

# Opening a New View to the Structure of Nuclei



*Isao Tanihata*

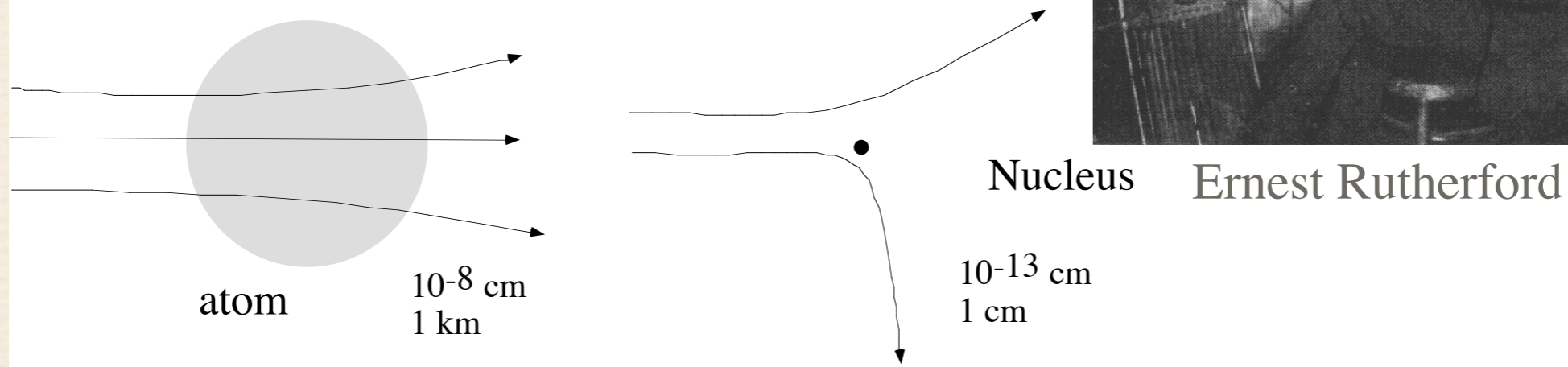
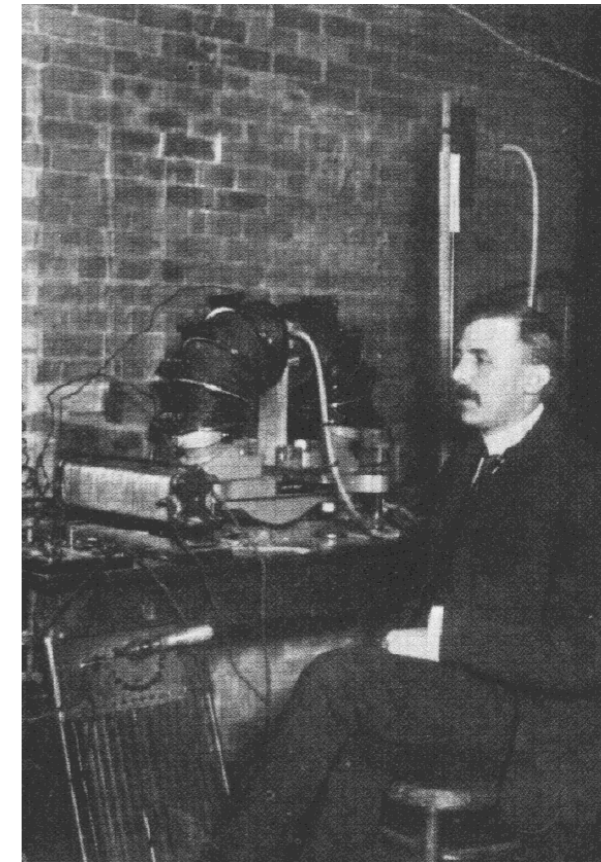
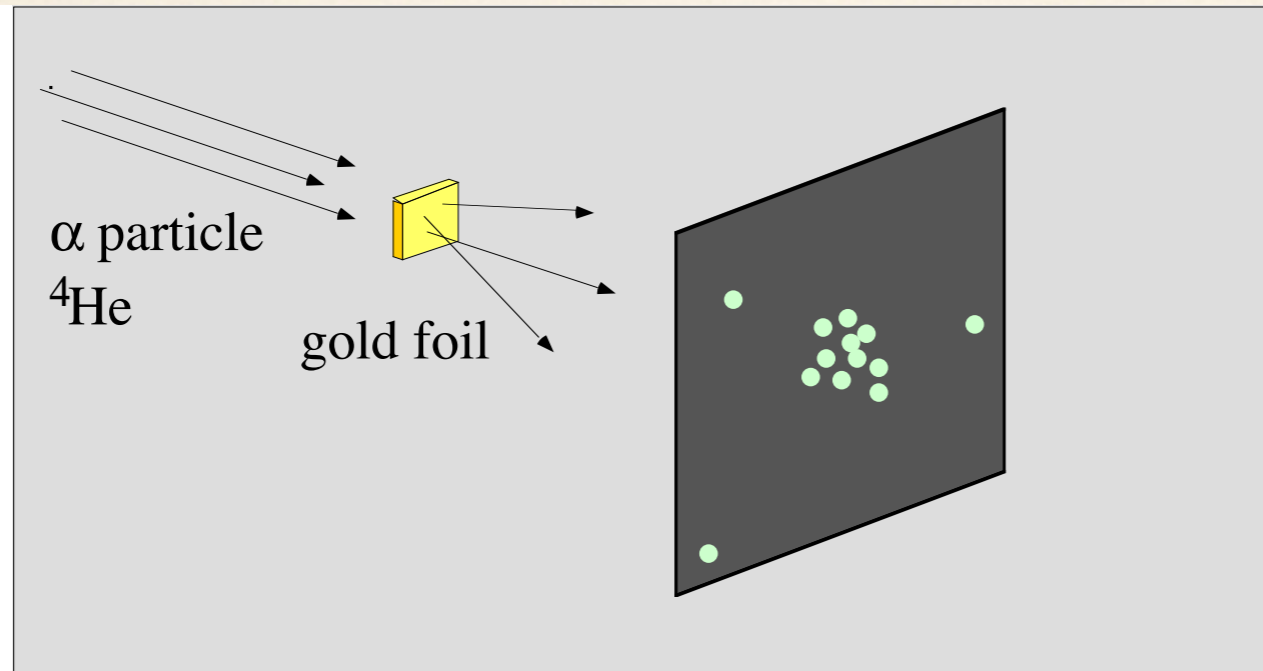
*IRCNPC, Beihang University, Beijing, China*

*RCNP, Osaka University*

Inaugural talk at Academia Europe, 2017 Budapest

# The first observation of nuclei and its size

in ~1910

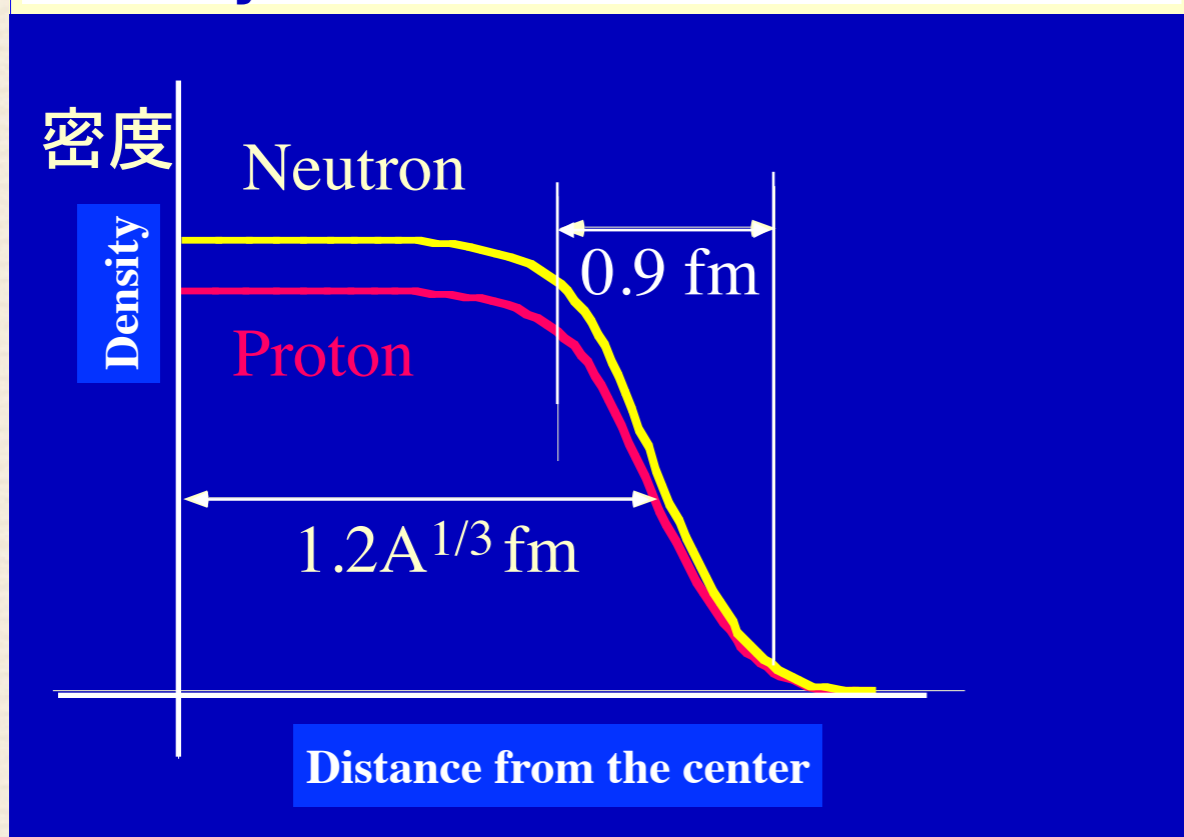


❖ **Observation of nuclei are mostly by scattering and decay!**

# “Common properties” obtained from e, p, .. scattering

- ❖ Nucleus is like a liquid drop with diffused surface

## Density distribution of stable nuclei

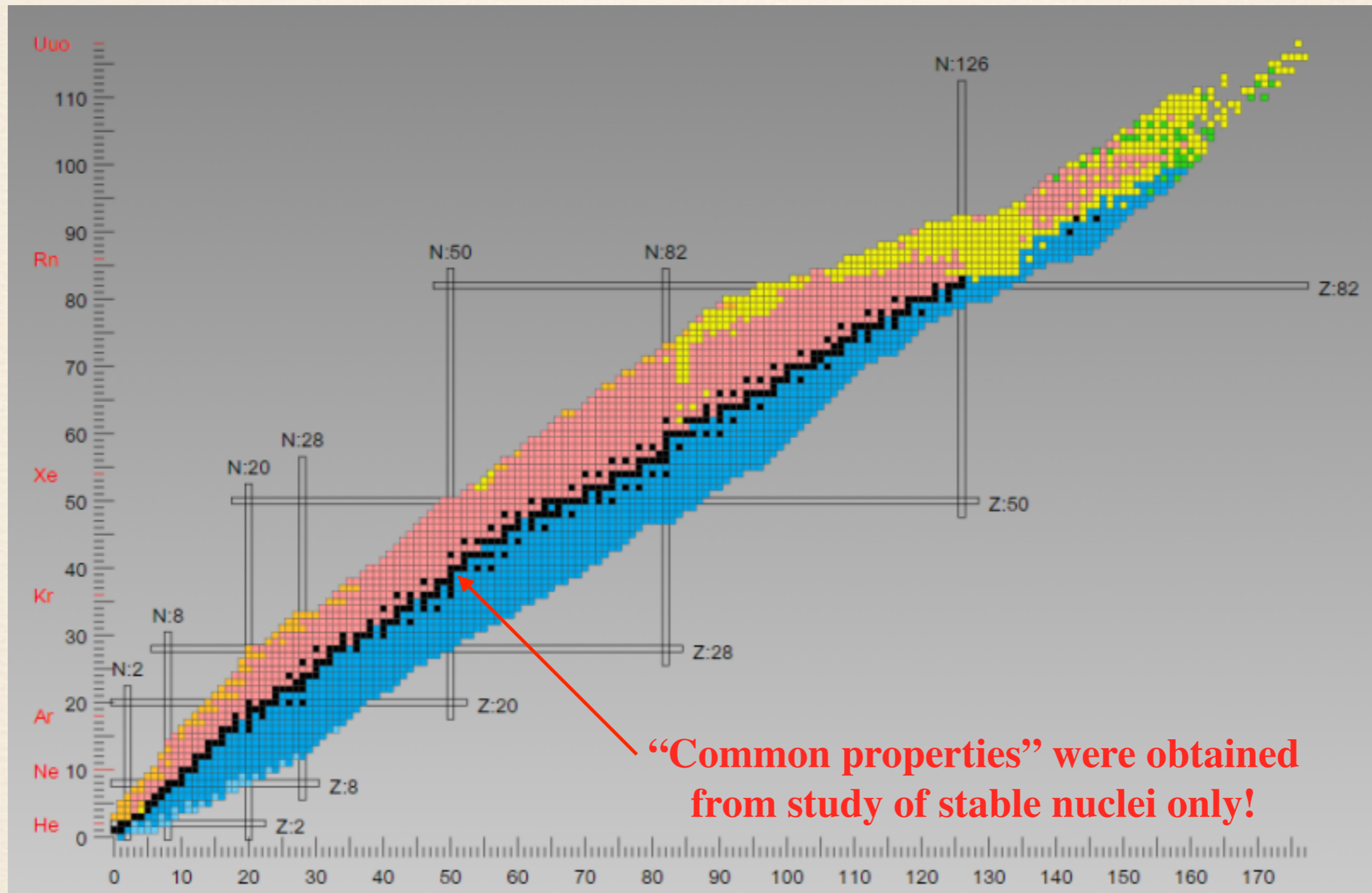


1.  $R = r_0 A^{1/3}$
2. Surface diffuseness is constant
3.  $\rho_p(r) \propto \rho_n(r)$

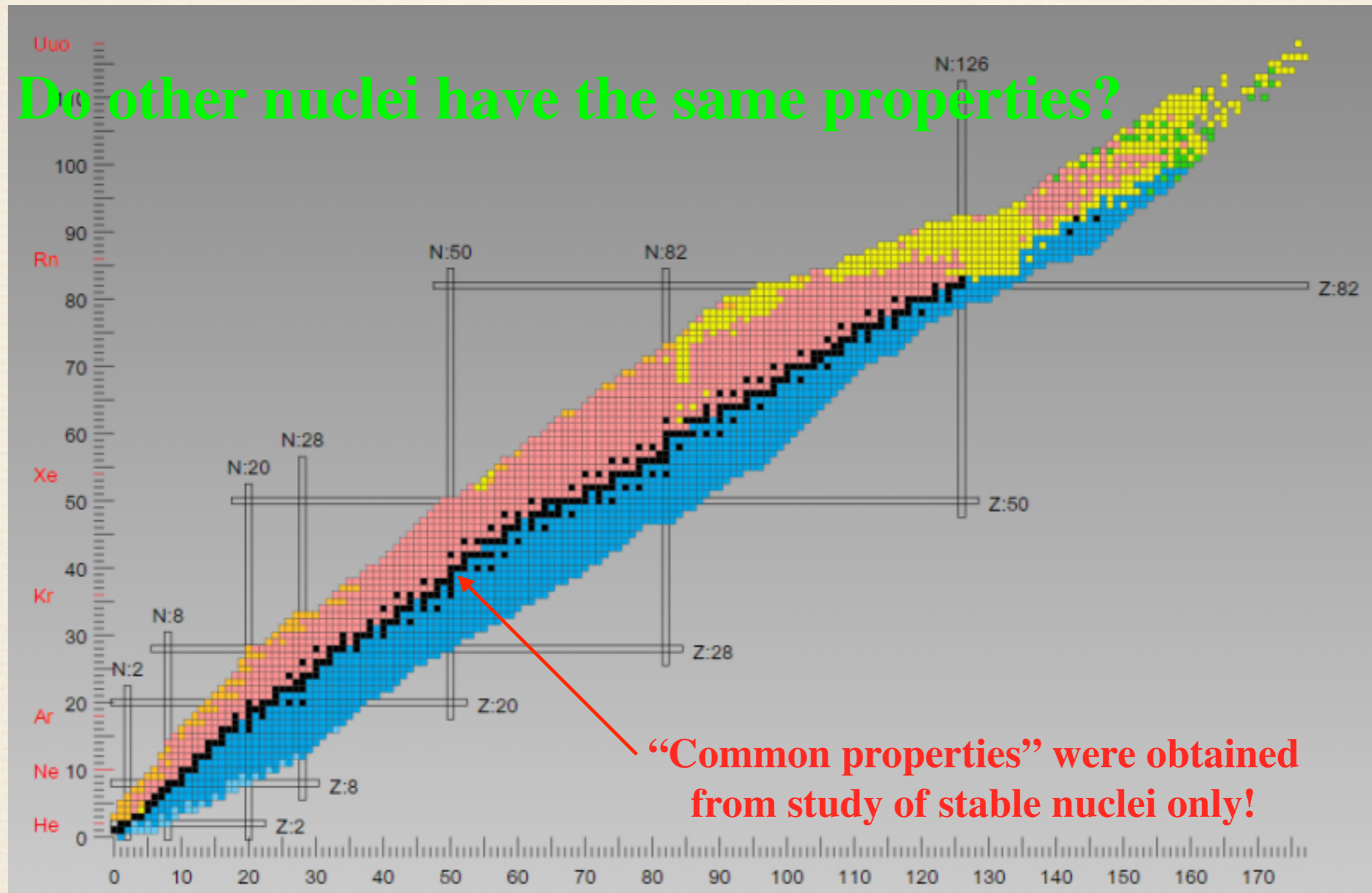
They hold both for  
 $^{40}\text{Ca}(Z=20, N=20)$   
 $^{208}\text{Pb}(Z=82, N=126)$

- ❖ 1. saturation of nuclear density,
- ❖ 2. range of nuclear forces,
- ❖ 3. strong proton-neutron attraction.

