Are tractor beams possible?

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Tractor beam

Is this possible?



- A wave is scattered
- The scatterer is pulled towards the source () () () () () ()

NRP

Radiation pressure

Scattering of light:

$$P = (1 + R^2 - T^2)W,$$
$$W = \epsilon_0 E_0^2/2$$



Can the pressure be negative? What objects can be pulled towards the radiation source? Unitarity (conservation of energy): $R^2 + T^2 = 1$, $P = 2WR^2 > 0$. How can *P* be negative? Intro Models 2D scat AB 2ch AB Summary

NRP

What is negative radiation pressure?

Mizraki, A. and Fainman, Y., *Opt. Lett.* **35** 3405 (2010) "gain medium": No unitarity!, T > 1



Is this really negative radiation pressure?

Intro Models 2D scat AB 2ch AB Summary

What is negative radiation pressure?

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NRP

Is this really negative radiation pressure?



Does the bullet pull the rocket?

Tricky beam

Sukhov, S. and Dogariu, A. Phys. Rev. Lett. 107 203602 (2011)





- Geometrical optics: very simple
- They have also solved Maxwell's equations (FEL)
- Also possible for general (absorption free) scatterers
- Examples: sphere, a number of tiny spheres
- One has to know scattering data of scatterer

NR

Two channel scattering

- Consider two channels with different momenta e.g., spin up and down particles in a magnetic field $B_z < 0$
- Incoming beam: smaller momentum channel, x direction
- Scattering between channels e.g., revert some spins, the scatterer can be a region with $B_x \neq 0$
- Outgoing ptcles carry more momentum
- Balanced by force acting on the scatterer in the direction opposite to that of the incoming beam
- Incoming plane wave acts as a tractor beam

A simple model: $k_i^2 = \omega^2 - m_i^2$, $m_u > m_d$ Scatterer: constant potential for $-L \le x \le L$

$$F_u/A_u^2 = k_u(1 + |R_{uu}|^2 - |T_{uu}|^2) + k_d(|R_{du}|^2 - |T_{du}|^2)$$

Wide range of parameters: effects of reflection dominated by scattering between channels

Outline



A strip of false vacuum I



- Vacuum: two modes: $k_a^2 + 4 = \omega^2$, $k_g^2 = \omega^2$
- False vacuum: two modes, $k_i^2 = \omega^2$
- Phase shift screen: $\pm \delta$ for the two modes

A strip of false vacuum II

- Analytical force formula $\frac{F_{\text{frs}}}{|a|^2} = k_a^2(1 + |R_a|^2 - |T_a|^2) + k_g^2(|R_g|^2 - |T_g|^2)$
- very simple for $L \rightarrow 0$

$$F_{L\to 0} = a^2 \frac{4k_a^2(k_a^2 - k_g^2)\sin^2 \delta}{k_a^2 + 6k_ak_g + k_g^2 - (k_a - k_g)^2\cos 2\delta}$$

and also neglecting multiple reflections

$$F_{\rm fvs} pprox |a|^2 rac{2k_a^2(k_a^2(k_i+k_g)^2+k_i^2(k_i^2+2k_ik_g-7k_g^2))}{(k_a+k_i)^2(k_i+k_g)^2}$$

• Oscillations: Fabry-Perot (or: antireflex coating)

A strip of false vacuum III



Index of refraction I

Similar to the false vacuum strip:



- Outside: two modes, "a" and "g"
- Inside: two modes, opposite phase shifts
- $\omega_i = c_i k_i$, $c_i = c/n_i$, $n_{a,g,b}$: indices of refraction

Index of refraction II

• analytical force formula for $L \rightarrow 0$:

$$F_{a,L=0} = rac{2(n_g - n_a)}{n_a^2(n_a + n_g)} \omega^2 \,, \qquad F < 0 ext{ if } n_a > n_g$$



Outline



2D scattering

Assumptions: free propagation far from the scatterer Scattering asymptotics: multichannel scattering

$$\psi_{a} \sim A_{a} \mathrm{e}^{\mathrm{i}k_{a}x} + \frac{\mathrm{e}^{\mathrm{i}k_{a}r}}{\sqrt{r}} \sum_{b} f_{ab}(\vartheta) A_{b}$$

