Suppression of anisotropic flow without viscosity

Adam Takacs^{*} University of Bergen (Norway) & Eotvos University (Hungary) Denes Molnar Purdue University (U.S.) Gergely G. Barnafoldi Wigner Institute (Hungary) Konrad Tywoniuk University of Bergen (Norway)



*takacs.adam@wigner.mta.hu



Emberi Erőforrások Minisztériuma Supported by the 18-3 New National Excellence Program of the Ministry of Human Capacities and OTKA K120660.

Balaton Workshop, 17-21. June 2019, Tihany

Outline

1. Motivation:

relativistic hydrodynamics to understand heavy ion collisions

2. Particlization:

basics, shortcomings, influence on observables

- 3. How much does f matter?
- 4. Results: 4-source model and 2+1D hydro
- 5. Conclusion and outlook



Motivation



For students: see Gabriel Denicol's talk about hydro from Hot Quarks 2018 <u>https://indico.cern.ch/event/7</u> 03015/contributions/3095199/



Motivation

D. Molnar & P. Huovien, J.Phys. G35 104125 (2008)
K. Dusling, G.D. Moore, D. Teaney, Phys. Rev. C81, 034907 (2010)
D. Molnar & Z. Wolff, Phys. Rev. C95, 024903 (2017)

- Shear viscosity is sensitive to the way of particlization
- There is no unique way to do the particlization
- Thermal equilibrium is an assumption



Motivation

D. Molnar & P. Huovien, J.Phys. G35 104125 (2008)
K. Dusling, G.D. Moore, D. Teaney, Phys. Rev. C81, 034907 (2010)
D. Molnar & Z. Wolff, Phys. Rev. C95, 024903 (2017)

- Shear viscosity is sensitive to the way of particlization
- There is no unique way to do the particlization
- Thermal equilibrium is an assumption



Questions:

- 1. How sensitive is ideal hydro?
- 2. Is thermal equilibrium a legitimate assumption?



Particlization

How is it done?



L. Csernai, Introduction Relativistic Heavy Ion Collisions (1994)

Conversion of fluid \rightarrow particles on a 3D hypersurface: $d\sigma^{\mu}$

Inside: fluid $(u^{\mu}(x), \varepsilon(x), P(x), n(x))$ $N^{\mu}(x) = (n(x), \vec{j}(x))$ $T^{\mu\nu}(x) = [\varepsilon(x) + P(x)]u^{\mu}u^{\nu} - P(x)g^{\mu\nu}$





L. Csernai, Introduction Relativistic Heavy Ion Collisions (1994)

Conversion of fluid \rightarrow particles on a 3D hypersurface: $d\sigma^{\mu}$

Inside: fluid
$$(u^{\mu}(x), \varepsilon(x), P(x), n(x))$$

 $N^{\mu}(x) = (n(x), \vec{j}(x))$
 $T^{\mu\nu}(x) = [\varepsilon(x) + P(x)]u^{\mu}u^{\nu} - P(x)g^{\mu\nu}$



Outside: particle f(x, p)

$$N^{\mu}(x) = \int \frac{d^{3}p}{p^{0}} p^{\mu} f(x, p)$$
$$T^{\mu\nu}(x) = \int \frac{d^{3}p}{p^{0}} p^{\mu} p^{\nu} f(x, p)$$



L. Csernai, Introduction Relativistic Heavy Ion Collisions (1994)

Conversion of fluid \rightarrow particles on a 3D hypersurface: $d\sigma^{\mu}$

Inside: fluid
$$(u^{\mu}(x), \varepsilon(x), P(x), n(x))$$

$$N^{\mu}(x) = (n(x), \vec{j}(x))$$

$$T^{\mu\nu}(x) = [\varepsilon(x) + P(x)]u^{\mu}u^{\nu} - P(x)g^{\mu\nu}$$



Outside: particle f(x, p)

$$N^{\mu}(x) = \int \frac{d^{3}p}{p^{0}} p^{\mu} f(x,p)$$
$$T^{\mu\nu}(x) = \int \frac{d^{3}p}{p^{0}} p^{\mu} p^{\nu} f(x,p)$$