# Chart of nuclei ( $\sim 3000$ nuclei observed)



# Chart of nuclei ( $\sim 3000$ nuclei observed)



# Two essential discoveries in 1970's

---

- ❖ **Discovery of projectile fragmentation at Berkeley Bevalac**
  - ❖ *H. H. Heckman, D. G. Greiner et al., 1972*
- ❖ **Efficient production of short lived nuclei by projectile fragmentations**
  - ❖ *G. Westfall, T.J.M Symons et al., 1979*

# Projectile fragmentation ( $^{16}\text{O} + \text{CH}_2 \rightarrow {}^A\text{C} + \dots$ )

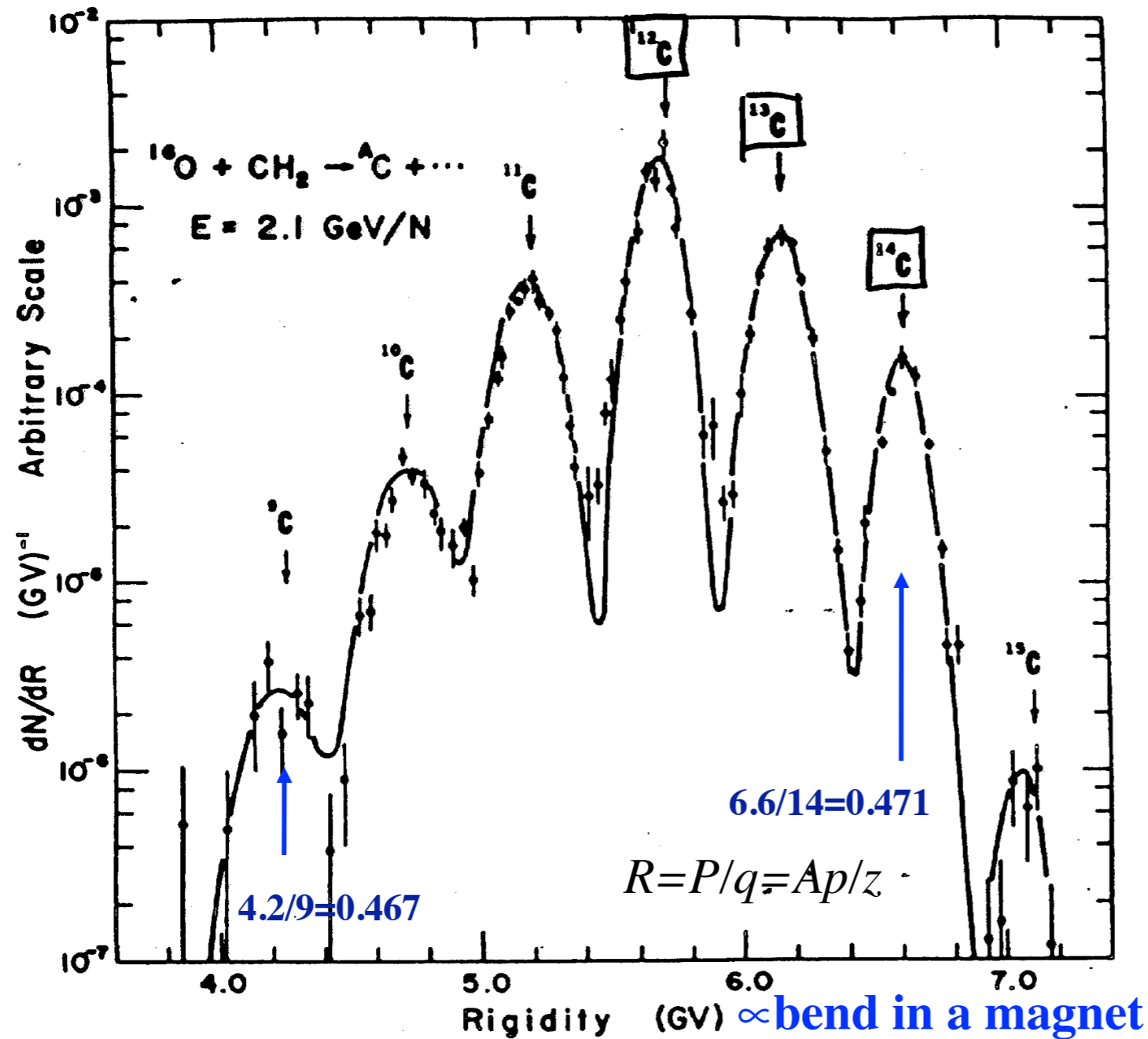


Figure 5 Rigidity spectrum of carbon isotopes produced by the fragmentation of  $^{16}\text{O}$  projectiles at 2.1 A GeV. Arrows indicate the rigidities for each isotope evaluated at beam velocity (87).

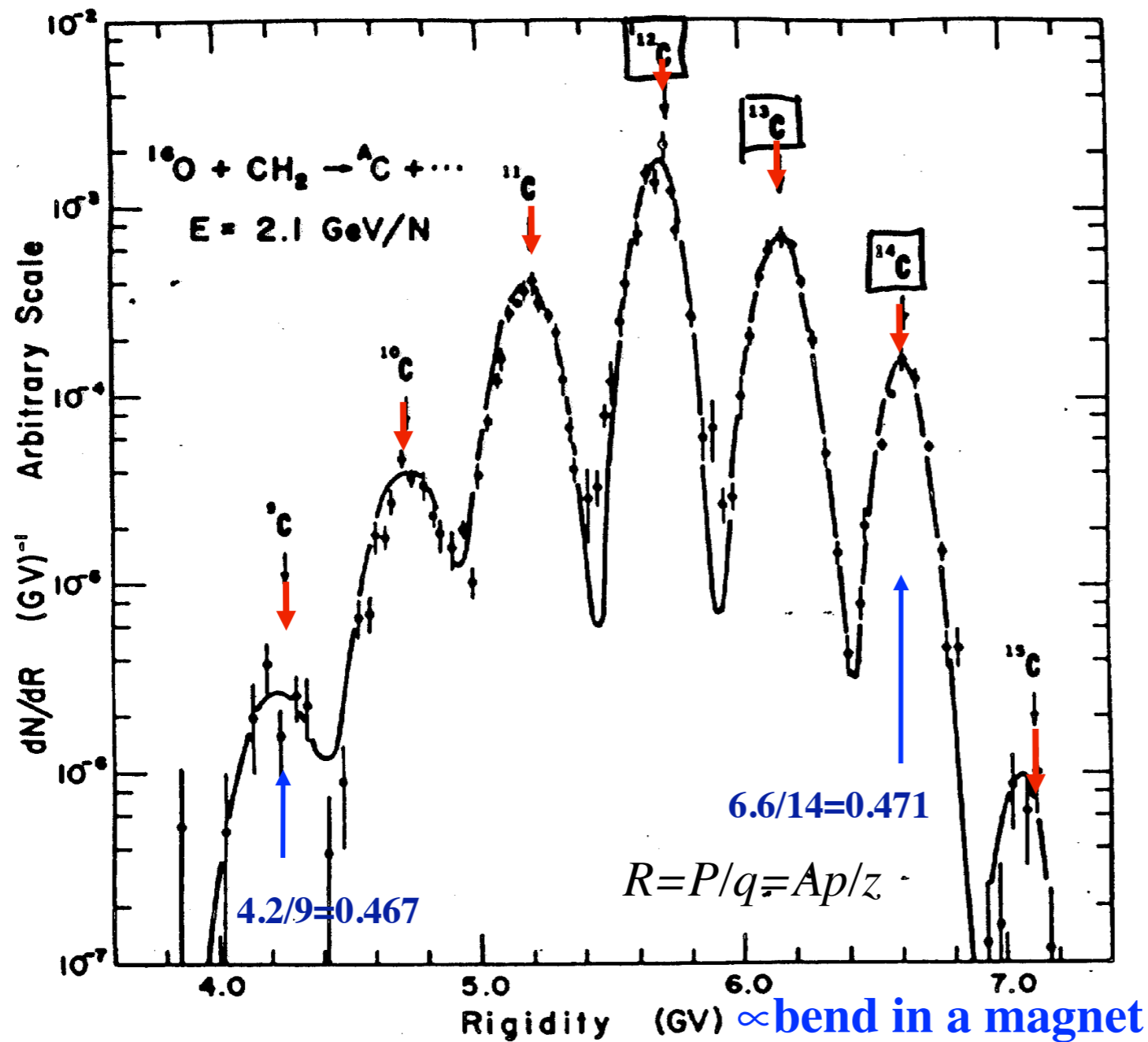
*D. Greiner et al.*

2.1 GeV/nucleon

$\approx 2.88 \text{ GeV/c/nucleon}$

2.83 GeV/c/nucleon

# Projectile fragmentation ( $^{16}\text{O} + \text{CH}_2 \rightarrow {}^A\text{C} + \dots$ )



2.1 GeV/nucleon  
 $\approx 2.88 \text{ GeV/c/nucleon}$

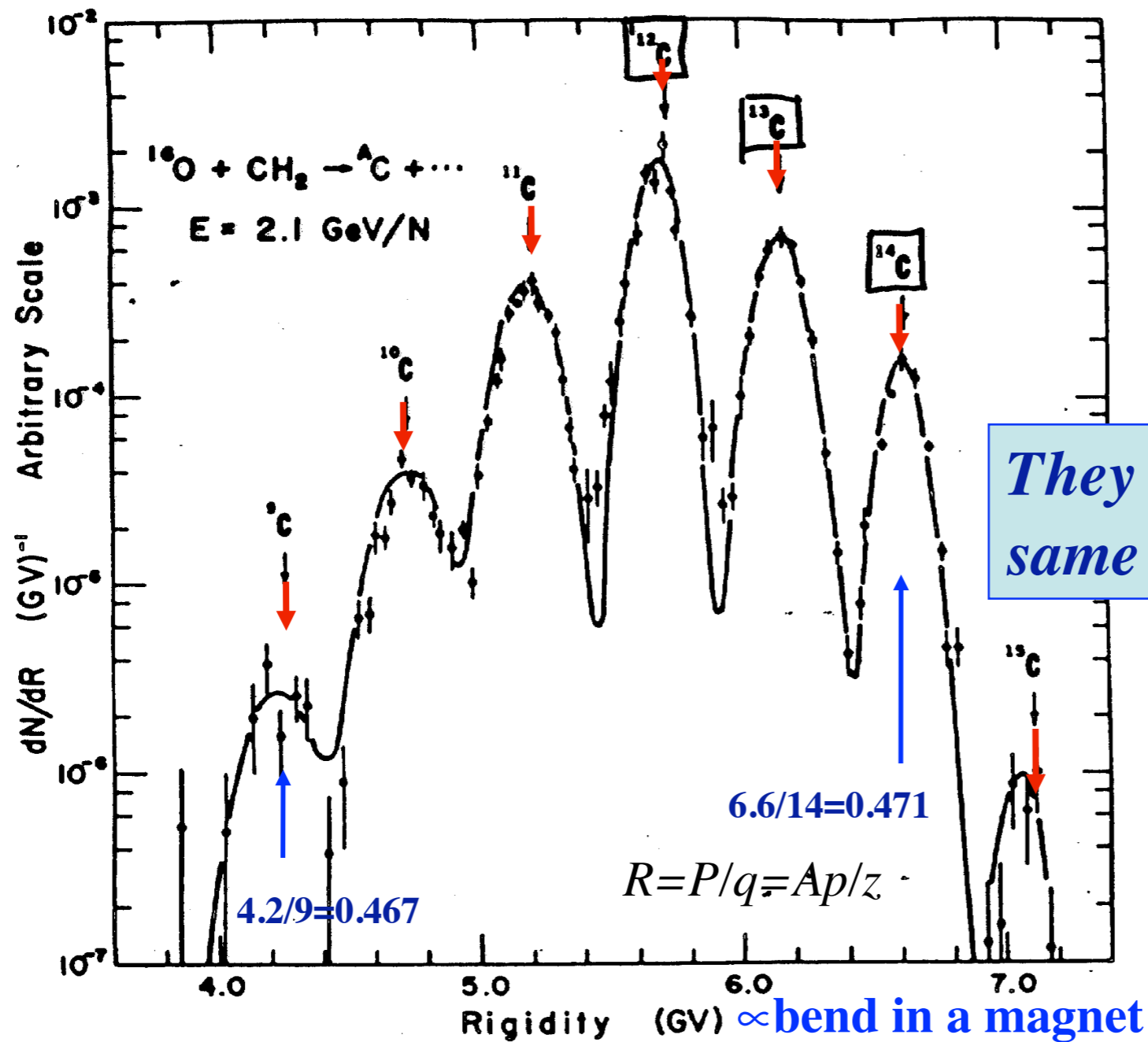
2.83 GeV/c/nucleon

Figure 5 Rigidity spectrum of carbon isotopes produced by the fragmentation of  $^{16}\text{O}$  projectiles at 2.1 A GeV. Arrows indicate the rigidities for each isotope evaluated at beam velocity (87).

