to ~ 2006-2008)


(Lattimer \& Prakash 2007)


## up to $\sim 2006-2008$ any valid EoS should predict

$$
M_{\max }[E o S]>1.4-1.5 M_{\odot}
$$

## Hyperons in NS

(up to $\sim 2006-2008$ )

(Lattimer \& Prakash 2007)

## Phenomenological:

## $\mathrm{M}_{\max }$ compatible with 1.4-1.5 $\mathrm{M}_{\odot}$



Microscopic : $\mathrm{M}_{\max }<1.4-1.5 \mathrm{M}_{\odot}$

(Schulze, Polls, Ramos \& IV 2006)

Recent measurements of high masses $\longrightarrow$ life of hyperons more difficult

- PSR J164-2230 (Demorest et al. 2010)

Shapiro delay:
$\checkmark$ binary sytem ( $\mathrm{P}=8.68 \mathrm{~d}$ )
$\checkmark$ eccentricity $\left(\mathrm{e}=1.3 \times 10^{-6}\right)$
$\checkmark$ companion mass: $\sim 0.5 M_{\odot}$
$\checkmark$ pulsar mass: $\quad M=1.97 \pm 0.04 M_{\odot}$


- PSR J0348+0432 (Antoniadis et al. 2013)
$\checkmark$ binary system $(\mathrm{P}=2.46 \mathrm{~h})$
$\checkmark$ very low eccentricity
$\checkmark$ companion mass: $0.172 \pm 0.003 M_{\odot}$
$\checkmark$ pulsar mass: $\quad M=2.01 \pm 0.04 M_{\odot}$


## Measured Neutron Star Masses (2015)



Observation of $\sim 2 \mathrm{M}_{\text {sun }}$ neutron stars


Dense matter EoS stiff enough is required such that

$$
M_{\max }[E O S]>2 M_{\odot}
$$

Can hyperons still be present in the interior of neutron stars in view of this constraint?

## The Hyperon Puzzle

"Hyperons $\rightarrow$ "soft (or too soft) EoS" not compatible (mainly in microscopic approaches) with measured (high) masses. However, the presence of hyperons in the NS interior seems to be unavoidable."
$\checkmark$ can YN \& YY interactions still solve it ?
$\checkmark$ or perhaps hyperonic three-body forces?
$\checkmark$ what about quark matter?

## Can Hyperonic TBF solve this puzzle?

Natural solution based on: Importance of NNN force in Nuclear Physics (Considered by several authors: Chalk, Gal, Usmani, Bodmer, Takatsuka, Loiseau, Nogami, Bahaduri, IV)


NNY, NYY \& YYY Forces


Can hyperonic TBF provide enough repulsion at high densities to reach $2 \mathrm{M}_{\odot}$ ?

## Two-meson exchange Hyperonic TBF



Vertices: consistent with YN and YY

Repulsion at high densities due to Z-diagram as in NNN

## Baryon-excitation contribution

 ( $\pi$-, $K$-exchange)$$
\begin{aligned}
V_{N N Y}^{M_{1} M_{2}, B}=C_{N N Y}^{M, M_{2}, B} & \left(\hat{O}_{I}\left\{X_{12}\left(\vec{r}_{12}\right), X_{23}\left(\vec{r}_{23}\right)\right\}\right. \\
& \left.+\hat{O}_{I I}\left[X_{12}\left(\vec{r}_{12}\right), X_{23}\left(\vec{r}_{23}\right)\right]\right)
\end{aligned}
$$