Cross section from channel b to a

$$rac{\mathrm{d}\sigma_{ab}}{\mathrm{d}\vartheta} = rac{k_a}{k_b} |f_{ab}(\vartheta)|^2$$

defined as energy flux into angle ϑ for unit flux incoming wave Scattering operator

$$S_{ab}(artheta) = \delta_{ab}\delta(artheta) + \sqrt{rac{ik_a}{2\pi}}f_{ab}(artheta)$$

mapping: incoming wave \mapsto outgoing wave

Partial waves

Easier to solve ODEs

$$\psi_{a} = \sum_{\ell=-\infty}^{\infty} \mathrm{e}^{i\ell\vartheta} R_{a,\ell}(r)$$

with radial functions $R_{a,\ell}$

S-matrix

Incoming plane wave $e^{ikx} = \sum_{\ell} i^{\ell} e^{i\ell\vartheta} J_{\ell}(kr)$; Asy for $r \to \infty$

 $R_\ell \sim (\dots) (\underline{S_{ab,\ell}} A_b H^{(1)} + H^{(2)} A_a), \quad H^{(1,2)} = J_\ell \pm i Y_\ell \sim \mathrm{e}^{\pm i k r} / \sqrt{r}$

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- Multiple (*n*) channels S_{ℓ} an $n \times n$ matrix
- Unitarity (energy conservation) $SS^{\dagger} = 1$, optical thm
- Symmetry: time reversal (in many cases) Perturbative problems: can be violated by the background

Force acting on the scatterer I

Momentum balance Assuming free propagation for $r \to \infty$, substitute into

$$\mathbf{F} = -\lim_{R \to \infty} R \int_{-\pi}^{\pi} \bar{\mathbf{T}} \mathbf{e}_{r} R \mathrm{d} \vartheta \, ,$$

T: stress tensor, **e**_r radial unit vector

Using the partial wave expansion a master formula is obtained:

$$F = F_x + iF_y = -4\sum_{\ell} \left\{ A^{\dagger} S_{\ell+1}^{\dagger} K S_{\ell} A - A^{\dagger} K A \right\} \,,$$

 $A = (A_1, ..., A_n)^T$ amplitude, $K = \text{diag}(k_1, ..., k_n)$ wave numbers Consequence of unitarity:

$$\operatorname{\mathsf{Re}}|A_a|^2k_a(1-S^*_{aa,\ell+1}S_{aa\ell})>0$$

one channel radiation pressure positive; NRP: $k_b > k_a$ necessary

Force acting on the scatterer II

Force expressed with cross sections

$$\begin{split} F_x &= -2k^2\sigma_{\parallel}|A|^2\,,\\ F_y &= -2k^2\sigma_{\perp}|A|^2\,, \end{split}$$

with $\sigma_{\parallel} = \int_{-\pi}^{\pi} d\sigma(\vartheta) (\cos \vartheta - 1)$ and $\sigma_{\perp} = \int_{-\pi}^{\pi} d\sigma(\vartheta) \sin \vartheta$ Also in QM: average momentum transfer

$$F = F_x + iF_y = -4\sum_{\ell} \left\{ A^{\dagger} S_{\ell+1}^{\dagger} K S_{\ell} A - A^{\dagger} K A \right\} \,,$$

- Neighbouring partial waves $(\ell, \ell + 1)$
- Positive contribution of diagonal elements (unitarity)
- More general than expression with cross-section (asymptotics of partial waves, not sum)

Outline



Aharonov–Bohm scattering I

The Aharonov–Bohm effect: Motion of a charged particle in a region with $\mathbf{B} = 0$ Double slit expreriment: Scattering:



- Both experiments show flux dependence
- Holonomy is also physical not just field strength $\mathbb{P}e^{i\int \mathbf{A}d\mathbf{r}}$
- Reaction force (deflected beam): Force acting on the scatterer

Aharonov–Bohm scattering II

Schrödinger-equation

$$-i\dot{\psi}=(\nabla-i\mathbf{A})^{2}\psi,$$

with electromagnetic vector potential

$$\mathbf{A}(r,artheta,z)=rac{A_0}{r}\mathbf{e}_artheta$$
 .