*D. Greiner et al.*



# Projectile fragmentation ( $^{16}\text{O} + \text{CH}_2 \rightarrow {}^A\text{C} + \dots$ )



2.1 GeV/nucleon  
 $\approx 2.88 \text{ GeV/c/nucleon}$

2.83 GeV/c/nucleon

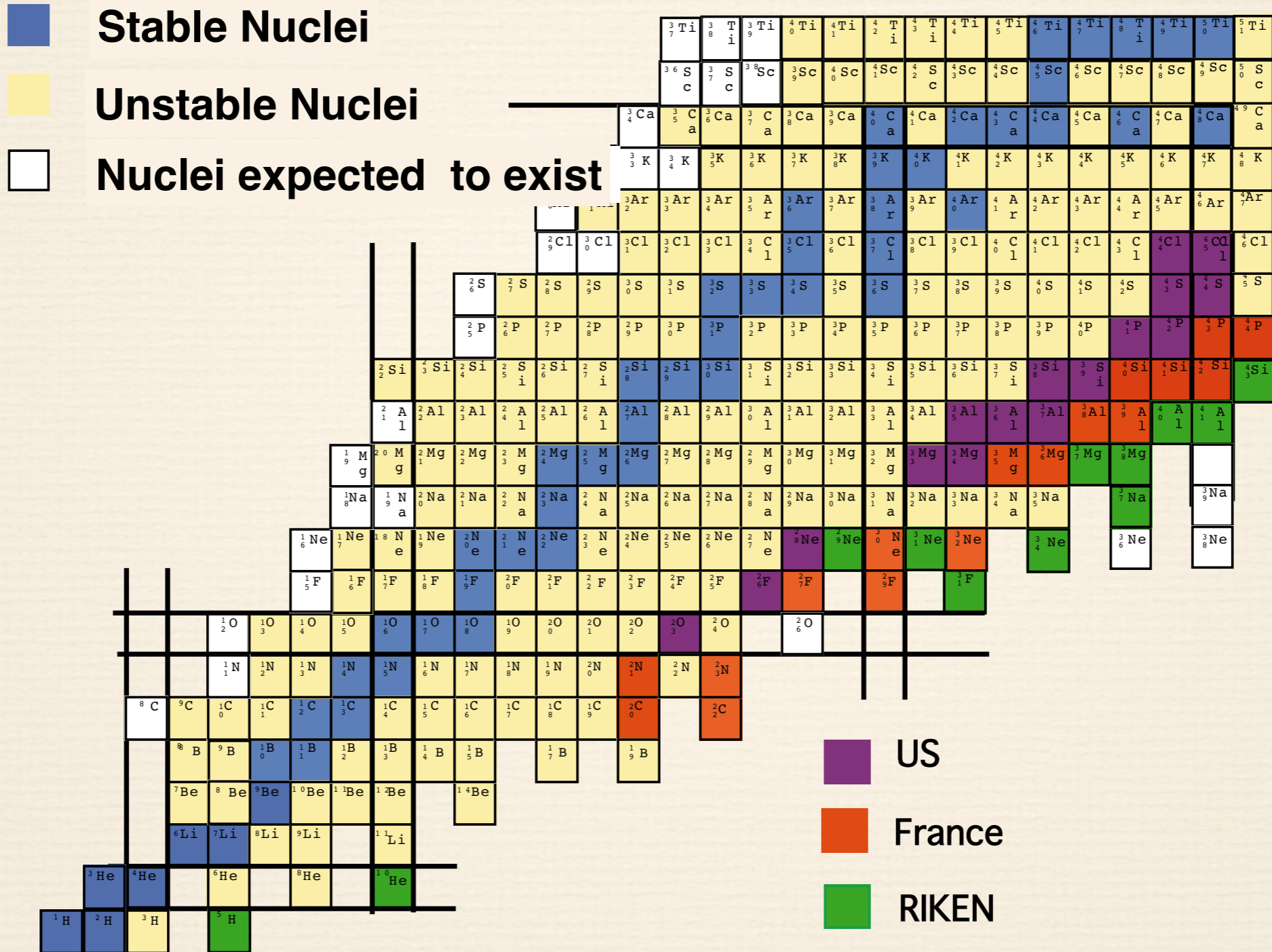
*They are emitted with same velocity!*

Figure 5 Rigidity spectrum of carbon isotopes produced by the fragmentation of  $^{16}\text{O}$  projectiles at 2.1 A GeV. Arrows indicate the rigidities for each isotope evaluated at beam velocity (87).

*D. Greiner et al.*

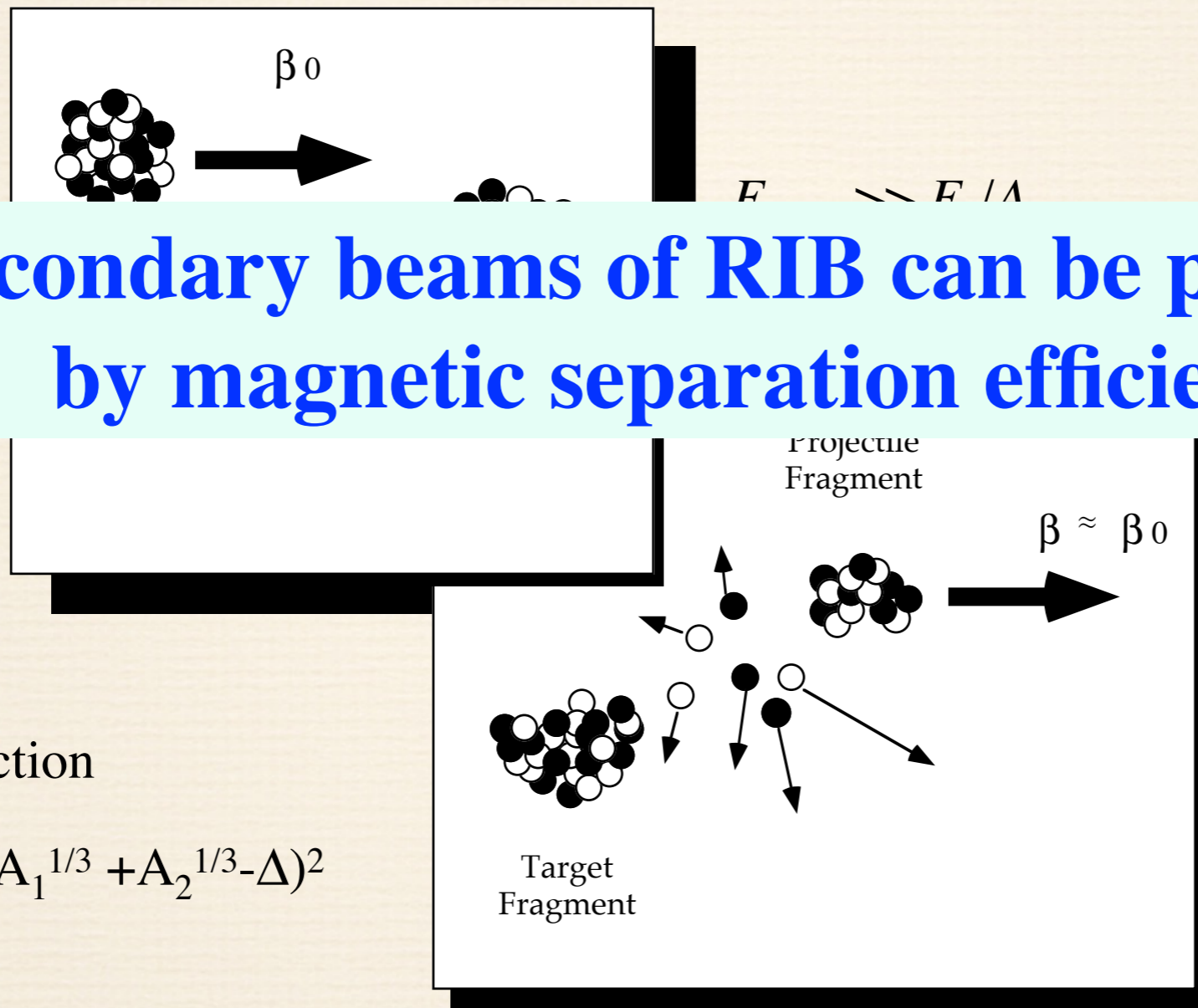
# Starting of new era

by abundant production of nuclei far from the stability line



# Projectile fragmentation and production of Radioactive Ion Beams (RIB)

## Projectile fragmentation



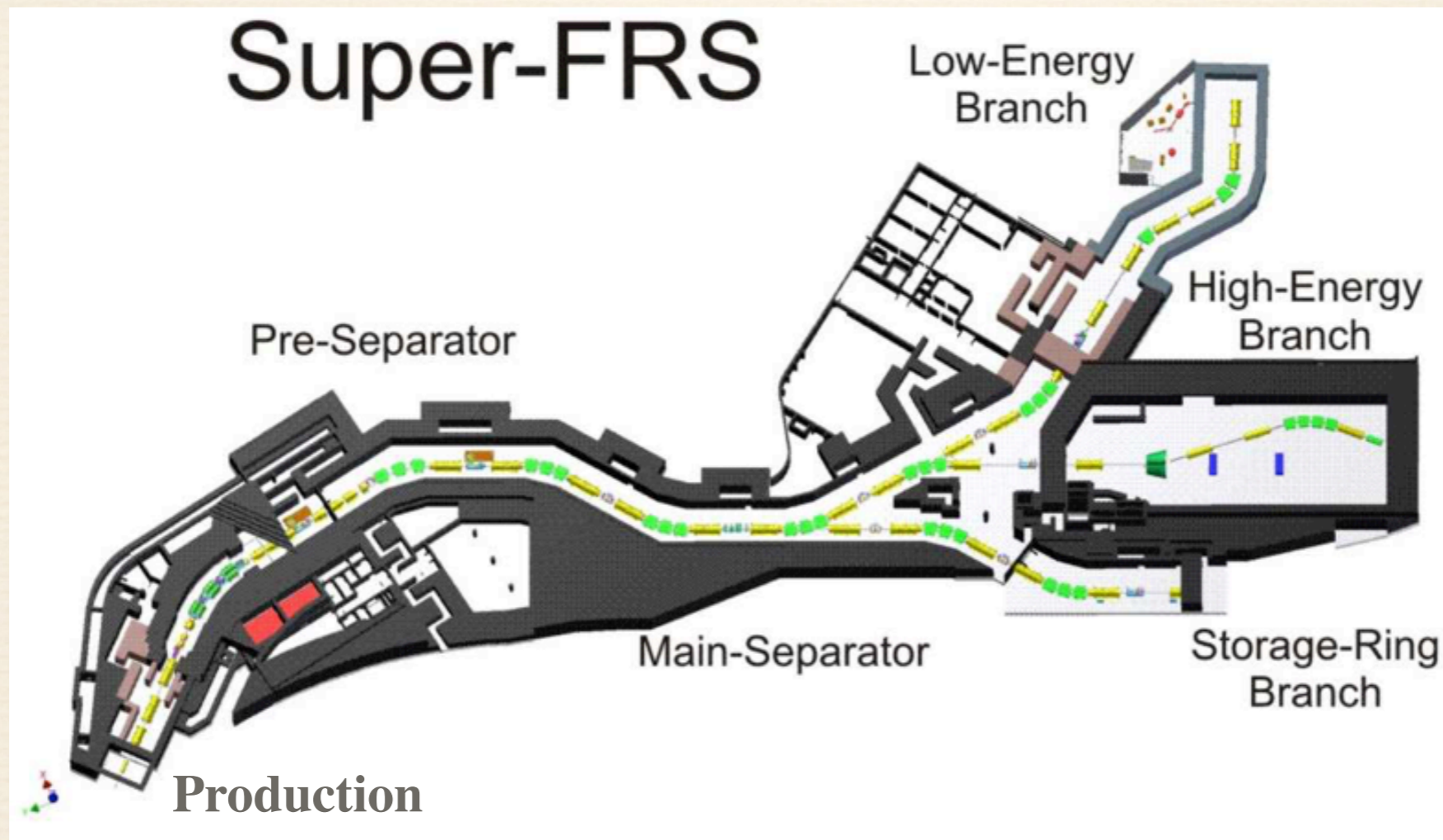
Secondary beams of RIB can be produced by magnetic separation efficiently.

Cross section

$$\sigma_I = \pi r_0^2 (A_1^{1/3} + A_2^{1/3} - \Delta)^2$$

# RIB is now used all over the world

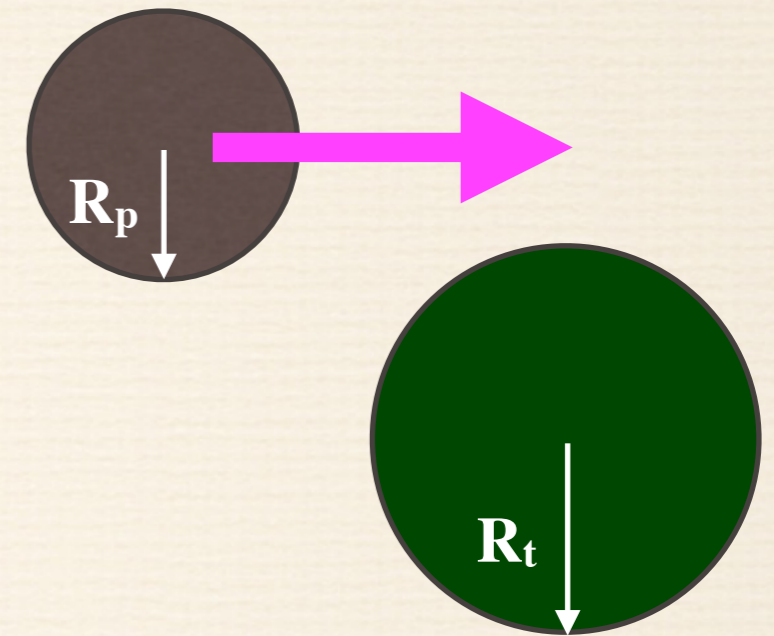
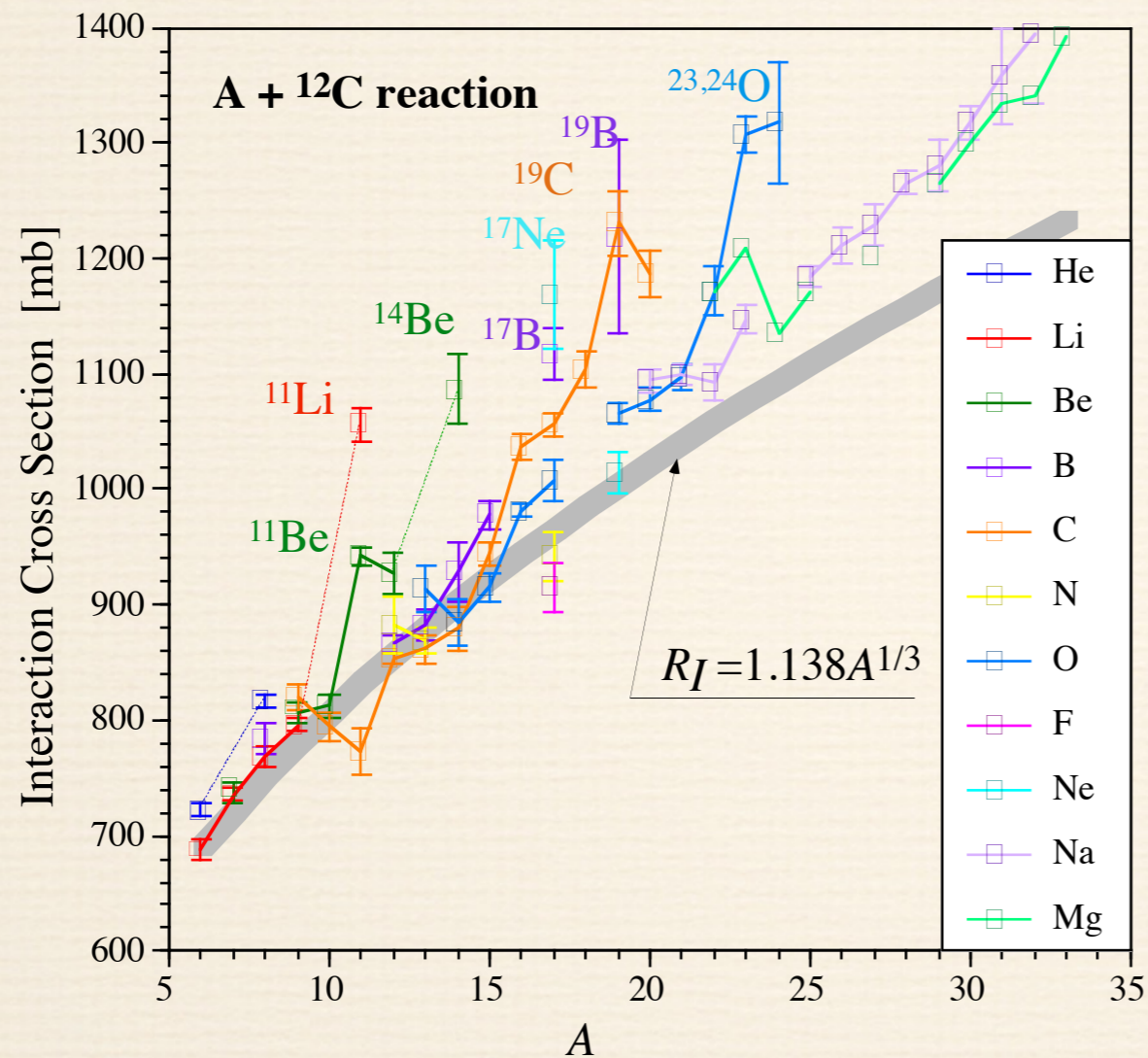
- ❖ New FAIR facility in GSI is one of the most advanced separator in the present stream.