$$
\hat{O}_{I}, \hat{O}_{I I} \rightarrow \text { isospin structure }
$$

$$
\begin{aligned}
& X_{i j}(\vec{x})=\vec{\sigma}_{i} \cdot \vec{\sigma}_{j} Y_{i j}(x)+\hat{S}_{i j}(\hat{x}) T_{i j}(x) \\
& Y_{i j}(x)=\frac{\partial^{2} Z_{i j}}{\partial x^{2}}+\frac{2}{x} \frac{\partial Z_{i j}}{\partial x}, T_{i j}(x)=\frac{\partial^{2} Z_{i j}}{\partial x^{2}}-\frac{1}{x} \frac{\partial Z_{i j}}{\partial x} \\
& Z_{12}(x)=\frac{4 \pi}{m_{M_{1}}} \int \frac{d \vec{k}}{(2 \pi)^{3}} e^{-i \vec{k} \cdot \vec{x}} k^{2} F_{M_{1}, B_{1} M_{1}}\left(k^{2}\right) F_{B_{2} B M_{1}}\left(k^{2}\right) \\
& Z_{23}(x)=\frac{4 \pi}{m_{M_{2}}} \int \frac{d \vec{q}}{(2 \pi)^{3}} \frac{e^{-i \vec{k} \cdot x}}{q^{2}+m_{M_{2}}^{2}} F_{B_{3}, B_{3} M_{2}}\left(q^{2}\right) F_{B_{2} B M_{2}}\left(q^{2}\right)
\end{aligned}
$$

## Isospin structure: operators $\hat{O}_{\mathrm{I}} \& \hat{O}_{\text {II }}$

$$
V_{N N Y}^{M_{1} M_{2}, B}=C_{N N Y}^{M_{1} M_{2}, B}\left(\hat{O}_{I}\left\{X_{12}\left(\vec{r}_{12}\right), X_{23}\left(\vec{r}_{23}\right)\right\}+\hat{O}_{I I}\left[X_{12}\left(\vec{r}_{12}\right), X_{23}\left(\vec{r}_{23}\right)\right]\right)
$$

| $V_{N N Y}^{M_{1} M_{2}, B}$ | $\hat{O}_{I}$ | $\hat{O}_{I}$ |
| :---: | :---: | :---: |
| $V_{N N \Lambda}^{\pi \pi, \Sigma^{*}}, V_{N N \Lambda}^{\pi \pi, \Sigma}, V_{N N \Sigma}^{\pi \pi, \Sigma^{*}}, V_{N N \Sigma}^{\pi \pi, \Lambda}, V_{N N \Sigma}^{K K, \Lambda}, V_{N N \Sigma \leftrightarrow N N \Lambda}^{\pi \pi, \Sigma^{*}}$ | $\overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{3}$ |  |
| $V_{N N \Sigma}^{\pi \pi, \Delta}$ | $\left\{\overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \vec{I}_{3}\right\}$ | $\frac{1}{4}\left[\overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \vec{I}_{3}\right]$ |
| $V_{N N \Sigma \leftrightarrow N N \Lambda}^{\pi \pi, \Delta}$ | $\left\{\overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \vec{\rho}_{3}\right\}$ | $\frac{1}{4}\left[\overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \vec{\rho}_{3}\right]$ |
| $V_{N N \Lambda}^{K K, \Sigma^{*}}$ | $\left\{\overrightarrow{1}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{1}_{3}\right\}$ | $-\frac{1}{2}\left[\overrightarrow{1}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{1}_{3}\right]$ |
| $V_{N N \Lambda}^{K K, \Sigma}$ | $\left\{\overrightarrow{1}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{1}_{3}\right\}$ | $\left[\overrightarrow{1}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{1}_{3}\right]$ |
| $V_{N N \Lambda}^{K K, \Lambda}$ |  |  |

## Isospin structure: operators $\hat{O}_{\mathrm{I}} \& \hat{O}_{I I}$ (cont')

$$
V_{N N Y}^{M_{1} M_{2}, B}=C_{N N Y}^{M_{1} M_{2}, B}\left(\hat{O}_{I}\left\{X_{12}\left(\vec{r}_{12}\right), X_{23}\left(\vec{r}_{23}\right)\right\}+\hat{O}_{I I}\left[X_{12}\left(\vec{r}_{12}\right), X_{23}\left(\vec{r}_{23}\right)\right]\right)
$$