 $2\pi A_0$ flux; outside **B** = 0. Fixed energy: $\psi(\mathbf{r}, t) = e^{-i\omega t}\psi(\mathbf{r})$. Scattering asymptotics (?)

$$\psi \sim \mathrm{e}^{ikx} + \frac{f(\vartheta)}{\sqrt{r}}\mathrm{e}^{ikr}$$

Cross sections: $d\sigma/d\vartheta = |f(\vartheta)|^2$ Scattering amplitude: $f(\vartheta) \sim \frac{\sin^2 \pi A_0}{2\pi} \frac{1}{\sin^2(\vartheta/2)}$.

Aharonov–Bohm scattering III

Partial waves: $s_{\ell} \propto J_{|\ell-A_0|}$ (note index shift wrt plane wave) Effect of scattering in the outgoing wave: phase shift δ_{ℓ} , $S_{\ell} = \exp(2i\delta_{\ell})$ $(J_{\nu}(z) \sim \cos(z - \nu\pi/2 - \pi/4)/\sqrt{2\pi z})$

$$\ell \ge 0 : \ \delta_{\ell} = \frac{A_0 \pi}{2} \,, \quad \ell \le -1 \;:\; \delta_{\ell} = \pi \ell - \frac{A_0 \pi}{2}$$

thus

$$\begin{split} F_x &= -|\phi_0|^2 4k \sum_{\ell=-\infty}^{\infty} \left\{ \cos\left[2(\delta_\ell - \delta_{\ell-1})\right] - 1 \right\} \\ &= -|\phi_0|^2 4k \left(\cos(2\pi A_0) - 1 \right) \approx 16 |\phi_0|^2 \pi^2 A_0^2 k , \\ F_y &= -|\phi_0|^2 4k \sum_{\ell=-\infty}^{\infty} \sin\left[2(\delta_\ell - \delta_{\ell-1})\right] = |\phi_0|^2 4k \sin(2\pi A_0) \approx 8 |\phi_0|^2 \pi A_0 k . \end{split}$$

Analog problem: force acting on a superfluid vortex (GPe)

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Aharonov–Bohm scattering IV

lordanskii force controversy C. Wexler, D.J. Thouless, *Phys. Rev.* **B58** R8897–R8900 (1998):

$$f(-\vartheta) = f(\vartheta) \quad \Rightarrow \quad F_y = 0$$

A.L. Shelankov *Europhys. Lett.* **43** (1998) 623 M.V. Berry *J. Phys. A: Math. Gen.* **32** (1999) 5627: scattering asymptotics does not hold in forward direction

$$\psi(\mathbf{r},t) = \mathrm{e}^{-i(\omega t - kx)}\phi(\mathbf{r})$$

and for $y \ll \sqrt{x}$

$$\phi(x > 0, y) \sim \cos(A_0 \pi) - \frac{2i^{1/2}}{\sqrt{\pi}} \sin(A_0 \pi) \sqrt{\frac{k}{2}} \frac{y}{\sqrt{x}}$$

transversal force F_y from this region (although not F_x)

Outline



Scattering on a vortex I

Consider two scalar fields, ϕ and ψ

- A vortex in φ: φ ~ ν exp(iϑ) (broken U(1) symmetry)
 e.g. φ⁴ potential V = λ(φ*φ − ν²)²/2
- Scattering of ψ
- Coupling $\Delta \mathcal{L} = \boldsymbol{g} \phi \psi^2 + \boldsymbol{c}. \boldsymbol{c}.$

Field equation

$$(\omega^2 - m^2 + \nabla^2)\psi = 2g\phi^*\psi^*$$
,

Coupling for $r \to \infty$: transformation to mass eigenstates

$$\begin{pmatrix} \psi \\ \psi^* \end{pmatrix} = \begin{pmatrix} \mathrm{e}^{-i\vartheta/2} & i\mathrm{e}^{-i\vartheta/2} \\ \mathrm{e}^{i\vartheta/2} & -i\mathrm{e}^{i\vartheta/2} \end{pmatrix} \begin{pmatrix} \mathsf{u} \\ \mathsf{d} \end{pmatrix}$$

- heavy mode $u (m_u^2 = m^2 + 2gv)$
- light mode $d (m_d^2 = m^2 2gv)$
- transformation matrix depends on ccordinate ϑ

Scattering on a vortex II

$$U = \frac{1}{\sqrt{2}} \begin{pmatrix} e^{-i\vartheta/2} & ie^{-i\vartheta/2} \\ e^{i\vartheta/2} & -ie^{i\vartheta/2} \end{pmatrix}, \quad \psi = U\rho, \quad \rho = \begin{pmatrix} u \\ d \end{pmatrix}$$