# Interaction Cross Sections $\sigma_I$

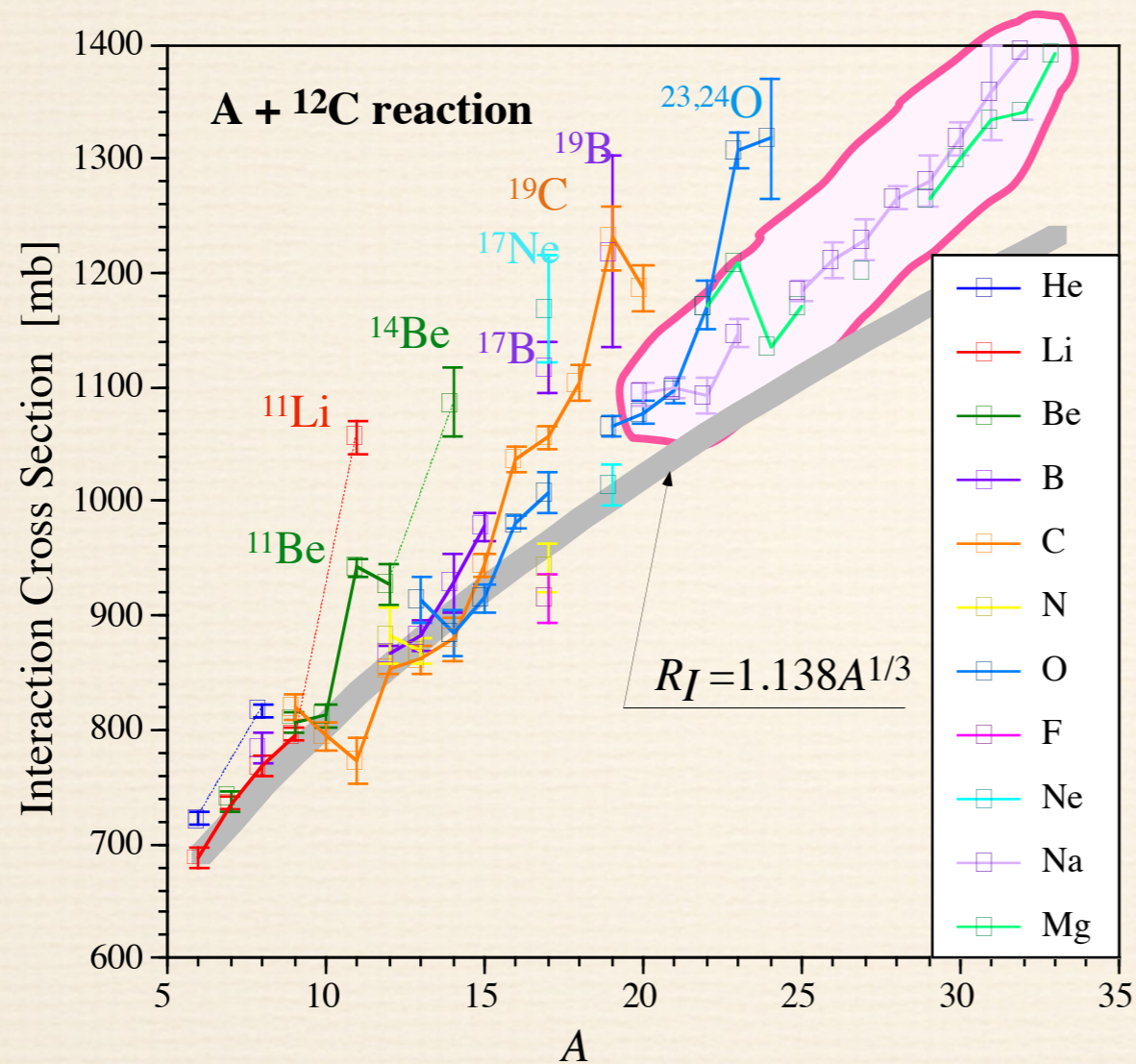
=the cross section is related to the size of a nucleus=