| $V_{N N Y}^{M_{1} M_{2}, B}$ | $\hat{O}_{I}$ | $\hat{O}_{I I}$ |
| :---: | :---: | :---: |
| $V_{N N \Sigma}^{K K, \Sigma^{*}}$ | $\left\{\overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{\boldsymbol{\tau}}_{3}\right\}$ | $-\frac{1}{2}\left[\overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{\boldsymbol{\tau}}_{3}\right]$ |
| $V_{N N \Sigma}^{K K, \Sigma}$ | $\left\{\overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{\boldsymbol{\tau}}_{3}\right\}$ | $\left[\overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{\boldsymbol{\tau}}_{3}\right]$ |
| $V_{N N \Sigma \leftrightarrow N N \Lambda}^{K K, \Sigma^{*}}$ | $\left\{\vec{\rho}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{\boldsymbol{\tau}}_{3}\right\}$ | $-\frac{1}{2}\left[\vec{\rho}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{\boldsymbol{\tau}}_{3}\right]$ |
| $V_{N N \Sigma \leftrightarrow N N \Lambda}^{K K, \Sigma}$ | $\left\{\vec{\rho}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{\boldsymbol{\tau}}_{3}\right\}$ | $\left[\vec{\rho}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}, \overrightarrow{\boldsymbol{\tau}}_{2} \cdot \overrightarrow{\boldsymbol{\tau}}_{3}\right]$ |
| $V_{N N \Sigma \leftrightarrow N N \Lambda}^{K K, \Lambda}$ | $\vec{\rho}_{1} \cdot \overrightarrow{\boldsymbol{\tau}}_{2}$ |  |

## Z-diagram contribution <br> ( $\sigma, \omega$-exchange)



- $\sigma \sigma$-exchange contribution

$$
\begin{aligned}
V_{N N Y}^{\sigma \sigma, \bar{B}}= & C_{N N Y}^{\sigma \sigma \cdot \bar{B}}\left(-4 Z_{12}\left(r_{12}\right) Z_{23}\left(r_{23}\right) \nabla_{r_{2}^{\prime}}^{2}-4 Z_{12}^{\prime}\left(r_{12}\right) Z_{23}\left(r_{23}\right) \hat{r}_{12} \cdot \nabla_{r_{2}^{\prime}}\right. \\
& -4 Z_{12}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right) \hat{r}_{23} \cdot \nabla_{r_{2}^{\prime}}-\left(Y_{12}\left(r_{12}\right) Z_{23}\left(r_{23}\right)+Z_{12}\left(r_{12}\right) Y_{23}\left(r_{23}\right)\right) \\
& -\hat{r}_{12} \cdot \hat{r}_{23} Z_{12}^{\prime}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right)-2 i\left(Z_{12}^{\prime}\left(r_{12}\right) Z_{23}\left(r_{23}\right) \vec{\sigma}_{2} \cdot \hat{r}_{12} \times \nabla_{r_{2}^{\prime}}\right. \\
& \left.\left.+Z_{12}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right) \vec{\sigma}_{2} \cdot \hat{r}_{23} \times \nabla_{r_{2}^{\prime}}\right)\right) \delta\left(\vec{r}_{1}-\vec{r}_{1}^{\prime}\right) \delta\left(\vec{r}_{2}-\vec{r}_{2}^{\prime}\right) \delta\left(\vec{r}_{3}-\vec{r}_{3}^{\prime}\right)
\end{aligned}
$$

- $\omega \omega$-exchange contribution

$$
\begin{aligned}
V_{N N Y}^{\omega \omega, \bar{B}} & =C_{N N Y}^{\omega \omega(\bar{B}}\left(\left(\left(1+\vec{\sigma}_{1} \cdot \vec{\sigma}_{2}+\vec{\sigma}_{1} \cdot \vec{\sigma}_{2}+\vec{\sigma}_{2} \cdot \vec{\sigma}_{3}\right) \hat{r}_{12} \cdot \hat{r}_{23}-\vec{\sigma}_{1} \cdot \hat{r}_{23} \vec{\sigma}_{2} \cdot \hat{r}_{12}-\vec{\sigma}_{2} \cdot \hat{r}_{23} \vec{\sigma}_{3} \cdot \hat{r}_{12}\right.\right. \\
& \left.-\vec{\sigma}_{1} \cdot \hat{r}_{23} \vec{\sigma}_{3} \cdot \hat{r}_{12}\right) Z_{12}^{\prime}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right)-2 i Z_{12}^{\prime}\left(r_{12}\right) Z_{23}\left(r_{23}\right)\left(\vec{\sigma}_{2}+\vec{\sigma}_{3}\right) \cdot \hat{r}_{12} \times \nabla_{r_{3}} \\
& \left.-2 i Z_{12}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right)\left(\vec{\sigma}_{2}+\vec{\sigma}_{3}\right) \cdot \hat{r}_{23} \times \nabla_{r_{2}^{\prime}}-4 Z_{12}\left(r_{12}\right) Z_{23}\left(r_{23}\right) \nabla_{r_{1}} \cdot \nabla_{r_{3}}\right) \\
& \delta\left(\vec{r}_{1}-\vec{r}_{1}^{\prime}\right) \delta\left(\vec{r}_{2}-\vec{r}_{2}^{\prime}\right) \delta\left(\vec{r}_{3}-\vec{r}_{3}^{\prime}\right)
\end{aligned}
$$