Boundary: conservation laws

$$[N^{\mu}d\sigma_{\mu}]_{\rm in-out} = 0$$
$$[T^{\mu\nu}d\sigma_{\mu}]_{\rm in-out} = 0$$

L. Csernai, Introduction Relativistic Heavy Ion Collisions (1994)

Conversion of fluid \rightarrow particles on a 3D hypersurface: $d\sigma^{\mu}$

Inside: fluid
$$(u^{\mu}(x), \varepsilon(x), P(x), n(x))$$

$$N^{\mu}(x) = (n(x), \vec{j}(x))$$

$$T^{\mu\nu}(x) = [\varepsilon(x) + P(x)]u^{\mu}u^{\nu} - P(x)g^{\mu\nu}$$



Outside: particle f(x, p)

$$N^{\mu}(x) = \int \frac{d^{3}p}{p^{0}} p^{\mu} f(x, p)$$
$$T^{\mu\nu}(x) = \int \frac{d^{3}p}{p^{0}} p^{\mu} p^{\nu} f(x, p)$$

Cooper-Frye formula: momentum spectrum (measurable)

$$E\frac{d^3N}{d^3p} = \int d\sigma_\mu p^\mu f(x,p)$$

L. Csernai, Introduction Relativistic Heavy Ion Collisions (1994)

Conversion of fluid \rightarrow particles on a 3D hypersurface: $d\sigma^{\mu}$

inside: fluid
$$(u^{\mu}(x), \varepsilon(x), P(x), n(x))$$

$$N^{\mu}(x) = (n(x), \vec{j}(x))$$

$$T^{\mu\nu}(x) = [\varepsilon(x) + P(x)]u^{\mu}u^{\nu} - P(x)g^{\mu\nu}$$

$$d\sigma_{\mu}$$

Outside: particle f(x, p) $N^{\mu}(x) = \int \frac{d^3p}{p^0} p^{\mu} f(x, p)$ $T^{\mu\nu}(x) = \int \frac{d^3p}{p^0} p^{\mu} p^{\nu} f(x, p)$

What is f(x, p)?

Cooper-Frye formula: momentum spectrum (measurable)

$$E\frac{d^3N}{d^3p} = \int d\sigma_\mu p^\mu f(x,p)$$

- $f(\boldsymbol{x},\boldsymbol{p})$ single-particle distribution
- We don't know.



$f(\boldsymbol{x},\boldsymbol{p})$ - single-particle distribution

• Described by some kinetic transport theory (assumption)



$f(\boldsymbol{x},\boldsymbol{p})$ - single-particle distribution

• Described by the <u>Boltzman Transport Equation</u> (assumption of assumption)



$f(\boldsymbol{x},\boldsymbol{p})$ - single-particle distribution

 0th solution of the BTE (assumption of assumption of assumption) Relativistic Boltzmann (thermal) distribution:

$$f_0(x,p) = \frac{g}{(2\pi)^3} e^{-\frac{p^{\mu} u_{\mu}(x)}{T(x)}}$$

One field describes everything: T(x) temperature.



 $f(\boldsymbol{x},\boldsymbol{p})$ - single-particle distribution

 0th solution of the BTE (assumption of assumption of assumption) Relativistic Boltzmann (thermal) distribution:

$$f_0(x,p) = \frac{g}{(2\pi)^3} e^{-\frac{p^{\mu} u_{\mu}(x)}{T(x)}}$$

One field describes everything: T(x) temperature.

• Change a little to "non-extensive": corrections

$$f(x,p) = A \left[1 + \frac{\alpha}{T_{\alpha}(x)} p^{\mu} u_{\mu}(x) \right]^{-\frac{1}{2}}$$

One field $T_{\alpha}(x)$, "temperature" + non-ext. parameter α .

If $\alpha \to 0$ we get back Boltzmann! Starts in exponential, end in power law!

 $f(\boldsymbol{x},\boldsymbol{p})$ - single-particle distribution

 0th solution of the BTE (assumption of assumption of assumption) Relativistic Boltzmann (thermal) distribution:

$$f_0(x,p) = \frac{g}{(2\pi)^3} e^{-\frac{p^{\mu} u_{\mu}(x)}{T(x)}}$$

One field describes everything: T(x) temperature.