$$\sigma_I \sim \pi(R_p + R_t)^2$$

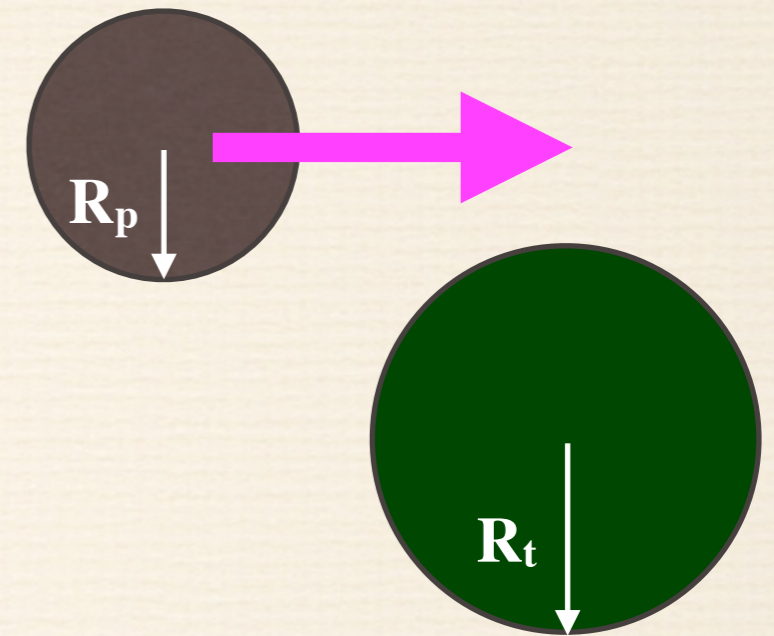


# Interaction Cross Sections $\sigma_I$

=the cross section is related to the size of a nucleus=

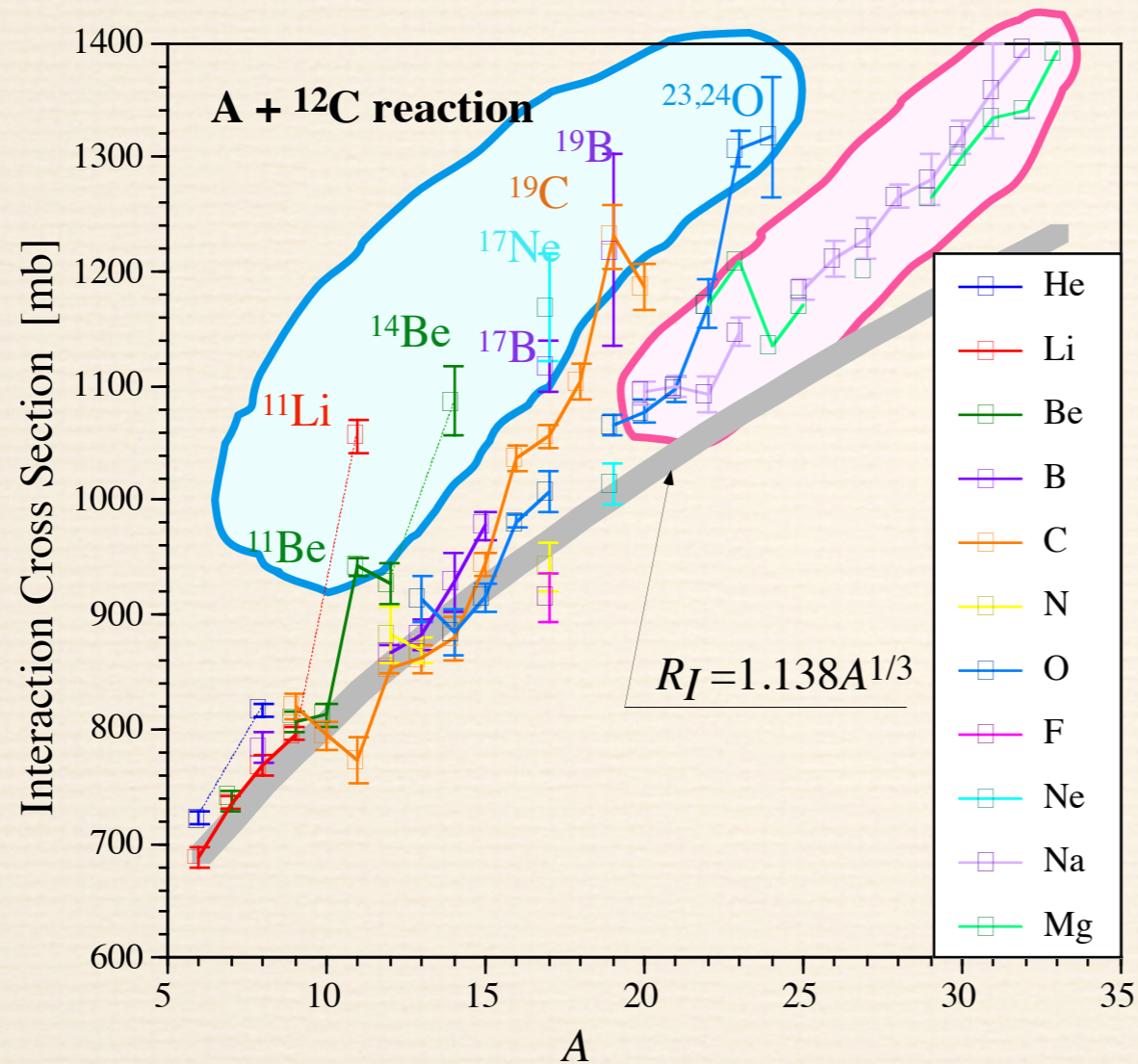


$$\sigma_I \sim \pi(R_p + R_t)^2$$

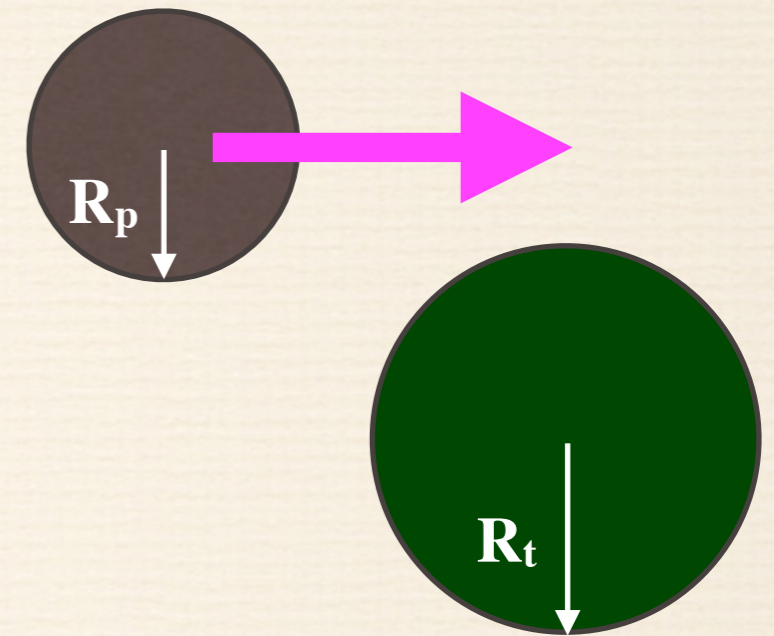


# Interaction Cross Sections $\sigma_I$

=the cross section is related to the size of a nucleus=

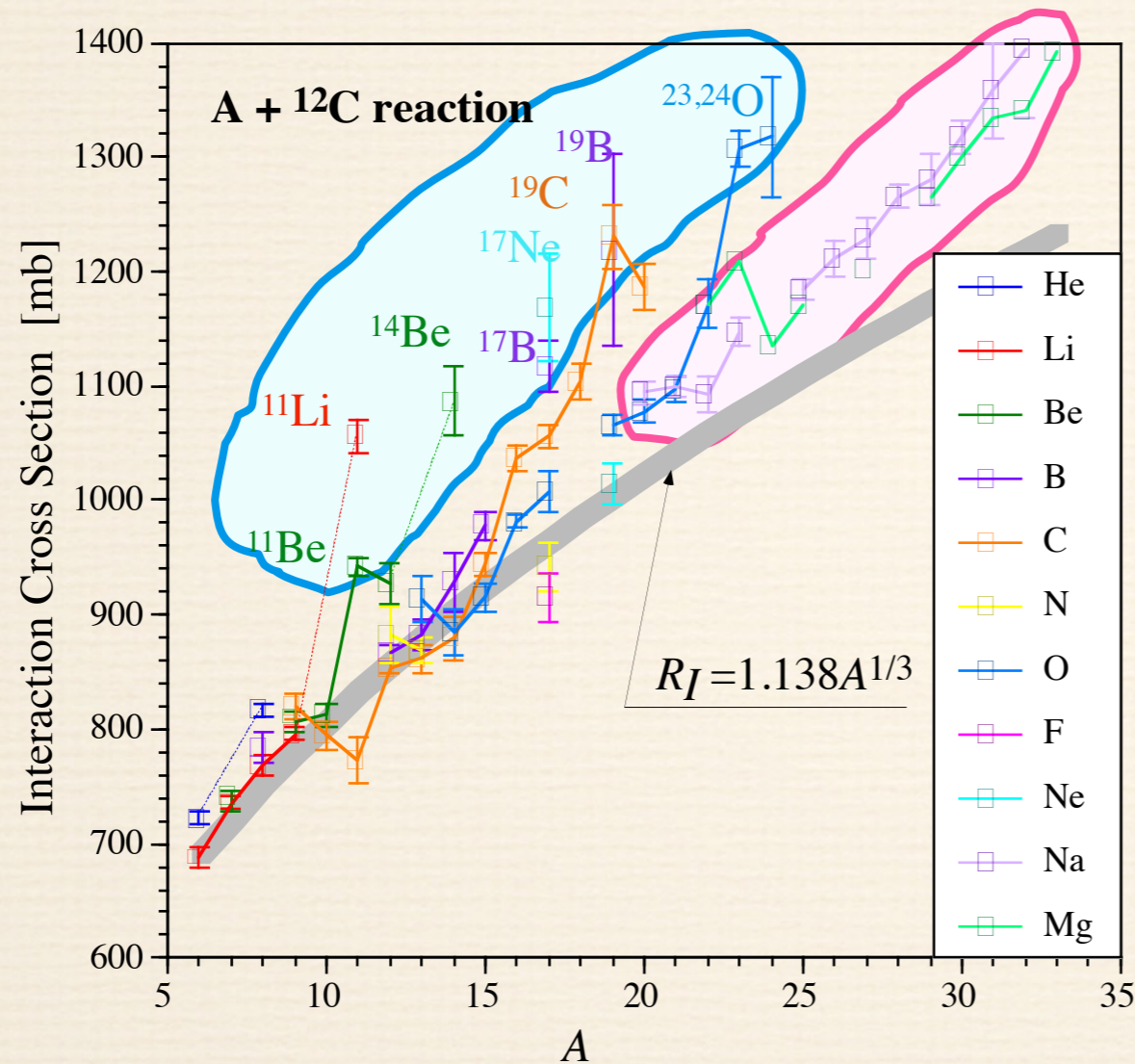


$$\sigma_I \sim \pi(R_p + R_t)^2$$

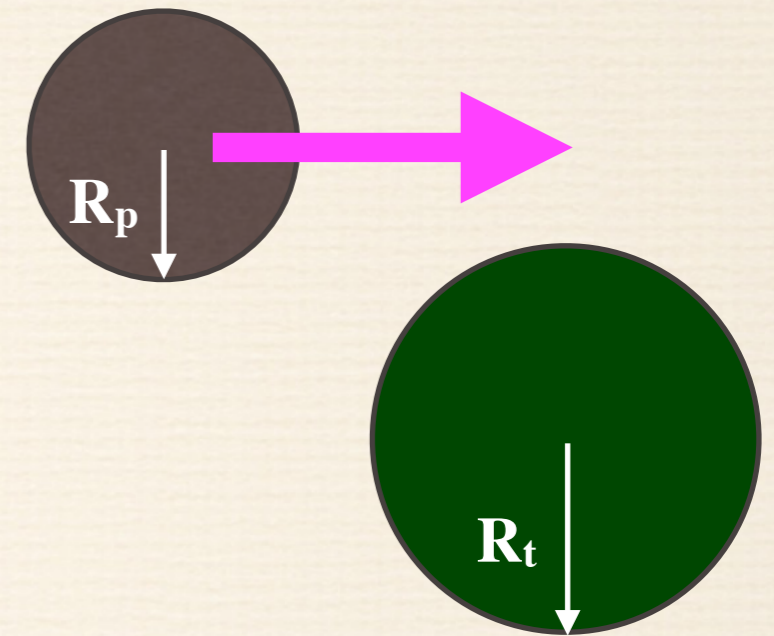


# Interaction Cross Sections $\sigma_I$

=the cross section is related to the size of a nucleus=



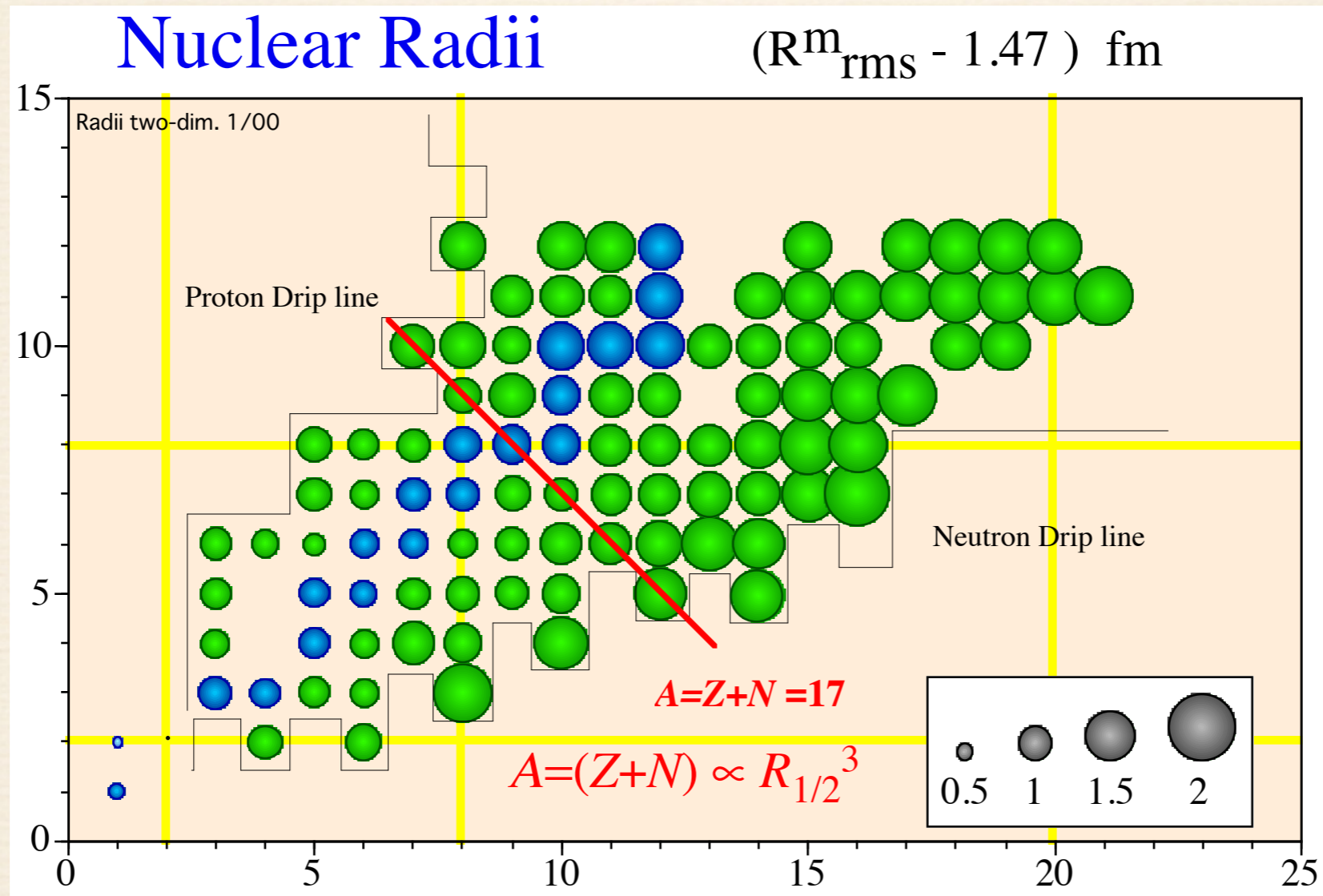
$$\sigma_I \sim \pi(R_p + R_t)^2$$



**Radioactive nuclei does not show the “common properties” !**

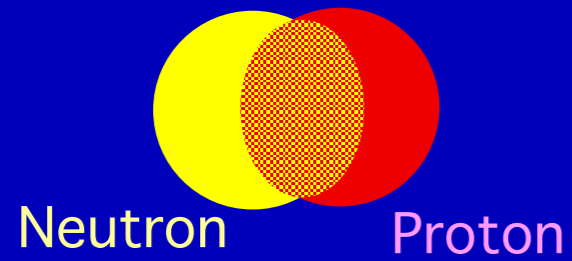


# Breaking of “general property #1”

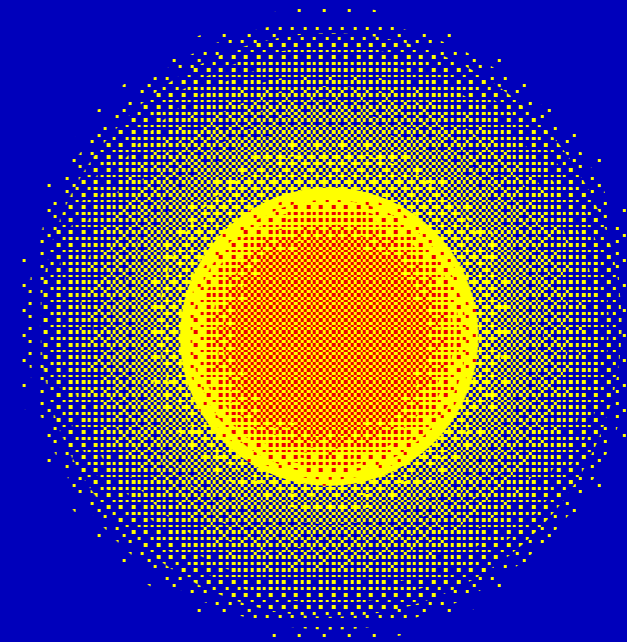
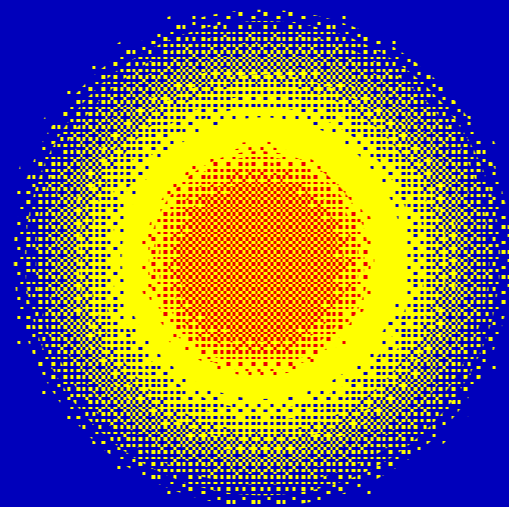
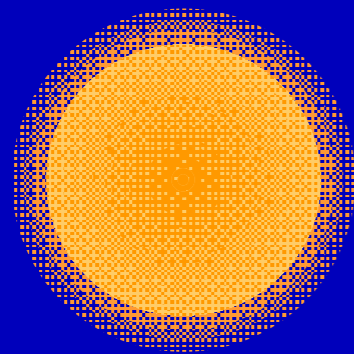


# New Density Distributions

p/n decoupling



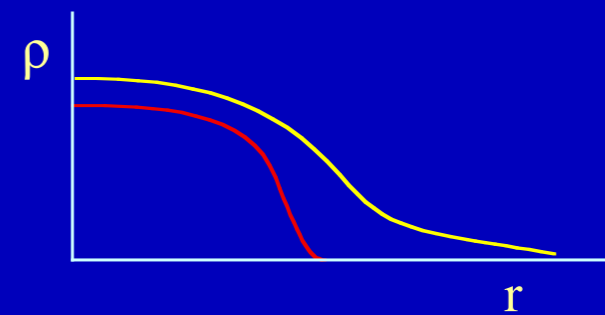
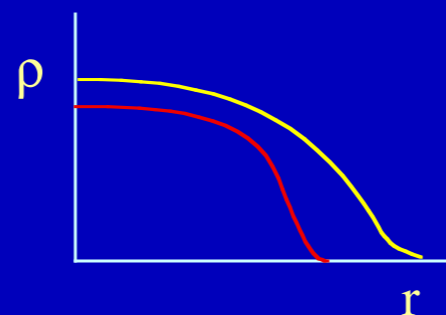
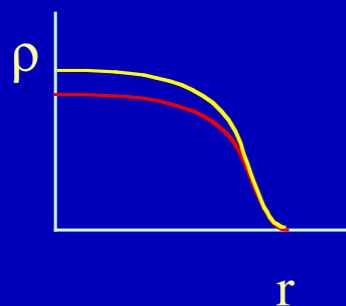
## Nuclear Density Distributions