- $\sigma \omega-$ exchange contribution

$$
\begin{aligned}
V_{N N Y}^{\sigma \omega \bar{B}}= & C_{N N Y}^{\sigma \omega, \bar{B}}\left(\left(\left(1+\vec{\sigma}_{2} \cdot \vec{\sigma}_{3}\right) Z_{12}\left(r_{12}\right) Y_{23}\left(r_{23}\right)-2 i Z_{12}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right)\left(\vec{\sigma}_{2}+\vec{\sigma}_{3}\right) \cdot \hat{r}_{23} \times \nabla_{r_{2}^{\prime}}\right.\right. \\
& +2 i Z_{12}^{\prime}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right)\left(\vec{\sigma}_{2}+\vec{\sigma}_{3}\right) \cdot \hat{r}_{12} \times \hat{r}_{23}+2 i Z_{12}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right) \vec{\sigma}_{2} \cdot \hat{r}_{23} \times \nabla_{r_{3}} \\
& +2 Z_{12}^{\prime}\left(r_{12}\right) Z_{23}\left(r_{23}\right) \hat{r}_{12} \cdot \nabla_{r_{3}^{\prime}}+2 Z_{12}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right) \hat{r}_{23} \cdot \nabla_{r_{3}}+4 Z_{12}\left(r_{12}\right) Z_{23}\left(r_{23}\right) \nabla_{r_{2}^{\prime}} \cdot \nabla_{r_{3}^{\prime}} \\
& \left.-\frac{1}{3}\left(\vec{\sigma}_{2} \cdot \vec{\sigma}_{3} Y_{12}\left(r_{12}\right)+\hat{S}_{23}\left(\hat{r}_{23}\right) T_{23}\left(r_{23}\right)\right) Z_{12}\left(r_{12}\right)\right) \\
& +D_{N N Y}^{\sigma \omega, \bar{B}}\left(-Y_{12}\left(r_{12}\right)+Y_{23}\left(r_{23}\right)-4 Z_{12}^{\prime}\left(r_{12}\right) Z_{23}^{\prime}\left(r_{23}\right)-3 Z_{12}^{\prime}\left(r_{12}\right) \nabla_{r_{2}^{\prime}} \cdot \hat{r}_{12}\right) \\
& +i \vec{\sigma}_{2} \cdot\left(2 \nabla_{r_{23}} \times \nabla_{r_{12}}-5 \nabla_{r_{2}} \times \hat{r}_{23}\right) Z_{12}\left(r_{12}\right) Z_{23}\left(r_{23}\right) \\
& \left.+\left(\vec{r}_{12} \leftrightarrow \vec{r}_{23}, \vec{r}_{1} \leftrightarrow \vec{r}_{1}^{\prime}, \vec{r}_{2} \leftrightarrow \vec{r}_{2}^{\prime}, \vec{\sigma}_{1} \leftrightarrow \vec{\sigma}_{3}\right)\right) \delta\left(\vec{r}_{1}-\vec{r}_{1}^{\prime}\right) \delta\left(\vec{r}_{2}-\vec{r}_{2}^{\prime}\right) \delta\left(\vec{r}_{3}-\vec{r}_{3}^{\prime}\right)
\end{aligned}
$$

# But that's only the beginning of the full story there are 

MANY, MANY, MANY more forces \& contributions ....