- trf matrix U depends on angular coordinate θ
- Commuting U and ∇ : artificial gauge potential induced

$$i\mathbf{A}rac{\sigma_2}{2} = rac{1}{r}U^\dagger rac{\partial U}{\partial artheta} \mathbf{e}_{artheta}$$

where $\mathbf{A} = \mathbf{e}_{\vartheta}/r$ (Kazan 1985; March-Russell, Preskill, Wilczek 1992)

Field equations

$$\left(
abla + i \mathbf{A} \frac{\sigma_2}{2}
ight)^2
ho - \mathcal{K}^2
ho = \mathbf{0} \,, \quad \sigma_2 = \begin{pmatrix} & -i \\ i & \end{pmatrix} \,,$$

- two channels, different masses
- coupling due to Aharonov–Bohm gauge potential

Solution of the scattering problem I

Partial waves

$$(u,d) = \sum_{\ell=-\infty}^{\infty} \mathrm{e}^{i(\ell+\gamma)artheta)} (u_\ell(r),d_\ell(r))$$

with $\gamma = 1/2$. The radial functions satisfy

$$egin{aligned} & u_\ell'' + rac{u_\ell'}{r} - rac{\eta_u^2}{r^2} u_\ell + rac{c}{r^2} d_\ell + k_u^2 u_\ell = 0 \,, \ & d_\ell'' + rac{d_\ell'}{r} - rac{\eta_d^2}{r^2} d_\ell + rac{c^*}{r^2} u_\ell + k_d^2 d_\ell = 0 \,, \end{aligned}$$

where $\eta_u^2 = \eta_d^2 = (\ell + 1/2)^2 + 1/4$ and $c = i(\ell + 1/2)$. We use rescaled variables such that v = 1, $m_u = 2$ (threshold at $\omega = 2$), and for simplicity sake, present numerical data for $m_d = 1$.

Approximate solution of radial equations I

Decoupled approximation: ignore coupling

$$d_\ell(r) = A_\ell J_{\eta d}(k_d r)$$

and thus

$$oldsymbol{S}_\ell = egin{pmatrix} \mathrm{e}^{2i\xi_u} & 0 \ 0 & \mathrm{e}^{2i\xi_d} \end{pmatrix} \quad \xi_{u,d} = rac{\pi}{2}(\ell-\eta_{u,d})$$

works well for $\omega \ll m_d$ (closed channel *u*), $F_y = 0$ Numerical solution (above the threshold)

- *u* channel: $F_{x,u} < 0$ tractor beam
- Phases depend weakly on frequency
- Scattering between channels important

•
$$F_y \neq 0$$

Approximate solution of radial equations II

Simplified perturbation theory

• calculate phase for *d* at the threshold (2nd order pert th)

$$\delta_d = \xi_d + \Delta$$
, with $\Delta = \arctan\left[\frac{|c|^2 \pi/4}{\eta_u^2 \eta_d + \eta_u \eta_d^2}
ight]$

calculate matrix elements with Bessel functions with the right phase

$$S_{\ell} pprox rac{1}{N} egin{pmatrix} \mathrm{e}^{2i\delta_u} & \mathrm{e}^{i(\delta_u+\delta_d)}i\pi cI \ \mathrm{e}^{i(\delta_u+\delta_d)}i\pi c^*I & \mathrm{e}^{2i\delta_d} \end{pmatrix}$$

with

$$I = \left(\frac{k_u}{k_d}\right)^{\nu_u} \Gamma(\bar{\nu}) \frac{{}_2F_1\left(\nu', \bar{\nu}, \nu_u + 1, k_u^2/k_d^2\right)}{2\Gamma\left(1 - \nu'\right)\Gamma(\nu_u + 1)} ,$$

where $\nu_u = \eta_u$, $\delta_u = \pi(\ell - \nu_u)/2$, $\bar{\nu} = (\nu_u + \nu_d)/2$, $\nu' = (\nu_u - \nu_d)/2$, and *N* is a normalization

Force acting on the scatterer I



Force acting on the scatterer II



How good is the solution



Errors for small *l*: origin

Effect of a vortex core

More realistic vortex configuration:

$$\phi(r) = f(r)e^{i\vartheta}$$
, $f(r \leq R_c) = vr/R_c$, $f(r \geq R_c) = v$



Lukács Á. Tractor beams possible?