Neutron skin

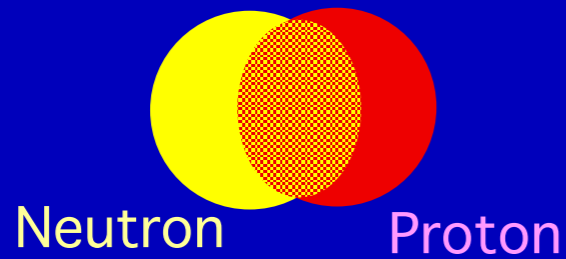
Neutron halo

Stable Nucleus

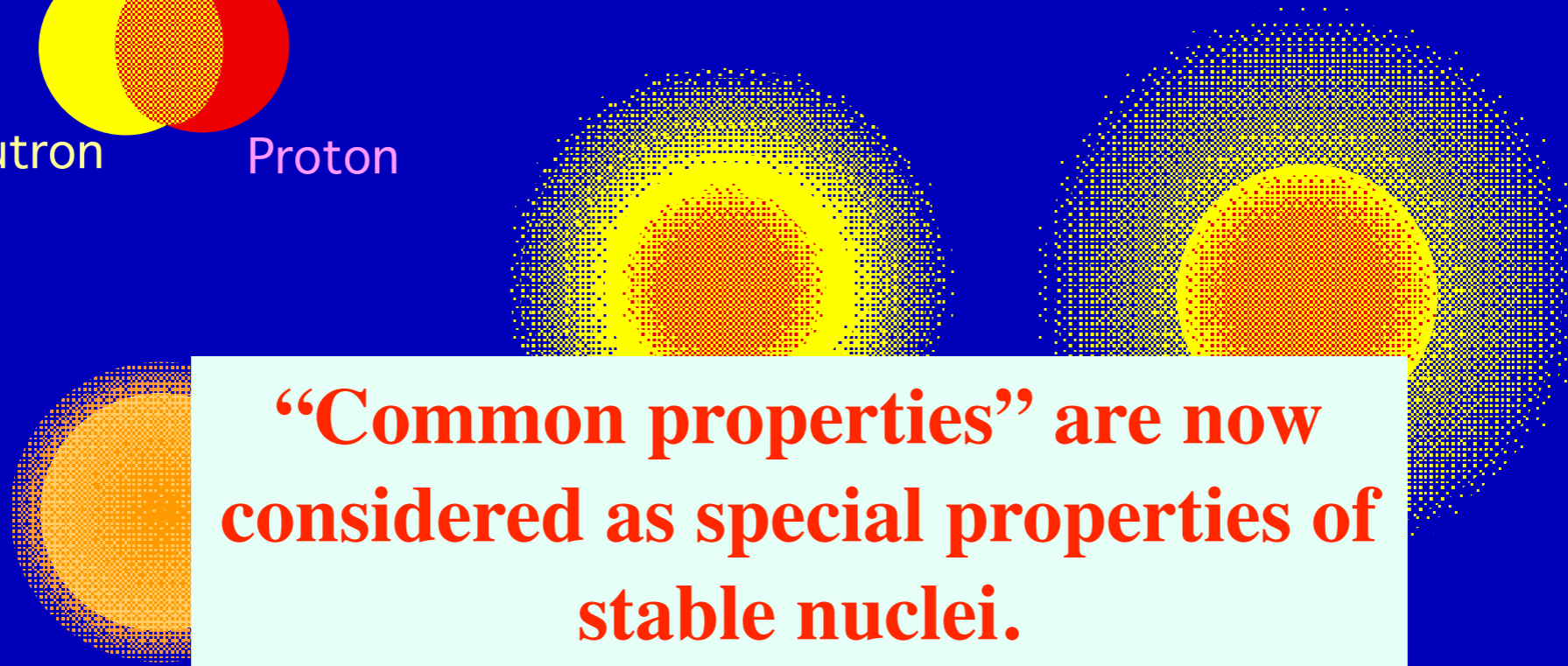


# New Density Distributions

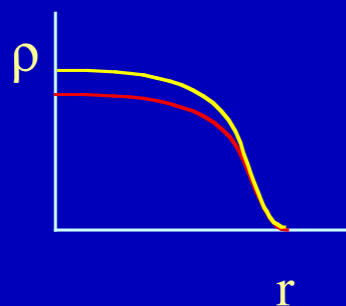
p/n decoupling



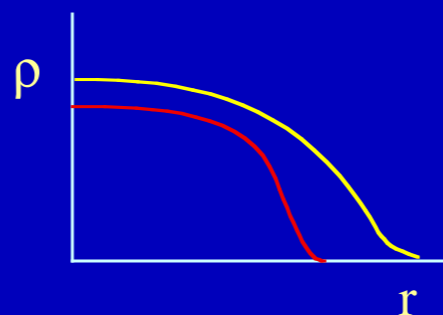
## Nuclear Density Distributions



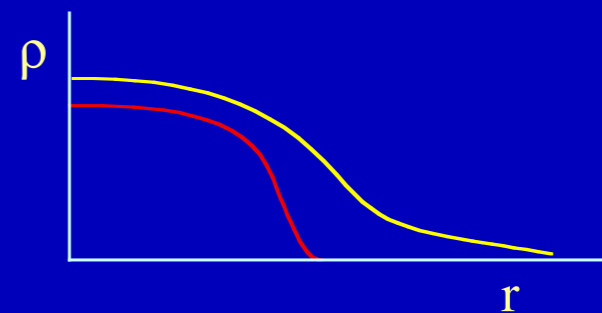
Stable Nucleus



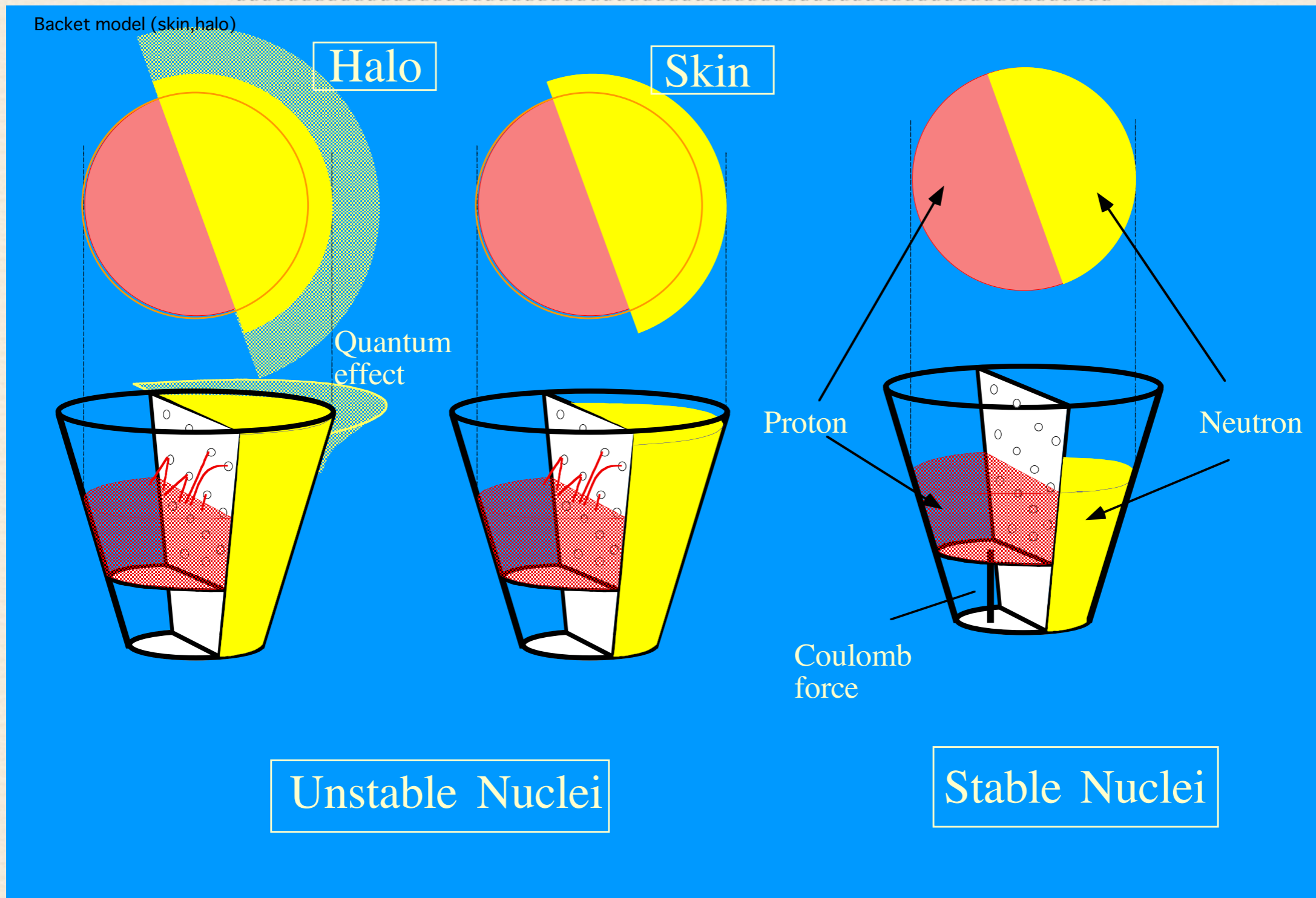
Neutron skin



Neutron halo



# Why such “common properties” in stable nuclei? =Bucket model of nuclear radii=



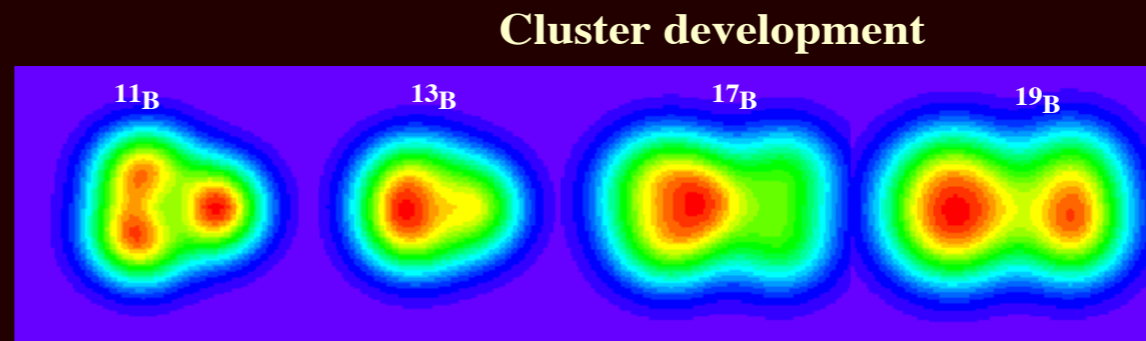
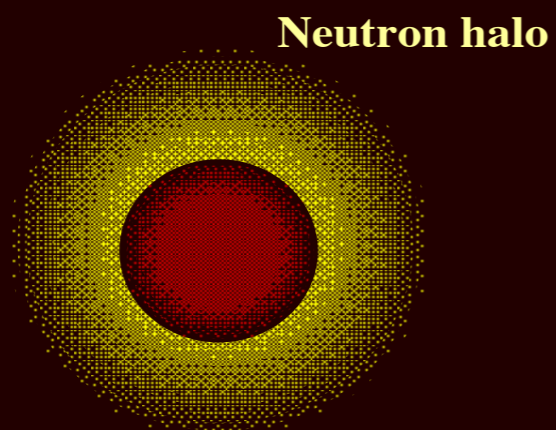
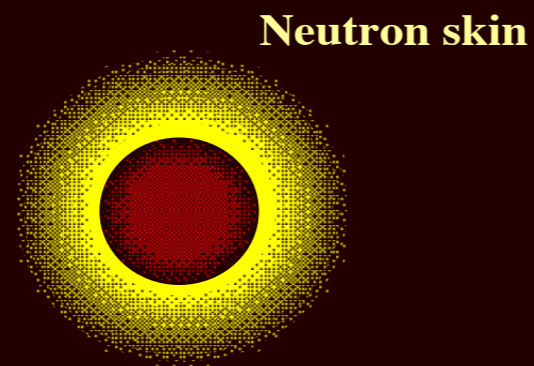
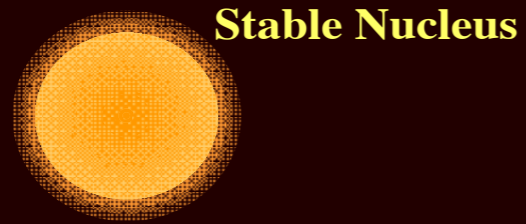
# **Not only the shops of nuclei, but other properties have to be restudied.**

---

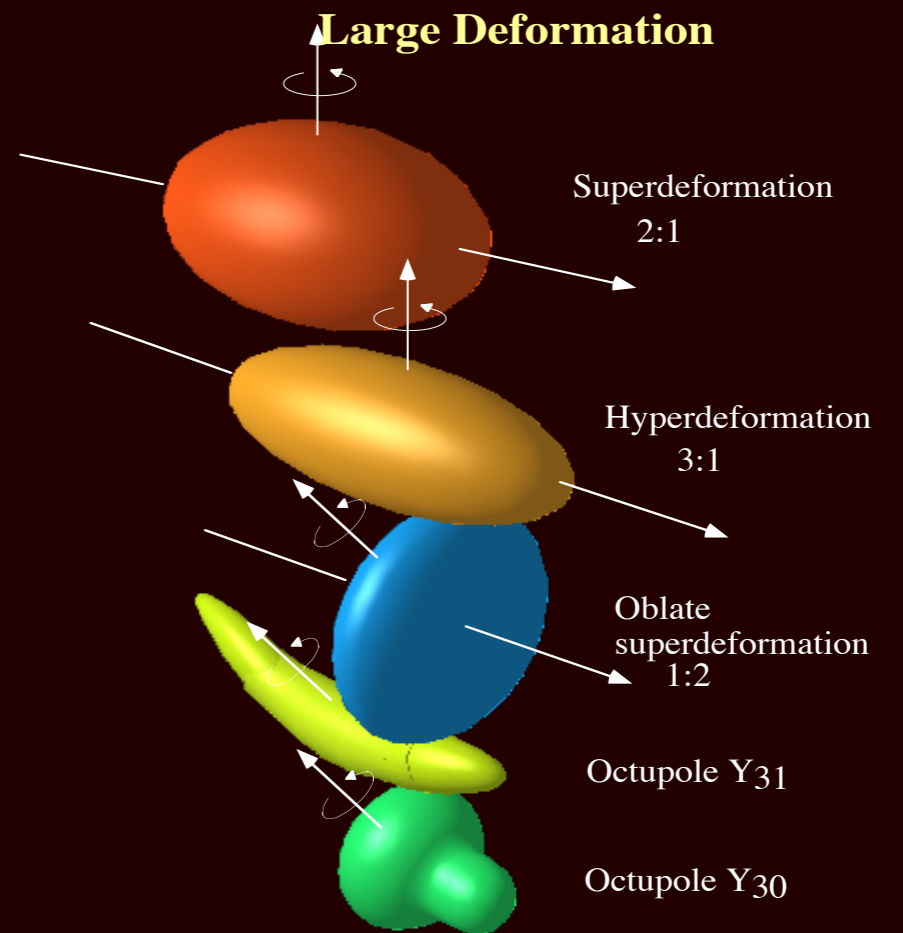
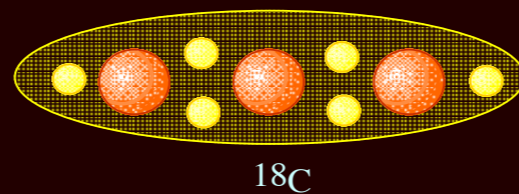
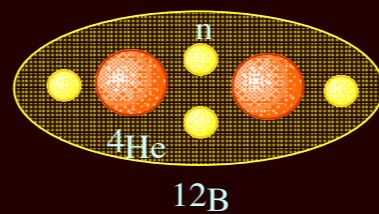
- ❖ **Weakly bound nuclei show new excitation modes.**
- ❖ **Large difference of proton and neutron number present new properties.**
- ❖ **Magic numbers changes far from the stability line.**
- ❖ **Clusters and molecular binding type nuclei.**
- ❖ **...**

# Other forms of Nuclei

Nuclear shapes



## Nuclear Polymer

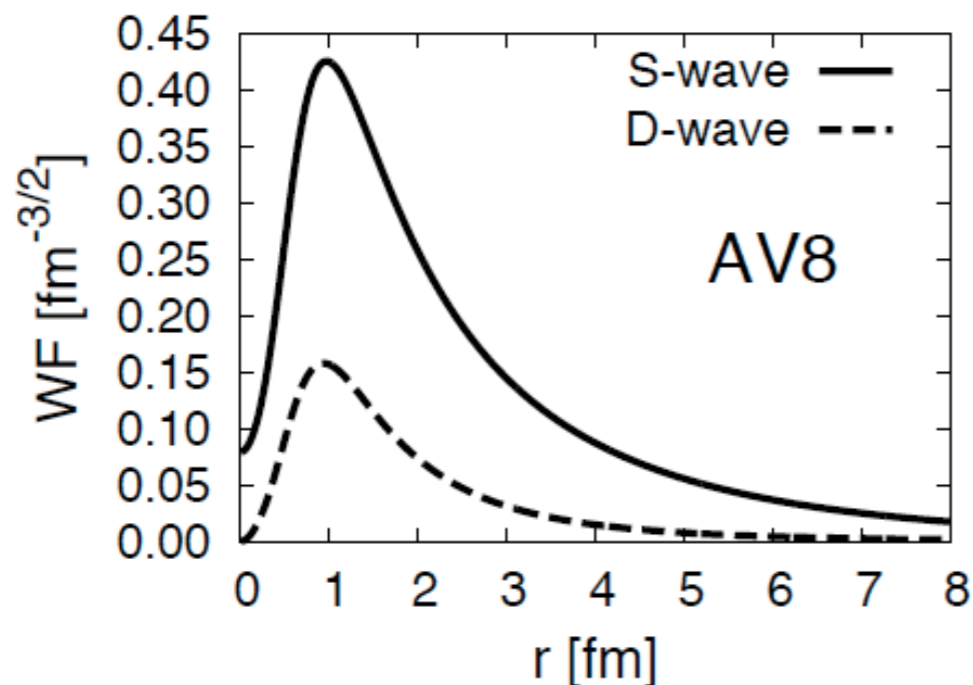
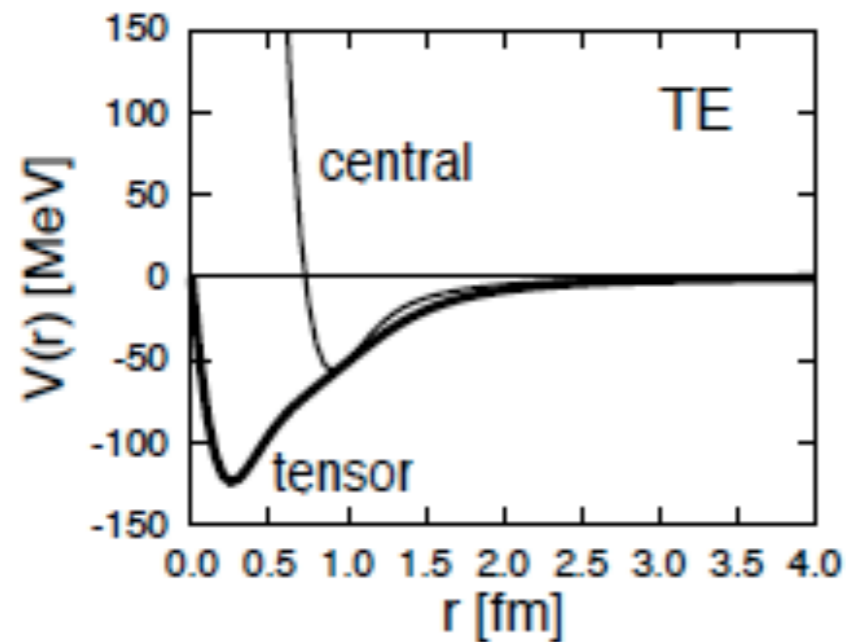


# For Nuclear Physicist

## The importance of tensor is clear in deuteron

$S=1$  and  $L=0$  or  $2$

### Binding of deuteron ( $1^+$ )



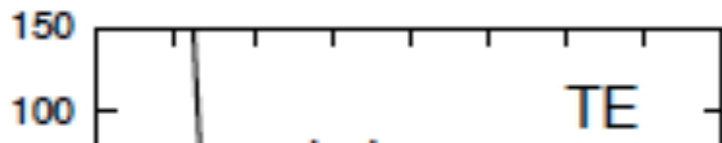
Energy	-2.24 [MeV]
Kinetic	19.88
(SS)	11.31
(DD)	8.57
Central	-4.46
(SS)	-3.96
(DD)	-0.50
Tensor	-16.64
(SD)	-18.93
(DD)	2.29
LS	-1.02
P(D)	5.78 [%]
Radius	1.96 [fm]
(SS)	2.00 [fm]
(DD)	1.22 [fm]

K.Ikeda, T.Myo, K.Kato, and H.Toki  
Lecture Notes in Phys.818(2010) 165.

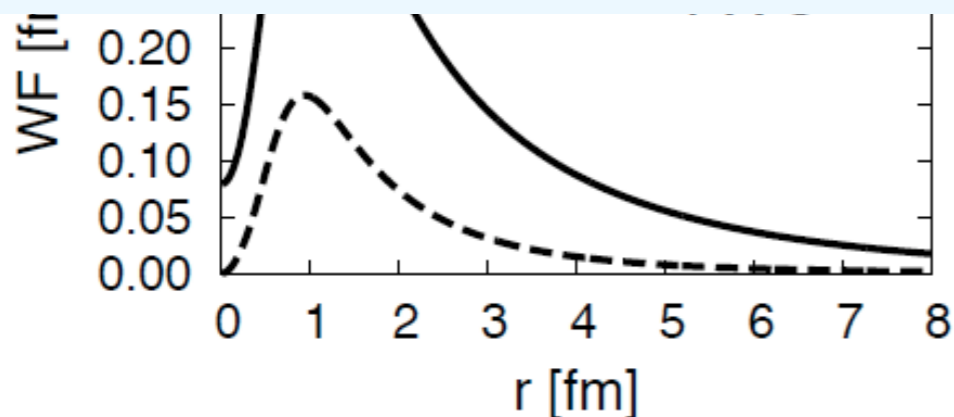
# The importance of tensor is clear in deuteron

$S=1$  and  $L=0$  or  $2$

Binding of deuteron ( $1^+$ )



1. Tensor interactions provides most of the binding energy.
2. It is due to the D-wave mixing through the tensor interactions
3. The binding energy by tensor interactions are not from  $D^2$  term but from SD cross term.
4. D wave has shorter range and thus has high-momentum.
5. High momentum nucleon are necessary to make binding.