Domenico in 2010


Domenico in 2013

## BHF approximation of Hyperonic Matter

$\diamond$ Energy per particle

- $\frac{E}{A}(\rho, \beta)=\frac{1}{A} \sum_{B} \sum_{k \leq k_{F_{B}}}\left(\frac{\hbar^{2} k^{2}}{2 m_{B}}+\frac{1}{2} \operatorname{Re}\left[U_{B}(\vec{k})\right]\right)$
$\diamond$ Bethe-Goldstone Equation
- $G(\omega)=V+V \frac{Q}{\omega-E-E^{\prime}+i \eta} G(\omega)$
- $E_{B}(k)=\frac{\hbar^{2} k^{2}}{2 m_{B}}+\operatorname{Re}\left[U_{N}(k)\right]+m_{B}$
- $U_{B}(k)=\sum_{B^{\prime}} \sum_{k^{\prime} \leq k_{f_{B}}}\left\langle\vec{k} \vec{k}^{\prime}\right| G\left(\omega=E_{B}(k)+E_{B^{\prime}}\left(k^{\prime}\right)\right)\left|\vec{k} \vec{k}^{\prime}\right\rangle$


Infinite sumation of two-hole line diagrams

Partial sumation of pp ladder diagrams
pur $=\rangle\langle+)_{Y}^{+}+\cdots+\cdots=$

$\checkmark$ Pauli blocking
$\checkmark$ Baryon dressing

## Three-Body Forces within the BHF approach

TBF can be introduced in our BHF approach by adding effective density-dependent two body forces to the baryon-baryon interactions V when solving the Bethe-Goldstone equation

$\mathrm{W}_{3}\left(\overrightarrow{\mathrm{r}}_{\mathrm{i}}, \vec{r}_{\mathrm{j}}, \vec{r}_{\mathrm{k}}\right)$

$V_{B_{i} B_{j}}^{\text {eff }}\left(\vec{r}_{i j}\right)$

$$
V_{B_{j} B_{j}}\left(\vec{r}_{i j}\right)=\frac{1}{\left(2 S_{B_{k}}+1\right)\left(2 I_{B_{k}}+1\right)} \operatorname{Tr} \int d^{3} \vec{r}_{k} \sum_{c y c} W_{3}\left(\vec{r}_{i}, \vec{r}_{j}, \vec{r}_{k}\right) n\left(\vec{r}_{i}, \vec{r}_{j}, \vec{r}_{k}\right)
$$

$$
W_{3}\left(\vec{r}_{i}, \vec{r}_{j}, \vec{r}_{k}\right): \text { genuine TBF } \quad n\left(\vec{r}_{i}, \vec{r}_{j}, \vec{r}_{k}\right): \text { three-body correlation function }
$$

we take: $\quad n\left(\vec{r}_{i}, \vec{r}_{j}, \vec{r}_{k}\right)=\rho_{B_{k}} g_{B_{i} B_{k}}^{2} g_{B_{j} B_{k}}^{2} \quad$ with $\quad g_{B_{m} B_{n}}$ : two-body correlation function

From the genuine NNN,NNY, NYY and YYY TBF ...

$$
\mathrm{NNN} \rightarrow \mathrm{NN}
$$



NYY $\rightarrow$ NY, YY

$\mathrm{NNY} \rightarrow \mathrm{NN}, \mathrm{NY}$


$$
\mathrm{YYY} \rightarrow \mathrm{YY}
$$




Effective NN density-dependent 2BF from NNY

- $V_{N N}^{\pi \tau Y, B}(\vec{r})=C_{N N Y}^{\pi \pi, B} \rho_{Y}\left[V_{S}^{\pi \pi}(\vec{r}) \vec{\sigma}_{1} \cdot \vec{\sigma}_{2}+V_{T}^{\pi \pi}(\vec{r}) S_{12}(\hat{r})\right] \vec{\tau}_{1} \cdot \vec{\tau}_{2}$
- $V_{N N}^{\sigma \sigma Y, \bar{B}}(\vec{r})=C_{N N Y}^{\sigma \sigma, \bar{B}}\left[\rho_{N} V_{C_{1}}^{\sigma \sigma}(\vec{r})+\rho_{N}^{5 / 3} V_{C_{2}}^{\sigma \sigma}(\vec{r})\right]$
- $V_{N N}^{\omega \omega Y}, \bar{B}(\vec{r})=C_{N N Y}^{\omega \omega, \bar{B}} \rho_{Y}\left[V_{C}^{\omega \omega}(\vec{r})+V_{S}^{\omega \omega}(\vec{r}) \vec{\sigma}_{1} \cdot \vec{\sigma}_{2}+V_{T}^{\omega \omega}(\vec{r}) S_{12}(\hat{r})\right]$
- $V_{N N}^{\sigma \omega Y, B}(\vec{r})=C_{N N Y}^{\sigma \omega, \vec{B}} \rho_{N} V_{C}^{\sigma \omega}(\vec{r})$