Cosmic strings

- Classical field theoretical solution
- Internal structure
- Thin, elongated object
- One cross section of a string: a vortex
- Localised energy density in the string

Important parameter: string tension

$$\mu = E/L$$

- Elevtroweak string: $G\mu \approx 10^{-32}$ (μ : 10 mg/Solar diam)
- GUT string: $G\mu \approx 10^{-6}$ (μ : Solar mass/Solar diam)

Cosmic strings are high energy localised objects that provide a link between astrophysics and particle physics.

Physics of cosmic strings I

Evolution of cosmic strings

- Formation: during phase transitions
- Evolution of a string network
 - expansion: the network becomes more diuted
 - collisions, interlinking
 - radiation (e.g. at cusps formed in collisions)
 - string tension contracts loops
- Signatures of cosmic strings
 - Scattering of material off strings: structure formation: galaxies, voids, filamets (fractal dimension: Murdzek 2007)
 - Contribution to CMB anisotropy: best fit with GUT strings $G\mu = (2.04 \pm 0.13) \times 10^{-6}$, Contribution to multipole $\ell = 10: f_{10} = 0.11 \pm 0.05$ (Hindmarsh et al., 2007, 2008)
 - Gravitational lensing
 - Gravitational radiation

Physics of cosmic strings II

Analogies with condensed matter systems

- vortex filaments in superfluids
- Abrikosov vortices in superconductors
- defects in liquid crystals
- defect formation: Kibble–Zurek mechanism



(Kép forrása: Martins and Shellard,

http://www.damtp.cam.ac.uk/research/gr/public/cs_evol.html)



(Srivastava, Testing Cosmic Defect Formation Theories in Liquid Crystal Experiments, COSLAB 2005, Smolenice)

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Spontaneous symmetry breaking

• More symmetries at higher energies

 $\cdots \rightarrow \text{GUT} \rightarrow \textit{SU}(3)_{\text{color}} \times \textit{SU}(2)_L \times \textit{U}(1)_Y \rightarrow \textit{SU}(3)_{\text{color}} \times \textit{U}(1)_{\text{em}}$

transition from a symmetric state to one with broken symmetry: phase transition

- GUT $\sim 10^{15}~GeV$ (at $10^{-36}~s),$ electroweak: $\sim 250~GeV$ (at $10^{-12}~s)$
- Mass terms for gauge potentials: coupling to scalar (Higgs) field

$$m^2 A_\mu A^\mu o \Phi^\dagger \Phi A_\mu A^\mu$$

- Higgs effect: $\langle \Phi \rangle \neq 0$ in the broken symm phase
- Scalar field self interaction

$$V(\Phi) = \beta (\Phi^{\dagger} \Phi - \eta^2)^2$$

"Mexican hat", minima: $\Phi^{\dagger}\Phi=\eta^2$

• ground state degenerate: "phase" of $\langle \Phi \rangle$ undetermined

Scattering on a cosmic string I

Field equations: $\ddot{\phi} - \nabla^2 \phi + 2(\phi^* \phi - 1)\phi$ Similar to Gross–Pitaevskii, but dynamics is of second order



- Perturbations of the scalar field scattered
- Fits into the previous framework with $e_1 = e_2 = 0$, q = -1, $k_u^2 = \omega^2 4$, $k_d = \omega$.
- Two modes: amplitude/phase perturbed
- Vortex core is needed for NRP

Force acting on the cosmic string



- *F_x* vs ω for two channels: NRP: scattering of massive mode into massless one
- Preliminary time-evolution data (Romańczukiewicz): qualitative agreement
- Iordanskii force vanishes $F_y = 0$

Outline



Summary

Negative radiation pressure

- "Tractor beam" fairly generic in multichannel scattering
- many models in optics (beam prepared for the scattered)
- field theoretical models: surplus momentum in forward scattered wave
 - kink: nonlinearities
 - vortex: two channels
- models motivated by the case of scattering on the vortex: two channels
 - strip of false vacuum
 - different index of refraction
- effect expected in multi-channel scattering

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THANK YOU FOR YOUR ATTENTION!