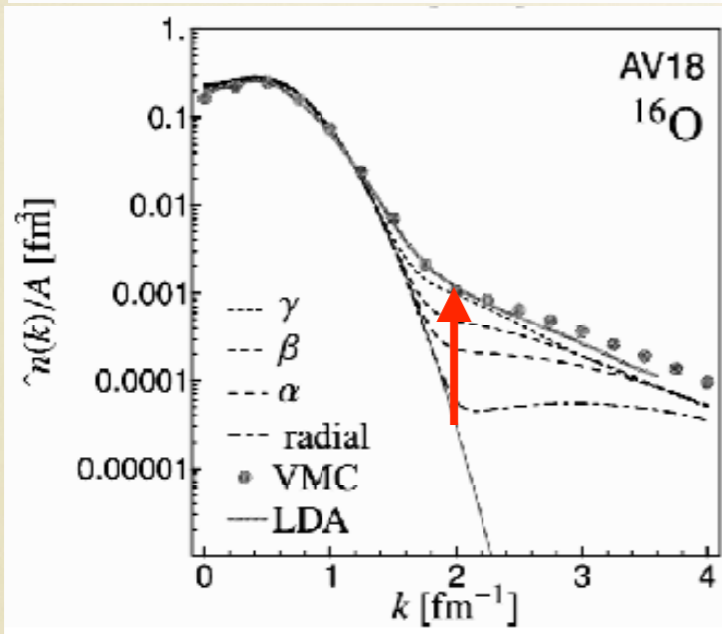
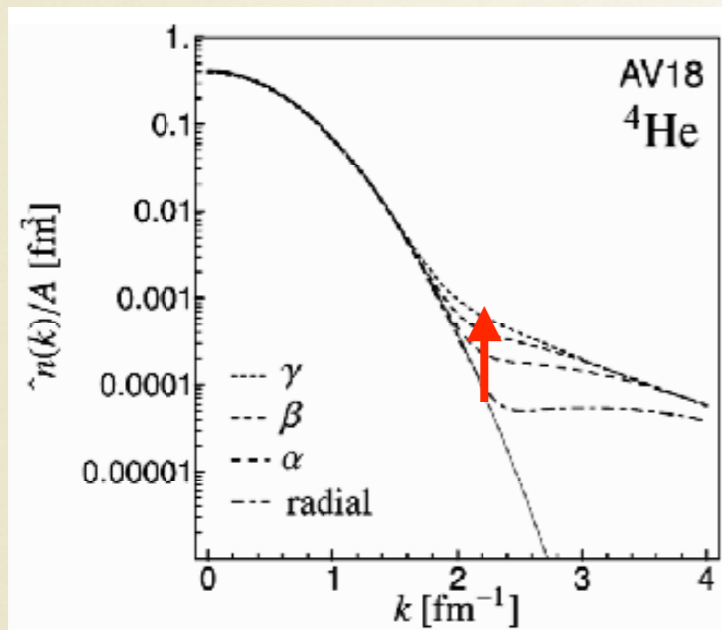


$P(D)$	5.78 [%]
Radius	1.96 [fm]
(SS)	2.00 [fm]
(DD)	1.22 [fm]

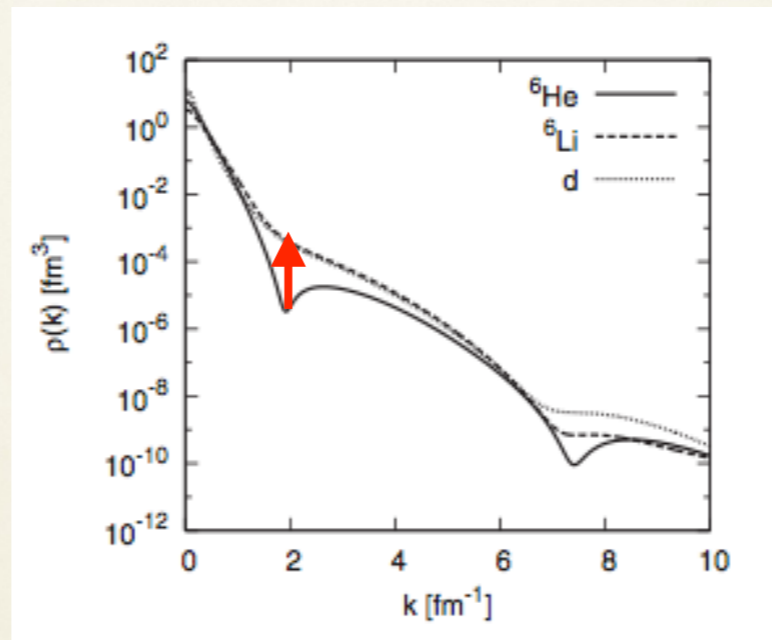
K.Ikeda, T.Myo, K.Kato, and H.Toki  
Lecture Notes in Phys.818(2010) 165.



# HIGH-MOMENTUM COMPONENTS (THEORETICAL PREDICTIONS)

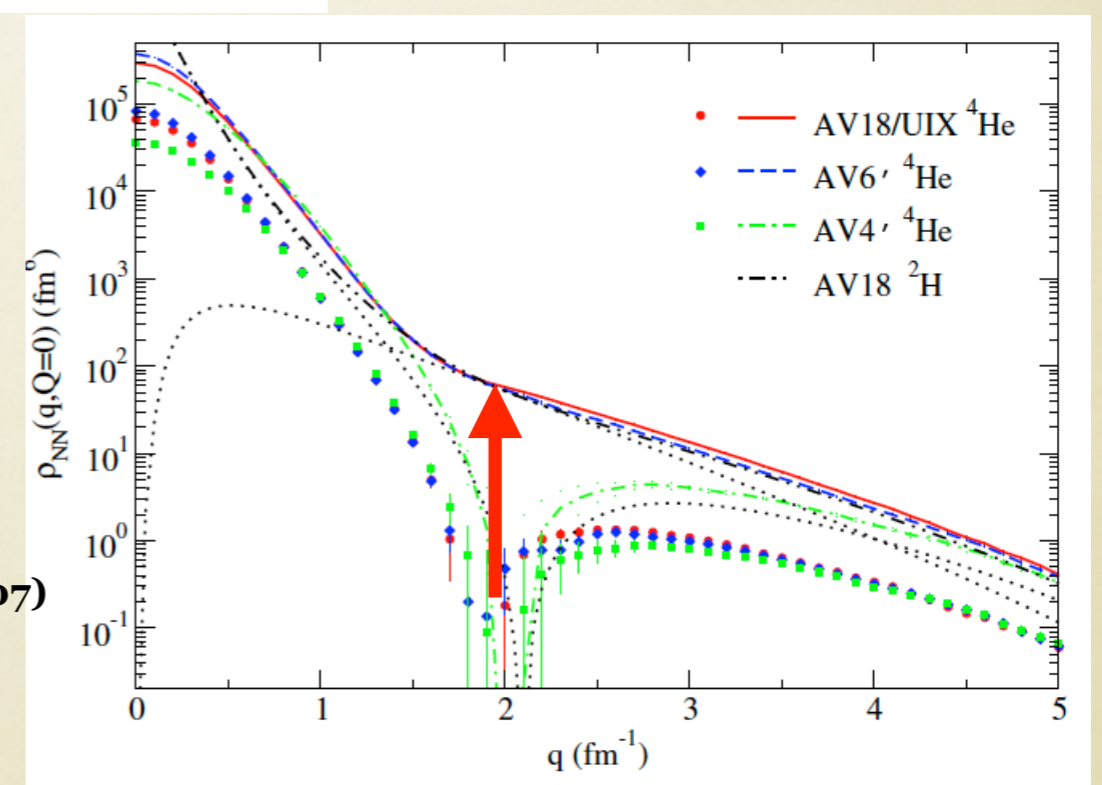


T. Neff and H. Feldmeier,  
NPA713, 311(2003)



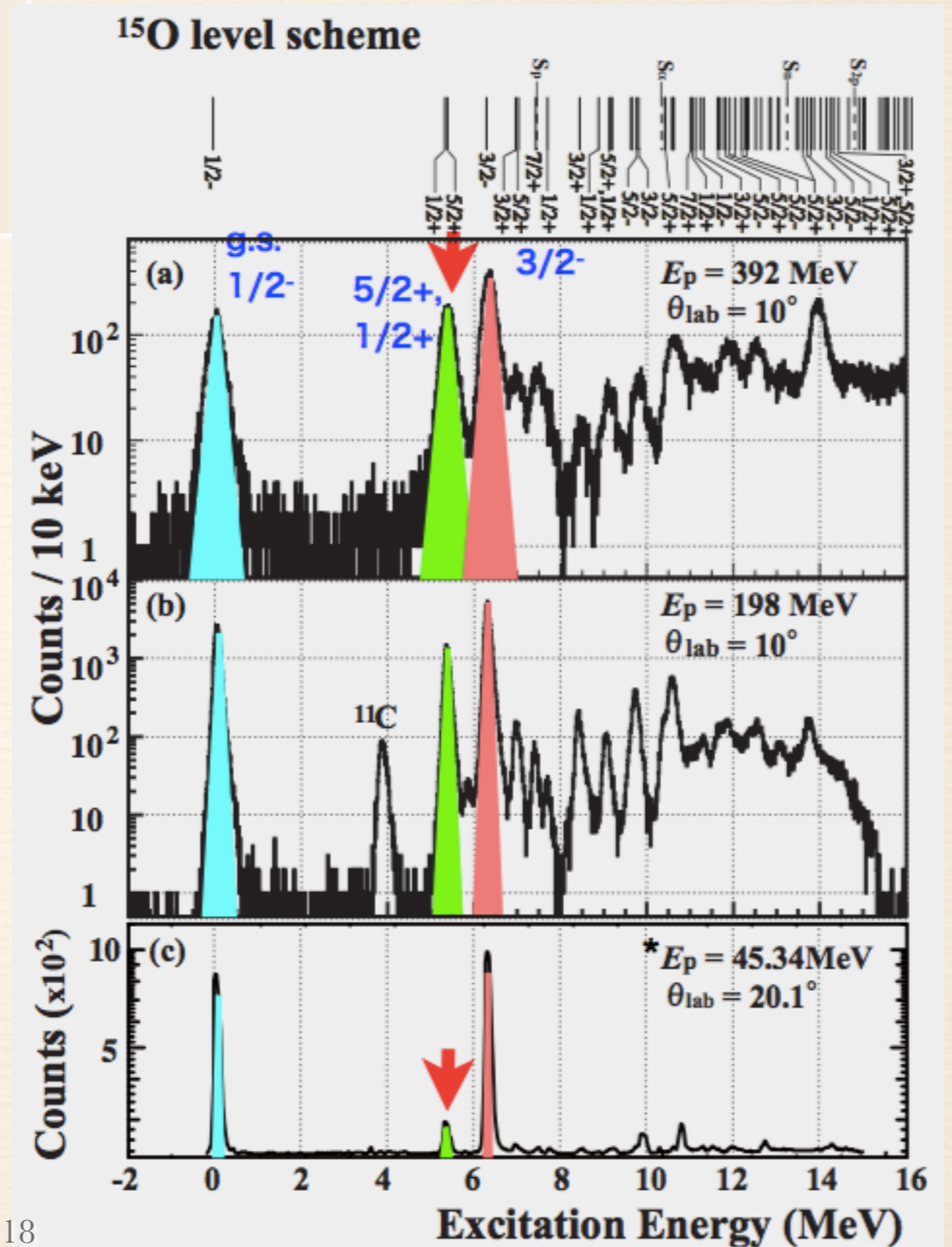
W. Horiuchi and Y. Suzuki,  
PRC76, 024311(2007)

R. Schiavilla et al.,  
PRL 98 132501 (2007)

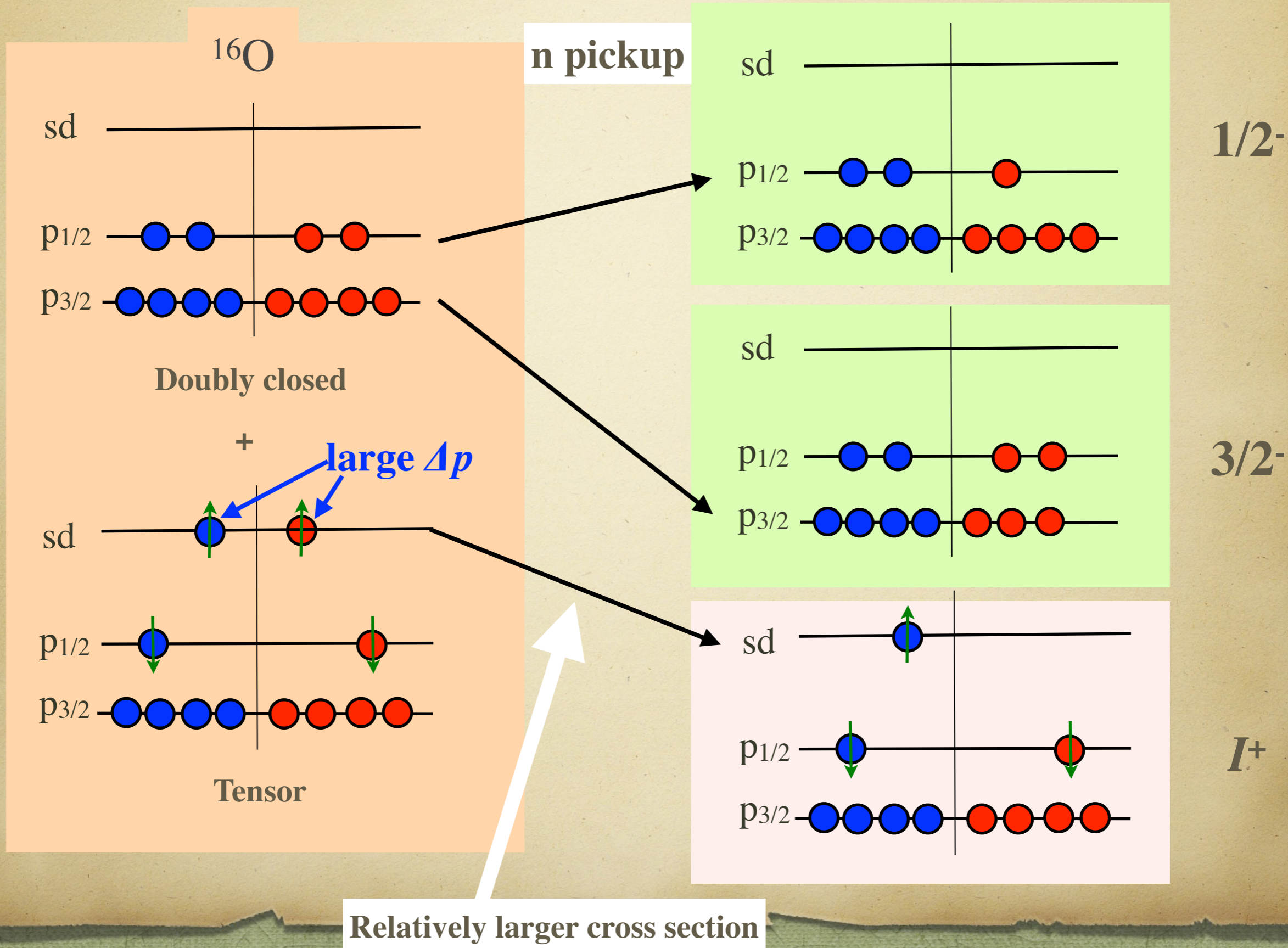


# State dependent effect by (p,d) reaction

- ❖ Another possibility
- ❖  ${}^6\text{He}(p,d){}^4\text{He}+X$   
(2n)