## Effective $\mathrm{N} \Lambda$ density-dependent 2BF from $\mathrm{NN} \Lambda$




- $V_{N \Lambda}^{K K N, \Lambda}(\vec{r})=C_{N N \Lambda}^{K K, \Lambda} \rho_{N}\left[V_{S}^{K K}(\vec{r}) \vec{\sigma}_{1} \cdot \vec{\sigma}_{2}+V_{T}^{K K}(\vec{r}) S_{12}(\hat{r})\right]$
- $V_{N \Lambda}^{K K N, \Sigma / \Sigma^{*}}(\vec{r})=C_{N N \Lambda}^{K K, \Sigma / \Sigma^{*}} \rho_{N}\left[V_{S}^{K K}(\vec{r}) \vec{\sigma}_{1} \cdot \vec{\sigma}_{2}+V_{T}^{K K}(\vec{r}) S_{12}(\hat{r})\right] \overrightarrow{\boldsymbol{\tau}}_{1} \cdot \overrightarrow{1}_{2}$
- $V_{N \Lambda}^{\sigma \sigma N, \bar{N}}(\vec{r})=C_{N N \Lambda}^{\sigma \sigma, \bar{N}}\left[\rho_{\Lambda} V_{C_{1}}^{\sigma \sigma}(\vec{r})+\rho_{\Lambda}^{5 / 3} V_{C_{2}}^{\sigma \sigma}(\vec{r})\right]$
- $V_{N \Lambda}^{\omega \omega N, \bar{N}}(\vec{r})=C_{N N \Lambda}^{\omega \omega, \bar{N}} \rho_{N}\left[V_{C}^{\omega \omega}(\vec{r})+V_{S}^{\omega \omega}(\vec{r}) \vec{\sigma}_{1} \cdot \vec{\sigma}_{2}+V_{T}^{\omega \omega}(\vec{r}) S_{12}(\hat{r})\right]$
- $V_{N \Lambda}^{\sigma \omega N, \bar{N}}(\vec{r})=C_{N N \Lambda}^{\sigma \omega, \bar{N}} \rho_{\Lambda} V_{C}^{\sigma \omega}(\vec{r})$



## Effective density-dependent transition $\mathrm{N} \Sigma$ - $\mathrm{N} \Lambda$ from NN $\Sigma$ - NN $\Lambda$






- $V_{N \Sigma \rightarrow N A}^{K K N / N / \Sigma Z^{*}}(\vec{r})=C_{N N \Sigma \rightarrow N N A}^{K K N / V V_{N}} \rho_{N}\left[V_{s}^{K K}(\vec{r}) \vec{\sigma}_{1} \cdot \vec{\sigma}_{2}+V_{T}^{K K}(\vec{r}) S_{12}(\hat{r})\right] \overrightarrow{1}_{1} \cdot \overrightarrow{1}_{2}$


## Effect of TBF on Mean Field \& E/A


$\checkmark$ Only NNY considered (preliminar)
$\checkmark$ Repulsion at high densities due to Z-diagram contribution as in NNN



Work is in progress, many more contributions have to be considered, but we can still try to estimate the effect of hyperonic TBF in NS

1-. Construct the hyperonic matter EoS within the BHF at 2 body level
(Av18 NN + NSC89 YN)

2-. Add simple phenomenological density-dependent contact terms that mimic the effect of TBF.