• Change a little to "non-extensive": corrections

$$f(x,p) = A \left[1 + \frac{\alpha}{T_{\alpha}(x)} p^{\mu} u_{\mu}(x) \right]^{-\frac{1}{\alpha}}$$

One field $T_{\alpha}(x)$, "temperature" + non-ext. parameter α .

If $\alpha \to 0$ we get back Boltzmann! Starts in exponential, end in power law!



How does spectrum and flow change?

• Spectrum: exponential like \rightarrow power-law tail (sounds good)



How does spectrum and flow change?

- Spectrum: exponential like \rightarrow power-law tail (sounds good)
- Flow:

$$v_n(p) = \frac{\int d\phi \cos(n\phi) E \frac{d^3 N}{d^3 p}}{\int d\phi E \frac{d^3 N}{d^3 p}}$$

Boltzmann:

$$v_n(p) \sim \frac{\mathrm{e}^{-ap}}{\mathrm{e}^{-bp}} = \mathrm{e}^{-(a-b)p} \to 1$$

Tsallis:

$$v_n(p) \sim \frac{(ap)^{-1/\alpha}}{(bp)^{-1/\alpha}} \to \left(\frac{a}{b}\right)^{-\frac{1}{\alpha}} \le 1$$
 (looks interesting!)



Hydrodynamic Simulation

- 1. Initial condition: Au+Au @ 200 GeV optical Glauber model $S_0=110$ fm⁻³.
- 2. 2+1D numerical ideal hydro with Azhydro.
- 3. Cooper-Frye freeze-out with Tsallis $\varepsilon(x), P(x)$ and q are fixed, $T_f = 140 \text{ MeV}$
- 4. Resonance decays are included.









- Power-law tail appears
- Better agreement with data!











- Suppression in flow!
- Better agreement with data!
- Mass ordering in the suppression, like shear viscosity!

Backup: changes in other observables





Viscous calculation:





- Non-ext. behaves like shear viscosity (no viscosity in the model)!
- Comparison to viscous calculation: $\alpha \approx 0.05 \leftrightarrow \eta/s = 0.05$
- From fits to spectra: $\alpha = 0 0.07$ K.Urmossy, G.G.Barnafoldi, T.S.Biro,

J.Phys. Conf.Ser. 612 012048 (2015)



COST Workshop Lund

28-02-2019

Summary

- Relativistic Boltzmann distribution (thermal equilibrium) is an approximation.
- To study a specific type of correction (using Tsallis distribution):
 - 1. Finite size effects and correlations.
 - 2. More suitable for the spectrum (power law tail) and the flow $(v_n < 1)$.
 - 3. Isotropization? Conformal theories do not require equilibrium.
- The correction has important effects:
 - Better spectra.
 - Suppressed flow
 - Mimic shear viscosity: $\alpha \approx 0.05 \leftrightarrow \eta/s = 0.05$
- Future:
 - Extend to viscous calculation.
 - Study kinetic transport for an exact correction. Better understand α

P.Arnold, J.Lenaghan, G.D.Moore, L.G. Yaffe, Phys. Rev.
Lett. 94, 072302 (2005)
M.Luzum & P.Romatschke. Phys.Rev. C78 034915 (2008)



Thank you for your attention!







How does spectrum and flow change?

- Spectrum: exponential like \rightarrow power law tail
- Flow:

$$v_n(p) = \frac{\int d\phi \cos(n\phi) E \frac{d^3 N}{d^3 p}}{\int d\phi E \frac{d^3 N}{d^3 p}}$$

• Simple 4-source model: 4-uniform fireballs, no long. expansion, boosted sym. $(\pm v_x, \pm v_y)$ and T=const freezeout

$$f^{4s}(p) = f|_{v_x} + f|_{-v_x} + f|_{v_y} + f|_{-v_y}$$



P. Houven et al Phys. Lett. B 503, 58 (2001)



29

How does spectrum and flow change?

Result from 4-source model: suppression in v_2 !



Backup: describing hydro fields with non-ext. distr.







Backup: non-ext. parameter vs. shear viscosity





Backup: v_4 with non-ext. f