# *$^{16}\text{O}$ and $(p,d)$ reaction*



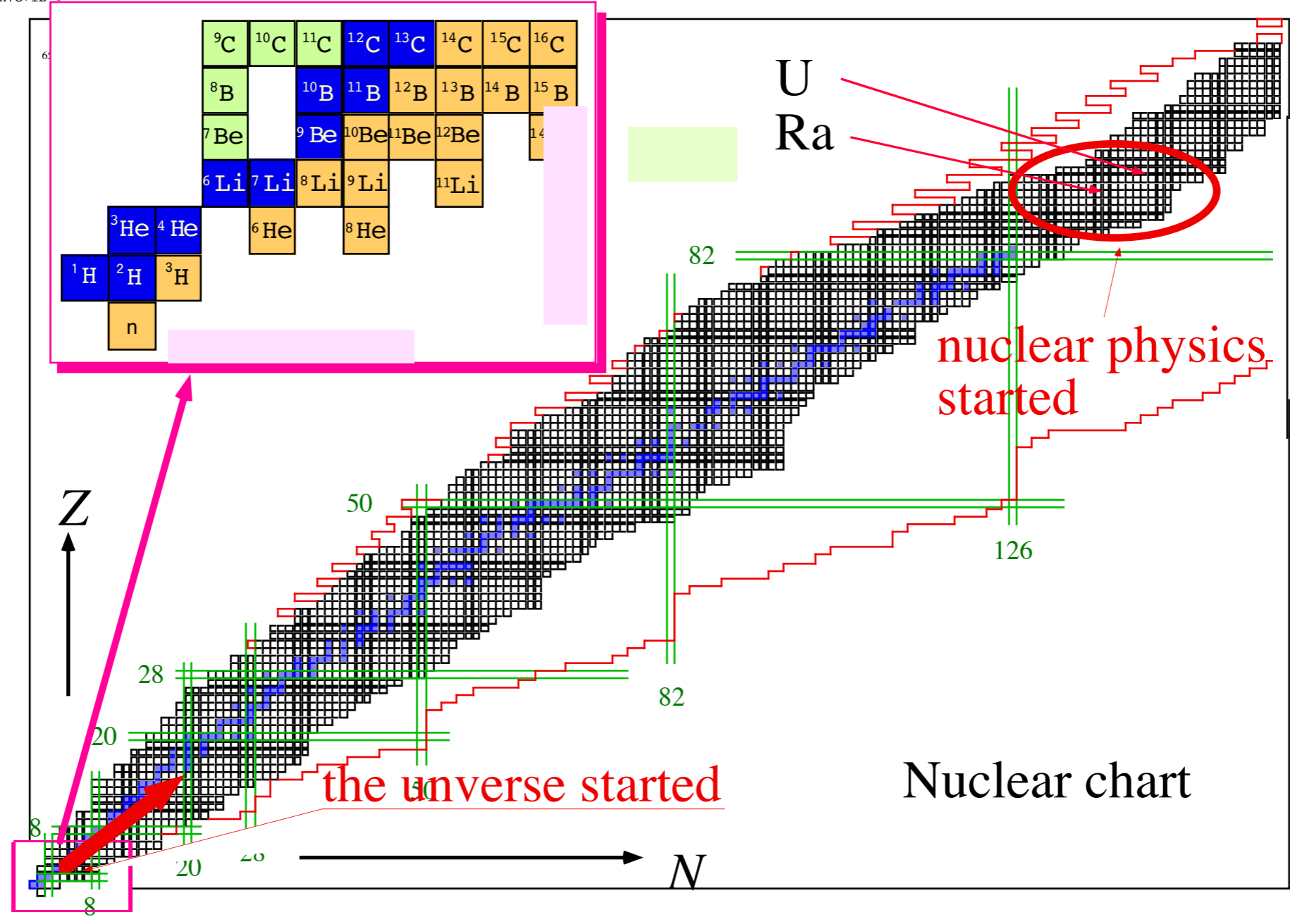
# Other use of RIB in research

---

- ❖ **Nuclear astrophysics, Cosmo-nuclear physics**
- ❖ **Implantation of RI**
- ❖ **...**

Why RIB

noeāncē, iē-1

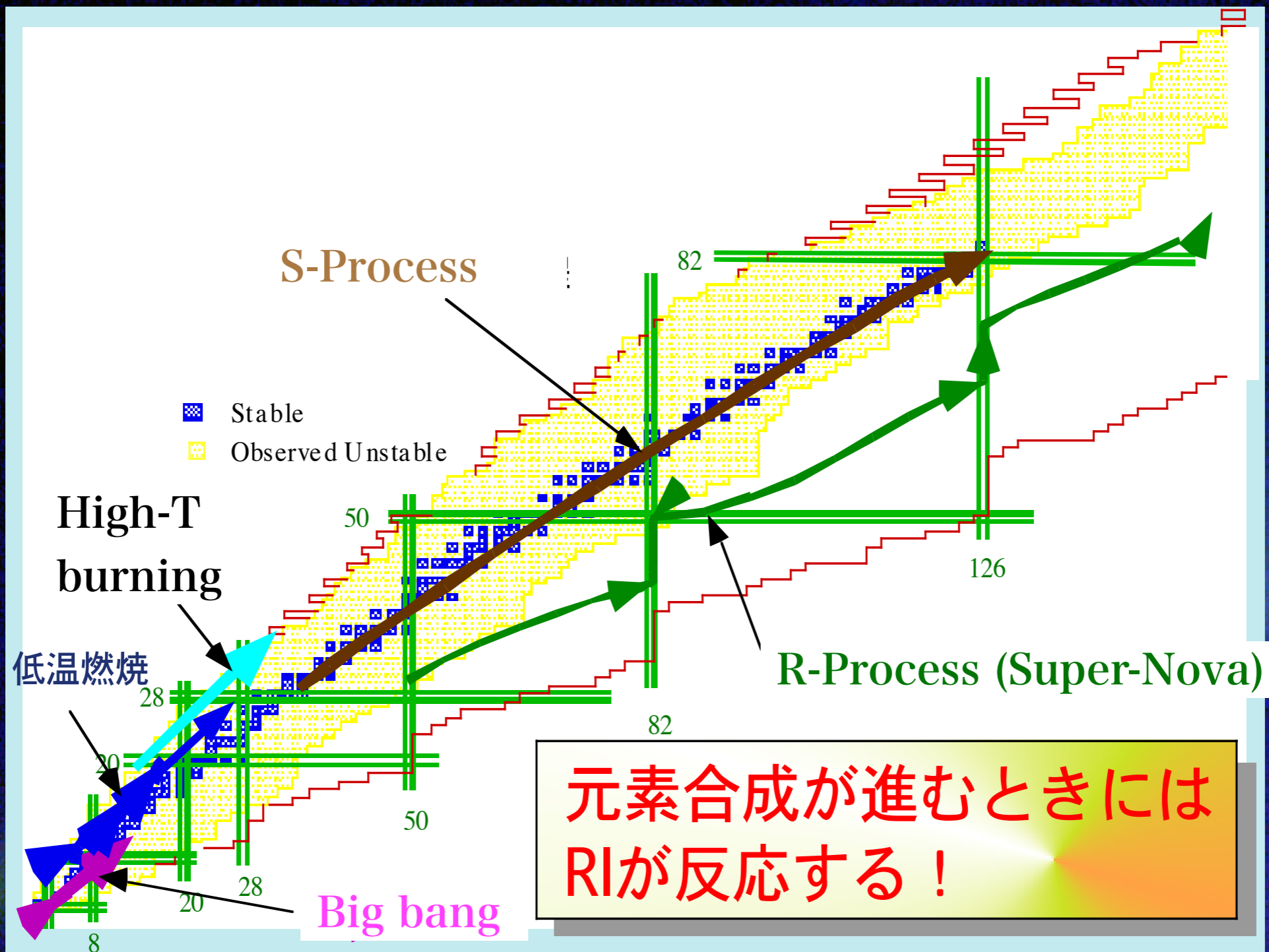


nuclear physics started

the universe started

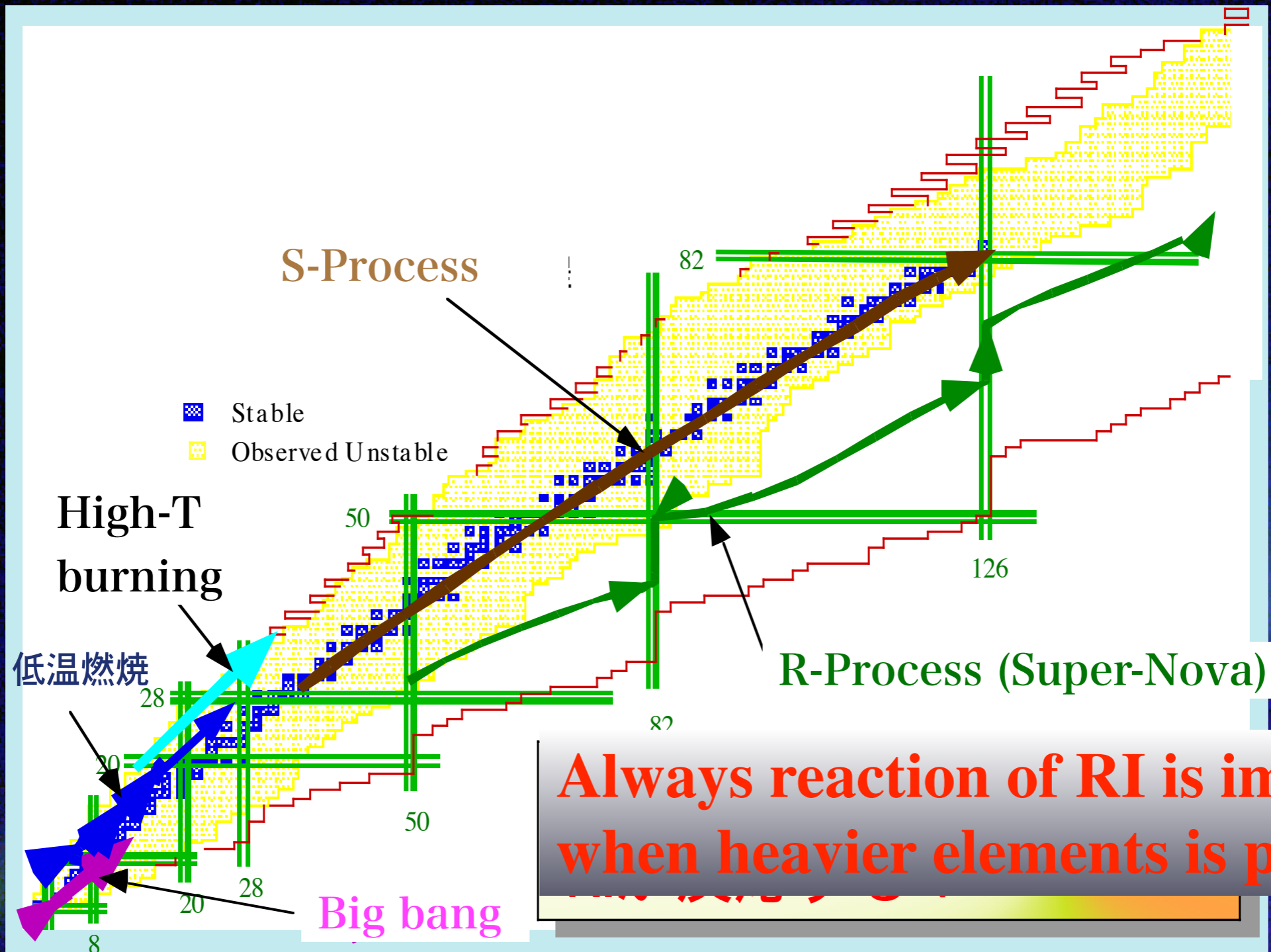
Nuclear chart

# Synthesis Paths



これらの反応の研究は  
RIビームで初めて可能

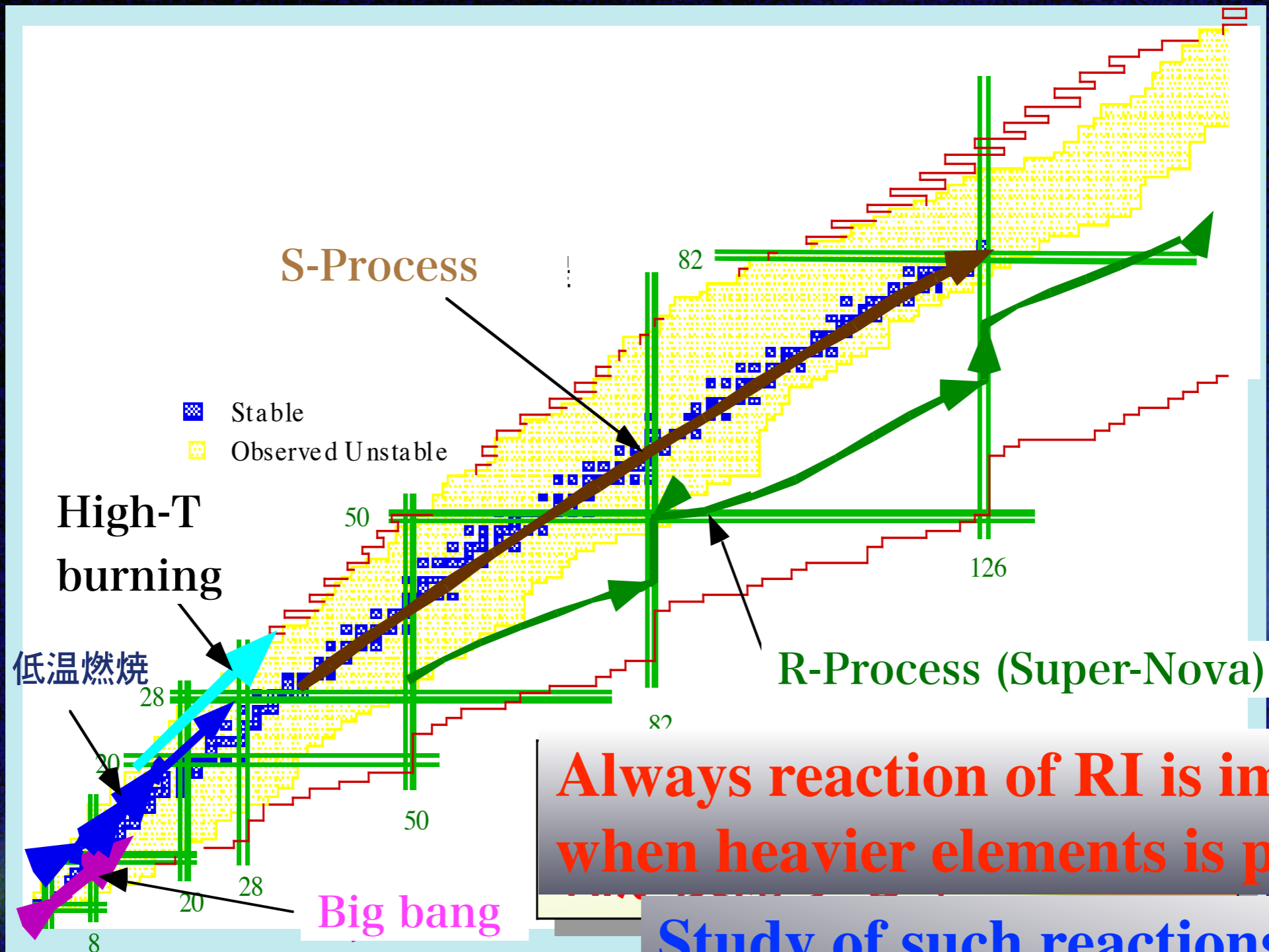
# Synthesis Paths



Always reaction of RI is important when heavier elements is produced.

これらの反応の研究は RIビームで初めて可能

# Synthesis Paths



Always reaction of RI is important when heavier elements is produced.

Study of such reactions are possible only using RI beams.



# RIB opened new era

---

- ❖ **in nuclear structure physics,**
  - ❖ *not only in nuclear size and shapes but also reveal other problems and raised new questions.*
- ❖ **in nuclear astrophysics and in cosmos-nuclear physics,**
  - ❖ *to understand the evolution of the universe and stars,*
  - ❖ *and nucleosynthesis in many stellar environments.*
- ❖ **more applications are expected.**