Density-dependent contact terms: (Balberg \& Gal 1997)
Potential of a baryon $\mathrm{B}_{\mathrm{y}}$ in a sea of baryons $\mathrm{B}_{\mathrm{x}}$ of density $\rho_{\mathrm{x}}$


Folding $\mathrm{V}_{\mathrm{y}}\left(\rho_{\mathrm{x}}\right)$ with $\rho_{\mathrm{x}}, \mathrm{V}_{\mathrm{x}}\left(\rho_{\mathrm{y}}\right)$ with $\rho_{\mathrm{y}}$ and combining with weight factors $\rho_{x} / \rho$ and $\rho_{y} / \rho$

$$
\varepsilon_{x y}\left(\rho_{x}, \rho_{y}\right)=a_{x y} \rho_{x} \rho_{y}+b_{x y} \rho_{x} \rho_{y}\left(\frac{\rho_{x}^{\gamma_{x y}}+\rho_{y}^{\gamma_{x y}}}{\rho_{x}+\rho_{y}}\right)
$$

attraction
repulsion
larger than 1

## Effect of hyperonic TBF on $\mathrm{M}_{\text {max }}$



IV, Logoteta, et al., (2011)

| $\gamma_{N N}$ | $x$ | $\gamma_{Y N}$ | Maximum Mass |
| :---: | :---: | :---: | :---: |
|  | 0 | - | $1.27(2.22)$ |
|  | $1 / 3$ | 1.49 | 1.33 |
| 2 | $2 / 3$ | 1.69 | 1.38 |
|  | 1 | 1.77 | 1.41 |
|  | 0 | - | $1.29(2.46)$ |
|  | $1 / 3$ | 1.84 | 1.38 |
| 2.5 | $2 / 3$ | 2.08 | 1.44 |
|  | 1 | 2.19 | 1.48 |
|  | 0 | - | $1.34(2.72)$ |
|  | $1 / 3$ | 2.23 | 1.45 |
| 3 | $2 / 3$ | 2.49 | 1.50 |
|  | 1 | 2.62 | 1.54 |
|  | 0 | - | $1.38 \widetilde{2.97}$ |
|  | $1 / 3$ | 2.63 | 1.51 |
| 3.5 | $2 / 3$ | 2.91 | 1.56 |
|  | 1 | 3.05 | 1.60 |

Hyperonic TBFs seem not to be the full solution of the "Hyperon Puzzle", although they probably contribute to its solution

$$
1.27<M_{\max }<1.6 M_{\odot}
$$

## A comment must be done at this point

Yamamoto et al. (2015)

## Lonardoni et al. (2015)

BHF with NN+YN+universal repulsive TBF (multipomeron exchange mecanism)


$$
M_{\max }>2 M_{\odot}
$$

First Quantum Monte Carlo calculation on neutron $+\Lambda$ matter


Some of the parametrizations of the $\Lambda n n$ force give maximum masses compatible with $2 \mathrm{M}_{\odot}$ but the onset of $\Lambda$ is above the largest density considered in the calculation $\sim 0.56 \mathrm{fm}^{-3}$

## Summary \& Conclusions

* Construction of two-meson exchange hyperonic TBF

Repulsion is obtained at high densities (Z-diagram)
D. Logoteta, Ph.D. Thesis (Univ. Coimbra, Sept. 2013)

* Simple model to establish numerical lower and upper limits to the effect of hyperonicTBF on the maximum mass of NS.

Assuming the strength of hyperonic $\mathrm{TBF} \leq$ nucleonic TBF:
$1.27 \mathrm{M}_{\odot}<\mathrm{M}_{\max }<1.60 \mathrm{M}_{\odot} \quad$ compatible with 1.4-1.5 $\mathrm{M}_{\odot}$
but incompatible with observation of very massive NS

$$
\begin{aligned}
& \text { PSR J1903+0327 }(1.67 \pm 0.01) \mathrm{M}_{\odot} \\
& \text { PSR J1614-2230 }(1.97 \pm 0.04) \mathrm{M}_{\odot} \\
& \text { PSR J0348+0432 }(2.01 \pm 0.04) \mathrm{M}_{\odot}
\end{aligned}
$$

* There is not yet a general agreement between different approaches/models


## Take away message

Hyperonic Three-Body Forces seem not to be the full solution to the "Hyperon Puzzle", although they probably can contribute to it

- You for your time \& attention
- The sponsors for their support


