# Nuclear astrophysics research in ATOMKI

György Gyürky Institute of Nuclear Research (ATOMKI) Debrecen, Hungary

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## Research activities of the nuclear astrophysics group of



### The LUNA collaboration

- Italian-German-British-Hungarian collaboration
- Operates the only underground accelerator of the world at LNGS, Gran Sasso, Italy
- Measurement of extremely low cross sections of astrophysical reactions





#### **Reactions studied at LUNA**



## The importance of ${}^{3}\text{He}(\alpha,\gamma){}^{7}\text{Be}$

## pp-chain of Hydrogen burning, solar neutrinos

Υ SuperK, SNO 1012 1011 Bahcall-Serenelli 2005 pp→±1% Neutrino Spectrum  $(\pm 1\sigma)$ 1010 D/H  $\pm 10.5\%$ ²Be→ 10 9 <sup>7</sup>lux (cm<sup>-2</sup> s<sup>-1</sup>) 10 8 <sup>3</sup>He/H peþ 10 +16%10 <sup>7</sup>Be  $\pm 10.5\%$ 10 5 10 4 7Li/H hep→ ±16% 10<sup>3</sup> 10 <sup>2</sup> 10 <sup>1</sup> – 0.1 10 Neutrino Energy in MeV

big-bang nucleosynthesis, <sup>7</sup>Li problem



#### Old data of <sup>3</sup>He( $\alpha,\gamma$ )<sup>7</sup>Be S-**High uncertainty Ambiguous results** 0.700 Nagatani 1969 Parker 1963 Krawinkel 1982 0.600 Robertson 1983 Osborn 1984 S factor [keV barn] Osborn act. 1984 Hilgemeier 1988 0.500 Nara Singh 2004 on-line activation 0.400 0.300 0.200 500 1000 1500 2000 0 E<sub>c.m.</sub> [keV] pp-chain big bang 6

## The LUNA of ${}^{3}\text{He}(\alpha,\gamma){}^{7}\text{Be}$ measurements

## Low energies never reached before = 92 – 170 keV)

- With both activation and in-beam methods (good agreement found)
- ✓ With high precision (<5%)</p>

#### The LUNA results



D. Bemmerer *et al.*, Phys. Rev. Lett. **97** (2006) 122502
Gy. Gyürky *et al.*, Phys. Rev. C **75** (2007) 035805
F. Confortola *et al.*, Phys. Rev. C **75** (2007) 065803

#### ERNA <sup>3</sup>He( $\alpha$ , $\gamma$ )<sup>7</sup>Be experiment



#### **ERNA Results**



A. di Leva et al., Phys. Rev. Lett. **102** (2009) 232502

### <sup>3</sup>He( $\alpha$ , $\gamma$ )<sup>7</sup>Be: outlook

- new measurements in ATOMKI
- high energy activation and indirect measurements (ANC)





#### The p-process nucleosynthesis

- Secondary process initiated on s- and r-seeds
- Several sub-processes involved
- Most important:  $\gamma$ -process ( $\gamma$ -induced reactions)
- Models are not able to reproduce the observed p-isotope abundances



### Data needs for the p-process

- Astrophysics: stellar environment
  - temperature
  - density
  - seed abundances
  - time scale
  - □ etc...

- Nuclear physics
  - masses
  - decay properties
  - reaction rates!!!

#### **Reaction rates:**

- based on theory
- very few experimental data
- high uncertainty
- experiments highly needed

#### Experiments

- Alpha-induced reations: 8-15 MeV
  - $\Rightarrow$  Cyclotron

- Proton-induced reactions: 1-4 MeV
  - $\Rightarrow$  Van de Graaff





## Capture reaction cross section measurements

#### Comparison with theory



Z. Halász et al., in prep.



### Fine tuning of parameters



G.G. Kiss et al., Phys. Rev. Lett. 101 (2008) 191101



#### Direct determination of alphanucleus optical potential

- High precision elastic scattering experiments
- Low energies (around Coulomb-barrier)
- Comparison with global optical potentials
- Construction of local potentials
- Experiments: cyclotron of ATOMKI



# Capture and scattering experiments





#### P-process: Outlook

further needs for experimental data
 recognized by the scientific community



**European Research Council** 

-Supporting top researchersfrom anywhere in the world



#### For further information...



#### http://www.atomki.hu/atomki/IonBeam/nag/index\_en.html

Magyar verzió 💳

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Collaborations

ERC grant

#### Welcome to the Website of the

#### **Nuclear Astrophysics Group**

Welcome to the homepage of the Nuclear Astrophysics Group of ATOMKI. Our group has been founded by Prof. Endre Somorjai and became one of the two research groups of the Section of Ion Beam Physics. Since its foundation, NAG represents experimental nuclear astrophysics among the walls of ATOMKI. In these pages you will find detailed information about our past and present activities, collaborations and group members. Have a nice browsing!